

# HOT CONTINENTAL TECTONICS:

## DEFORMATION, FLOW, STRESS AND STRAIN

### INSIGHTS FROM NUMERICAL EXPERIMENTS

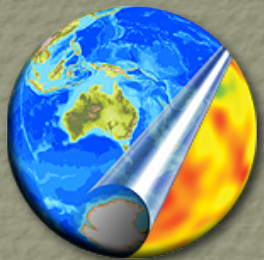
PATRICE F. REY,

NICOLAS THÉBAUD, GUILLAUME DUCLAUX, NICOLAS FLAMENT

CHRISTIAN TEYSSIER, DONNA L. WHITNEY

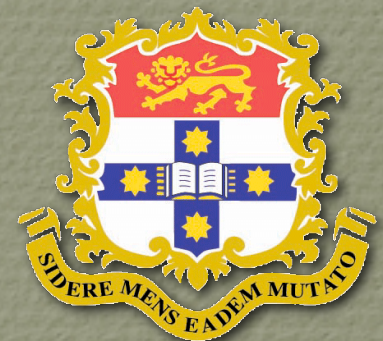
GREG HOUSEMAN, NICOLAS COLTICE

LOUIS MORESI, JULIAN GIORDANI, JOHN MANSOUR



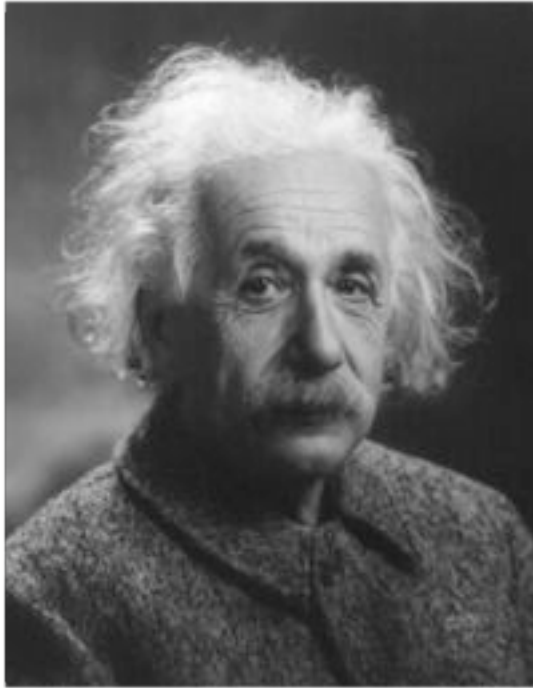
Earth **BYTE**

Linking observations to kinematic and dynamic models



The University of Sydney





“If at first the idea is not absurd,  
then there is no hope for it.”

- Deformation of hot continents consequences for early Earth
- Early continents and plate tectonics
- Strain regimes in hot crusts

# DEFORMATION OF HOT CONTINENTS



HOT?

$$T_{\text{MOHO}} > 0.85 T_{\text{MELT}}$$

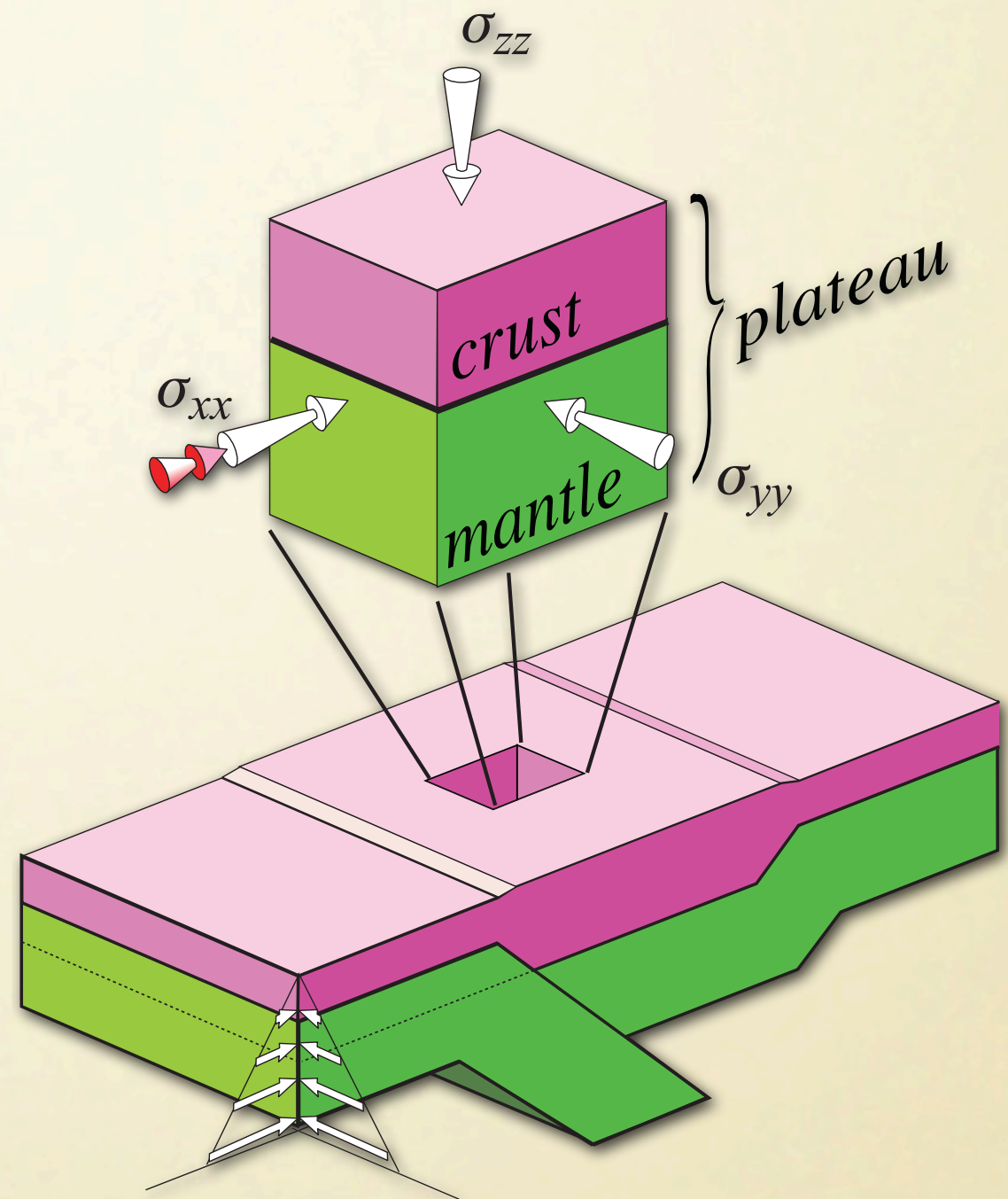
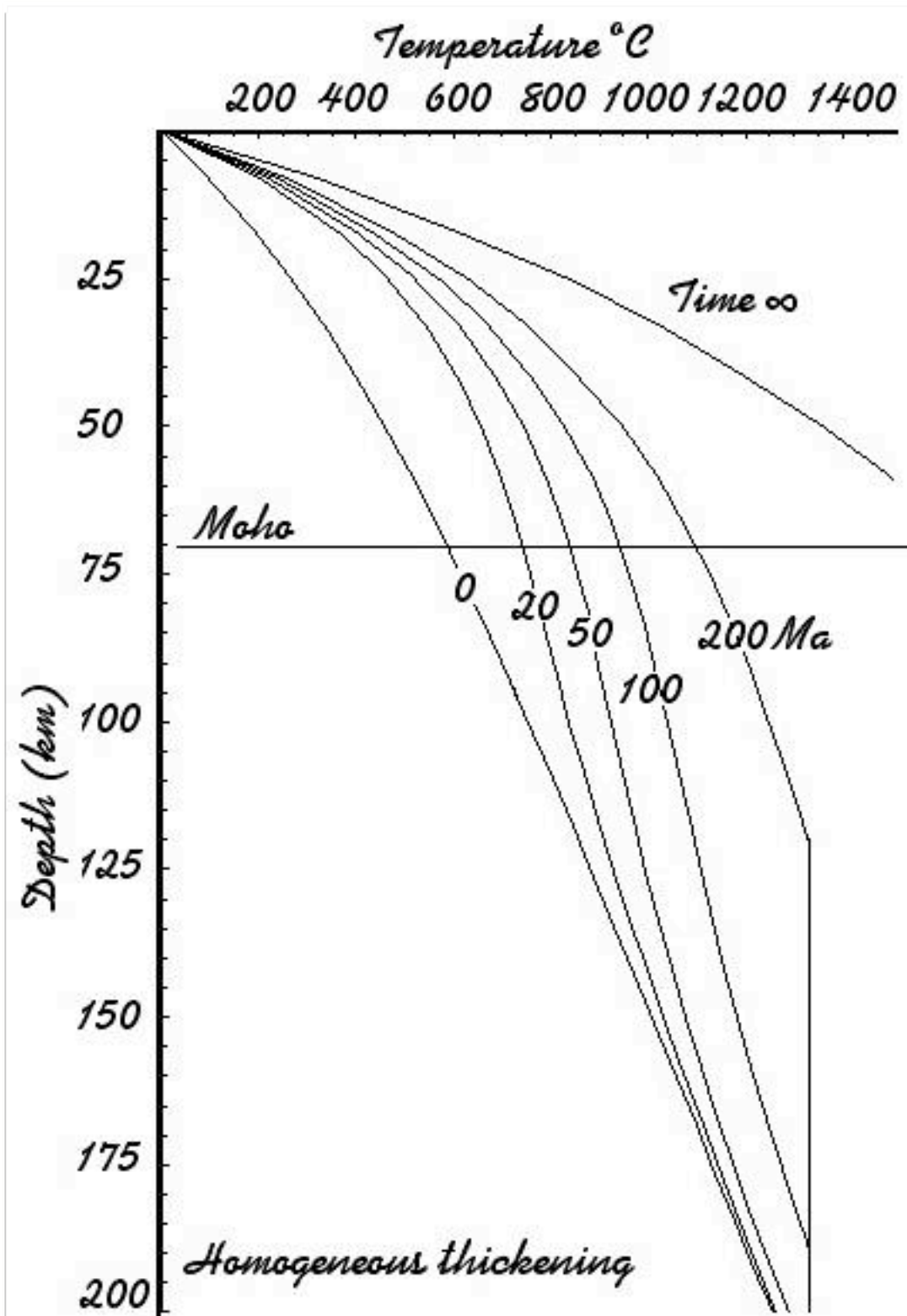


A faint, light-colored world map is visible in the background of the slide, centered behind the text.

# **WHERE DO HOT CONTINENTS COME FROM?**

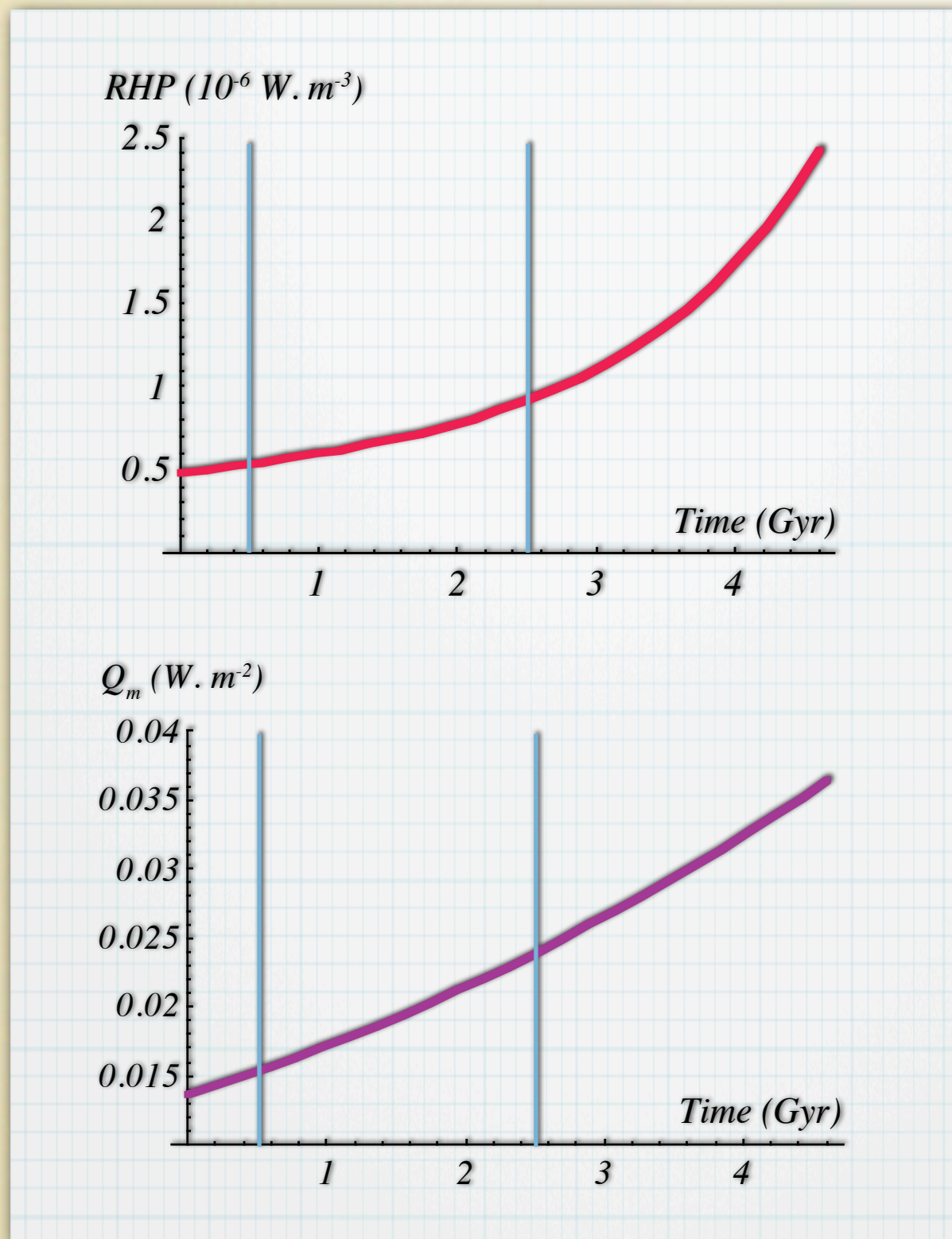


- Crustal thickening





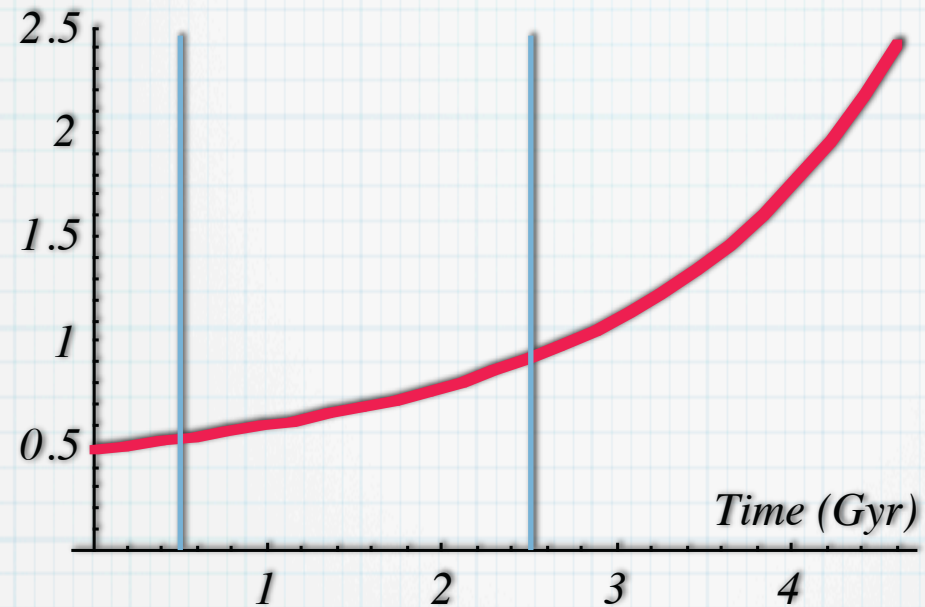
- Continental geotherm through time



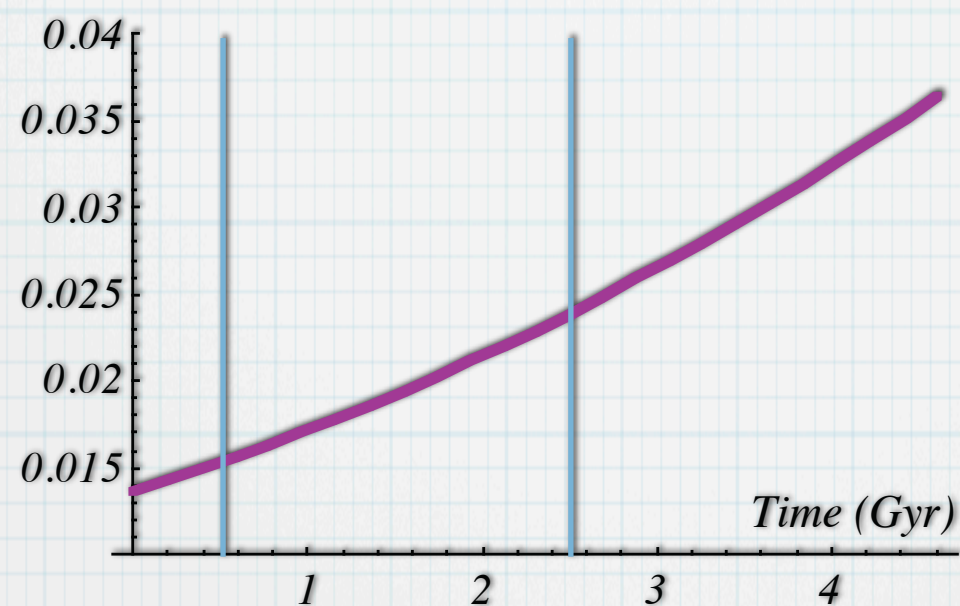


- Continental geotherm through time

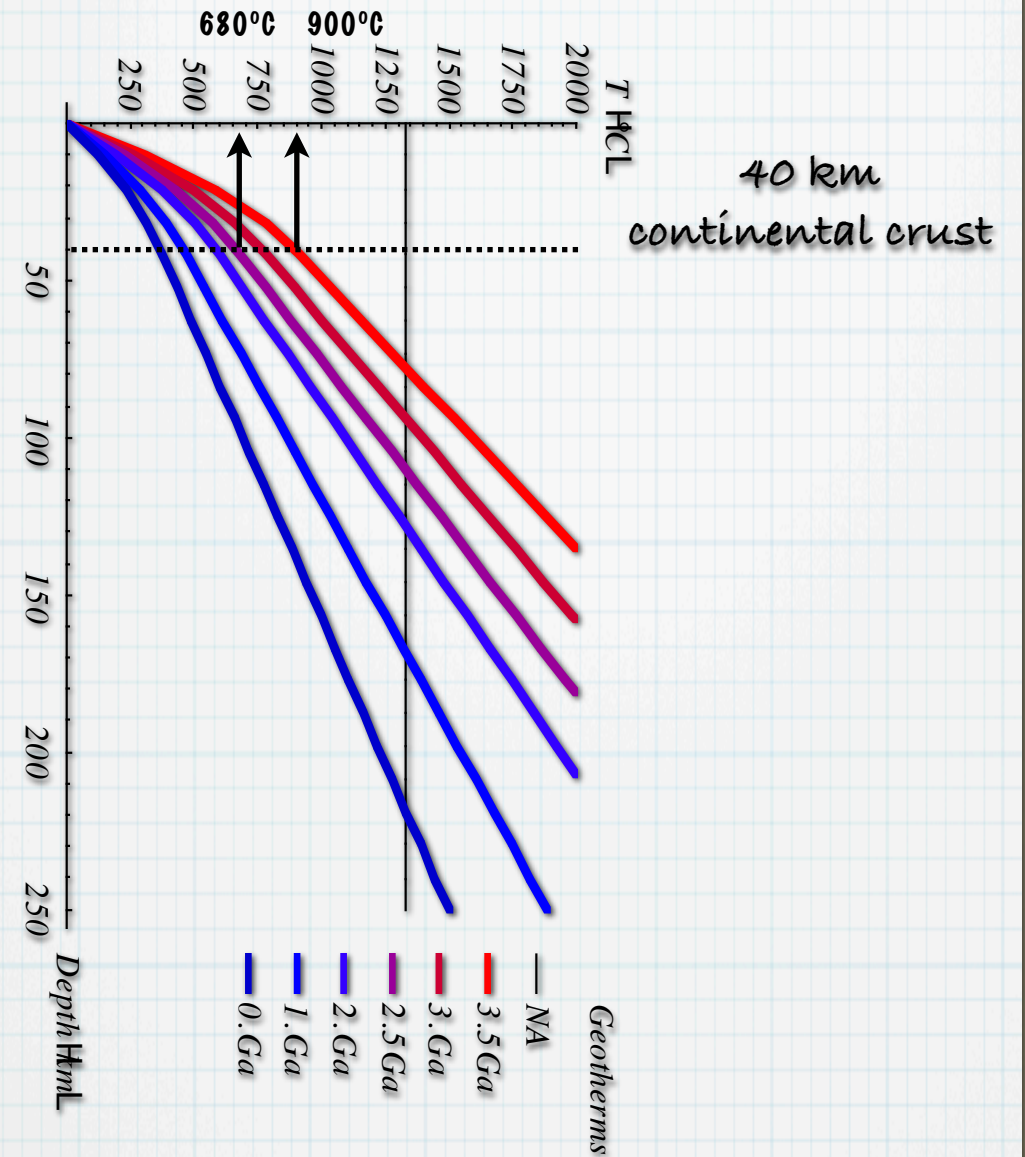
$RHP (10^{-6} W.m^{-3})$



$Q_m (W.m^{-2})$



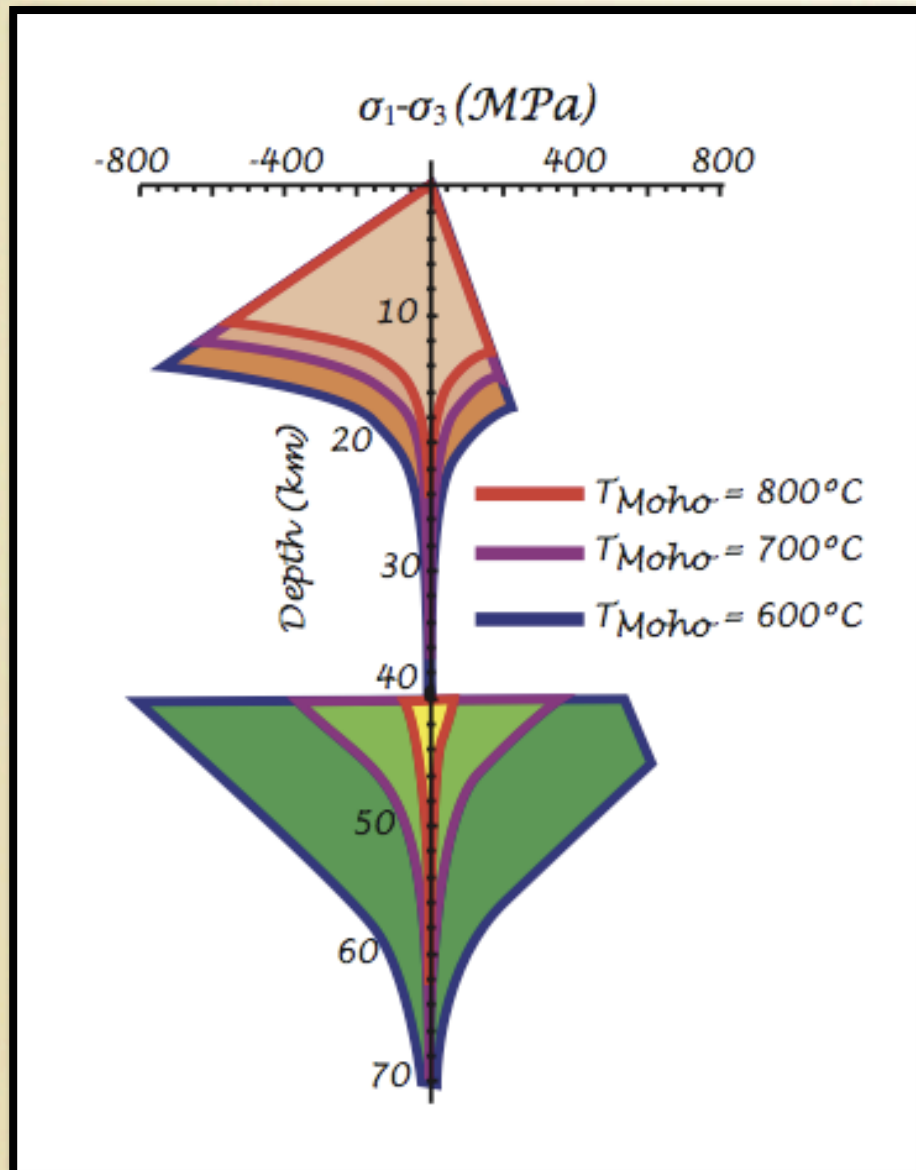
## Continental geotherm through time



@ 0.0 Ga:  $T_{moho} = 370^{\circ}C$ .  $Q_m = 0.013 W.m^{-2}$

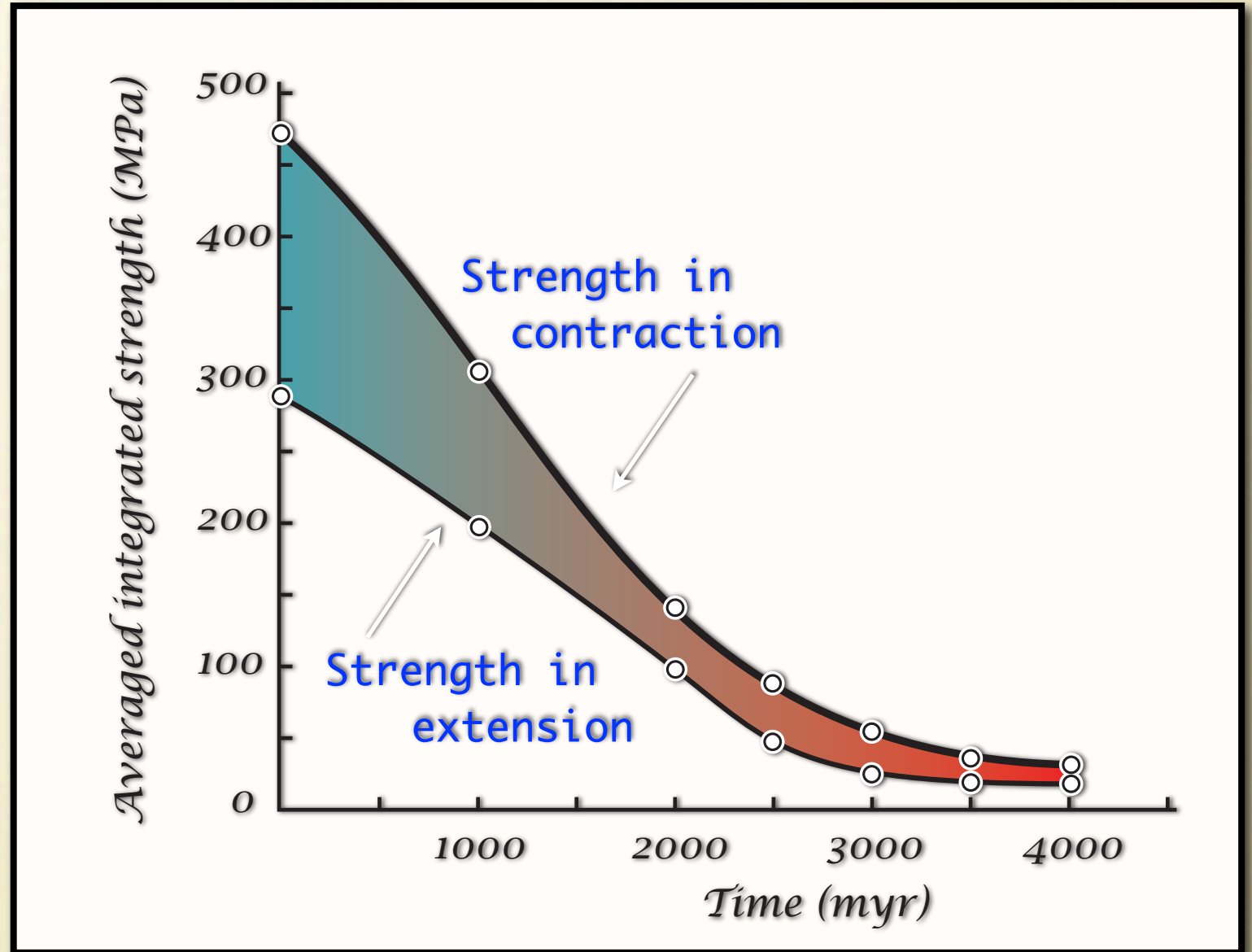
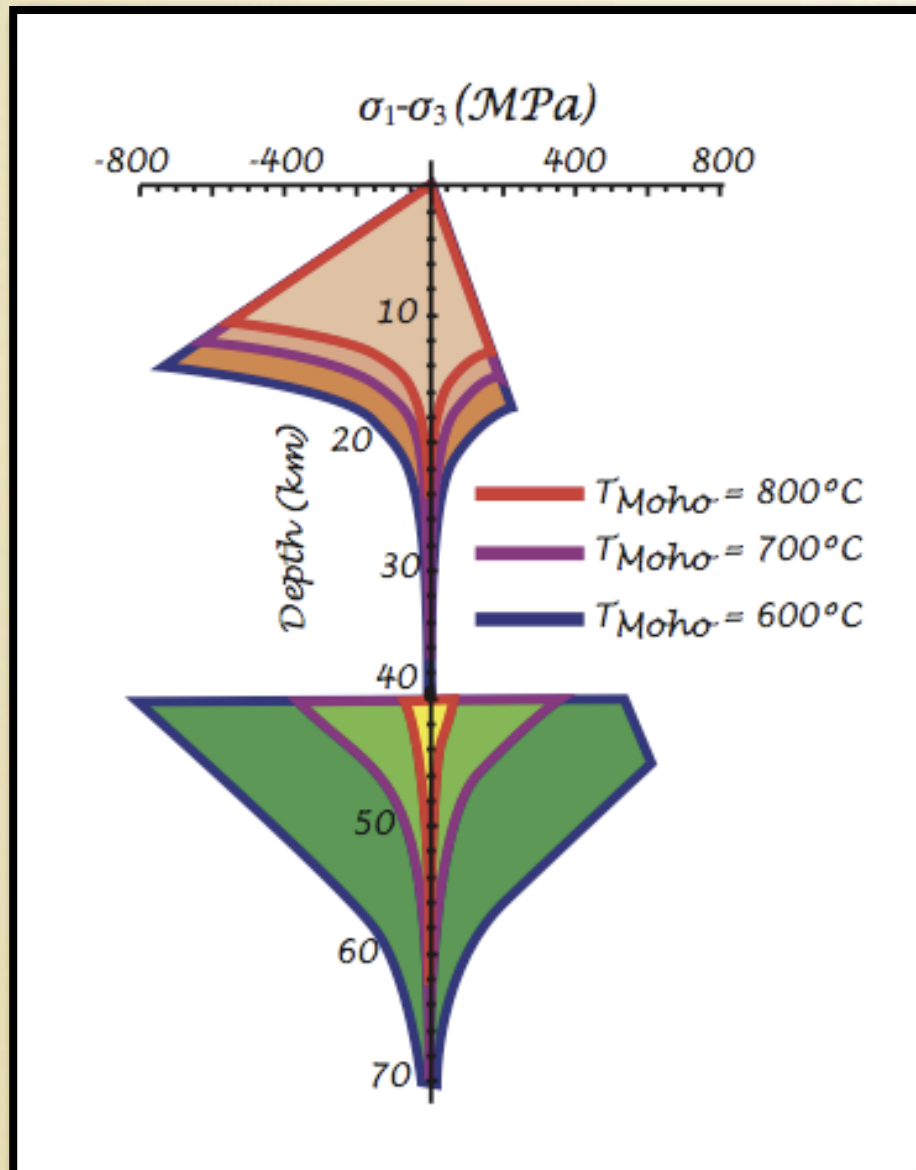
@ 2.7 Ga:  $T_{moho} = 700^{\circ}C$ ,  $Q_m = 0.025 W.m^{-2}$

# Viscous forces through time

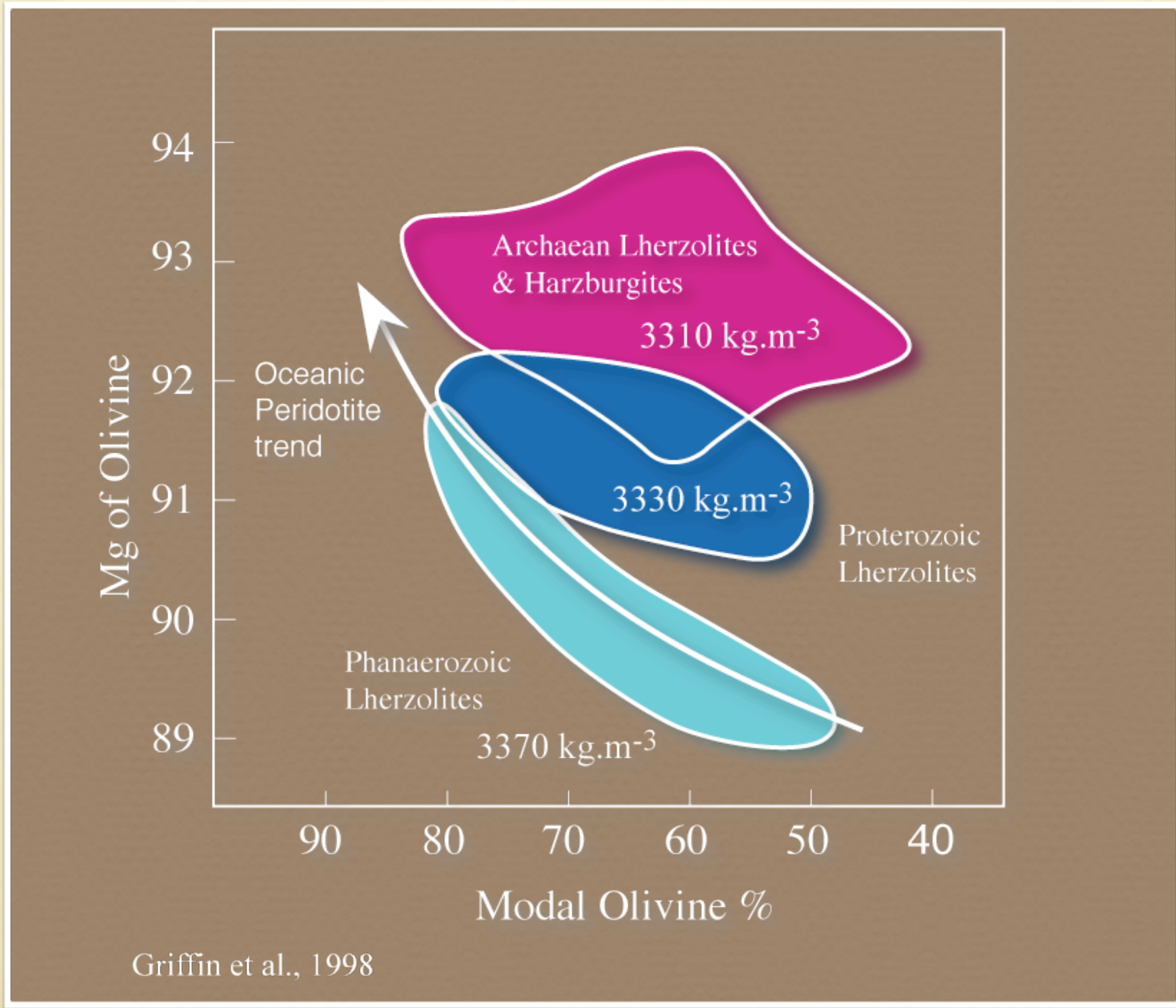




# Viscous forces through time

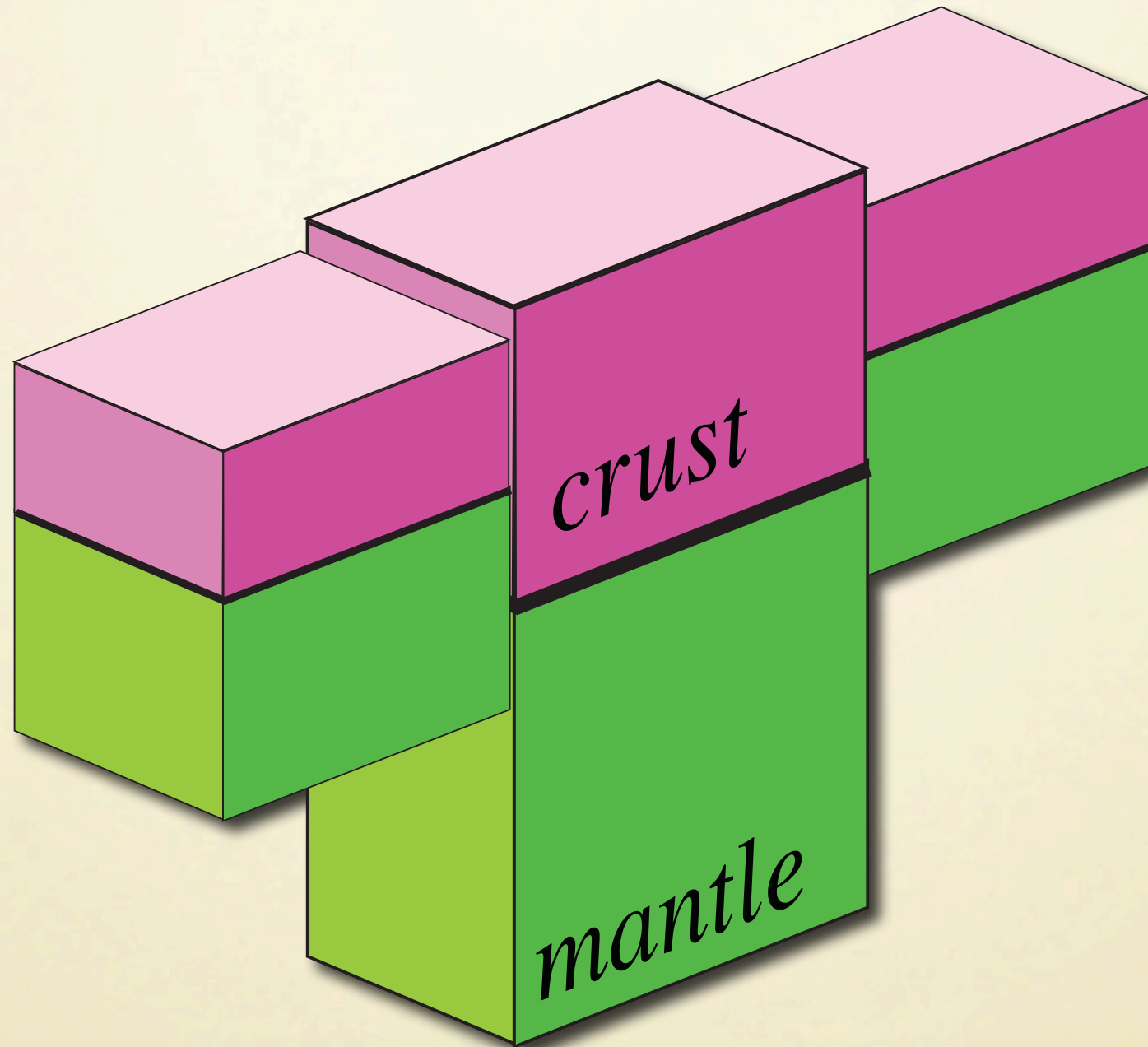


# Buoyancy of SCLM

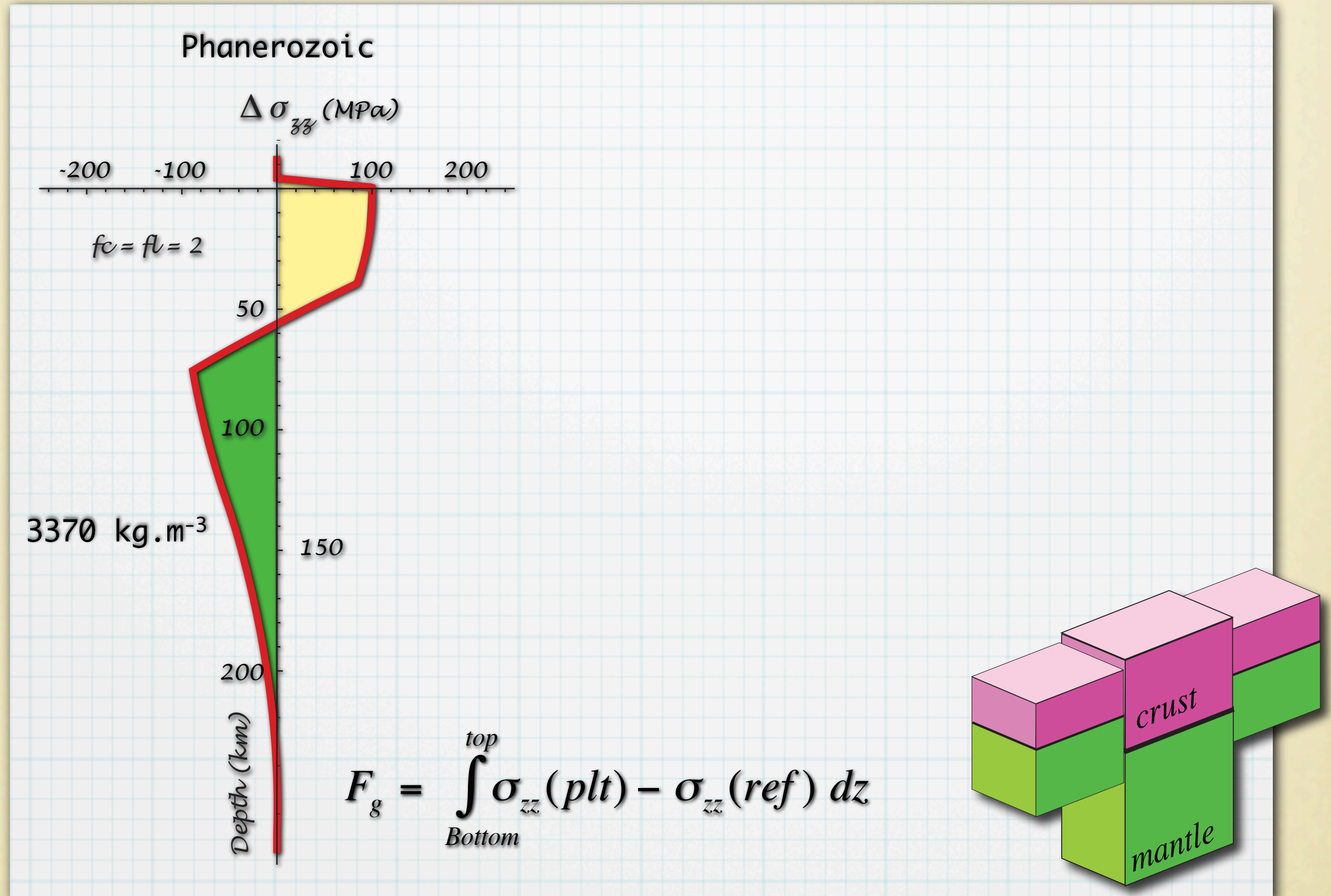




# Gravitational forces...

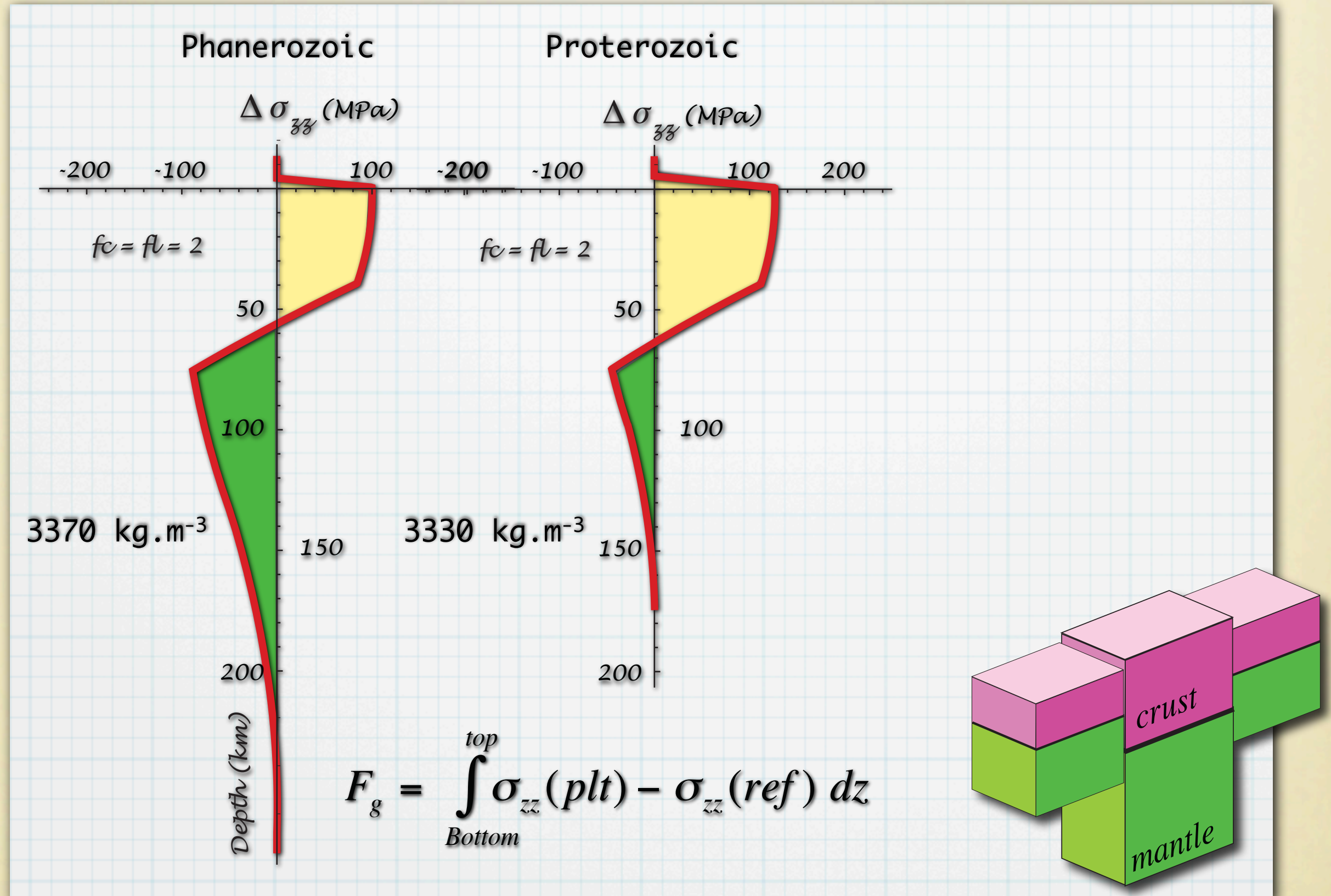


# Gravitational forces...

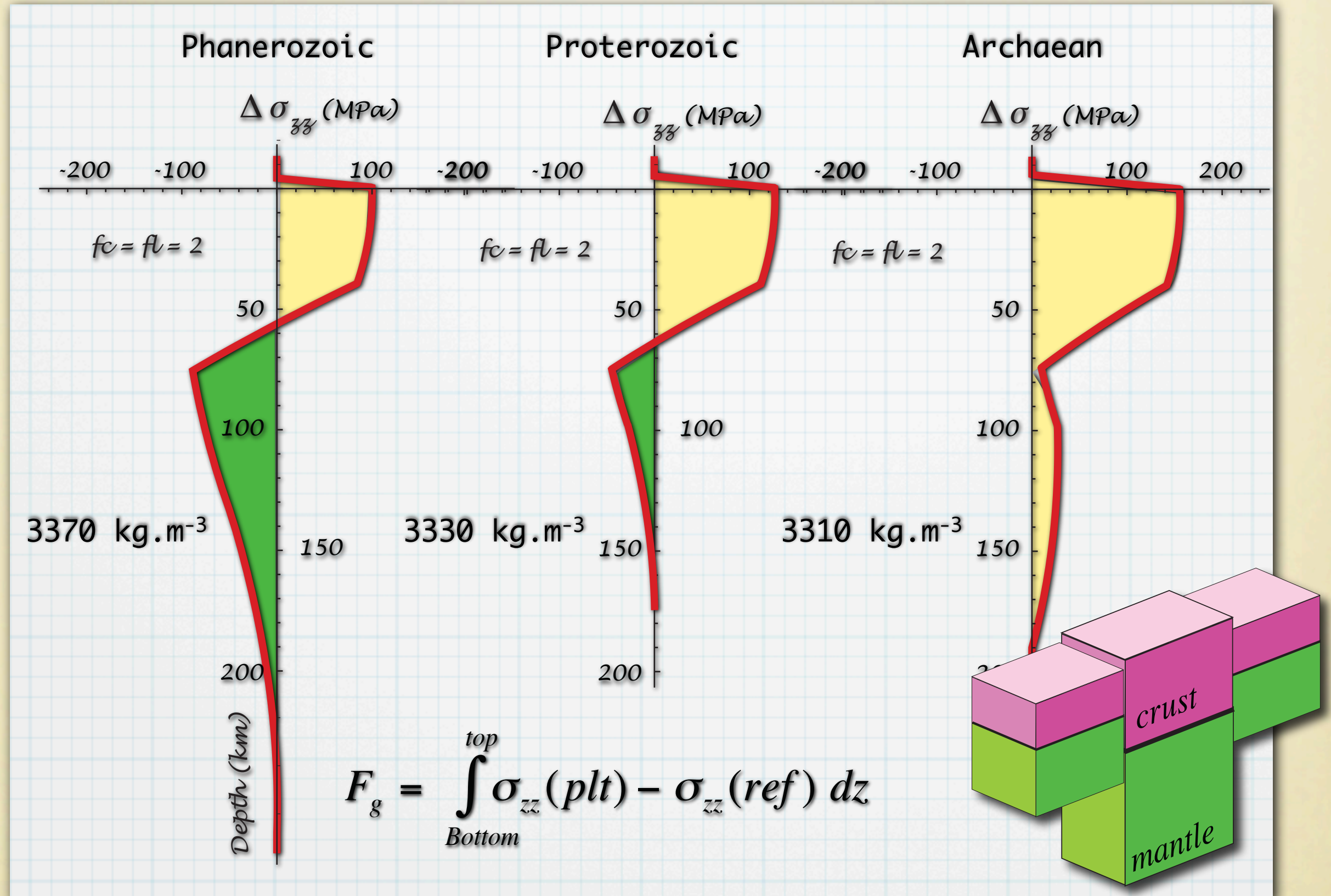




# Gravitational forces...



# Gravitational forces...





## Thin viscous sheet equations in a triaxial situation

$$\sigma_{xx} - \sigma_{yy} = 2 \eta (\dot{\epsilon}_{xx} - \dot{\epsilon}_{yy})$$

$$\sigma_{zz} - \sigma_{yy} = 2 \eta (\dot{\epsilon}_{zz} - \dot{\epsilon}_{yy})$$

$$\sigma_{zz} - \sigma_{xx} = 2 \eta (\dot{\epsilon}_{zz} - \dot{\epsilon}_{xx})$$

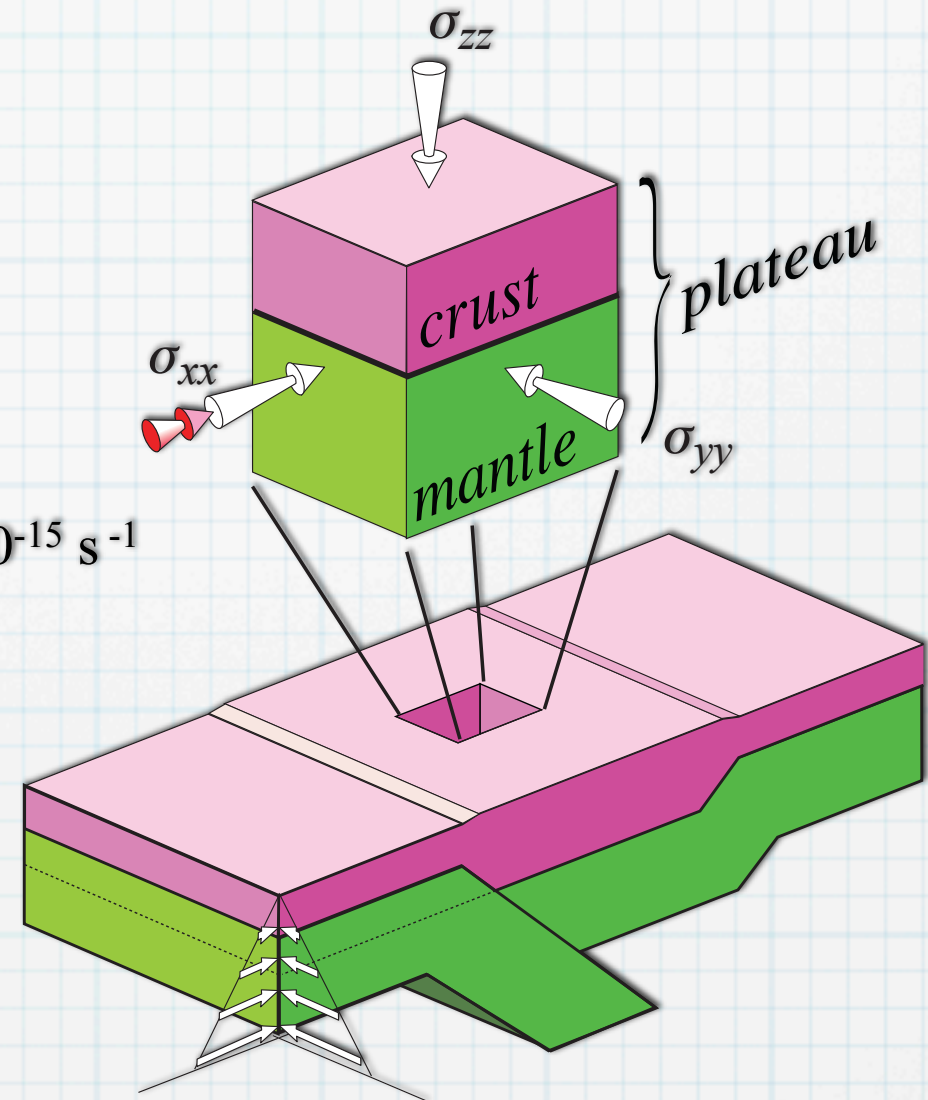
$$\dot{\epsilon}_{xx} + \dot{\epsilon}_{yy} + \dot{\epsilon}_{zz} = 0$$

$$\dot{\epsilon}_{zz} = \frac{1}{6 \eta} (2 \sigma_{zz} - \sigma_{xx} - \sigma_{yy})$$

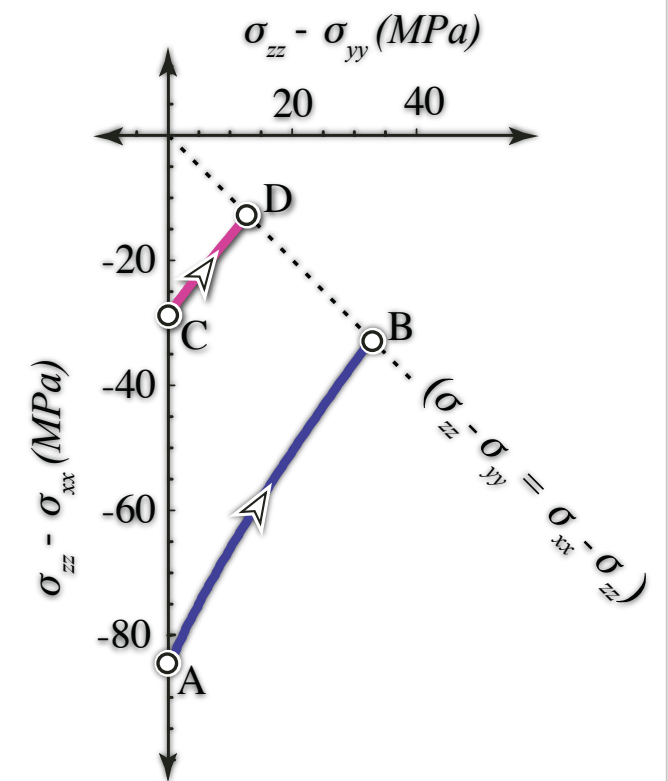
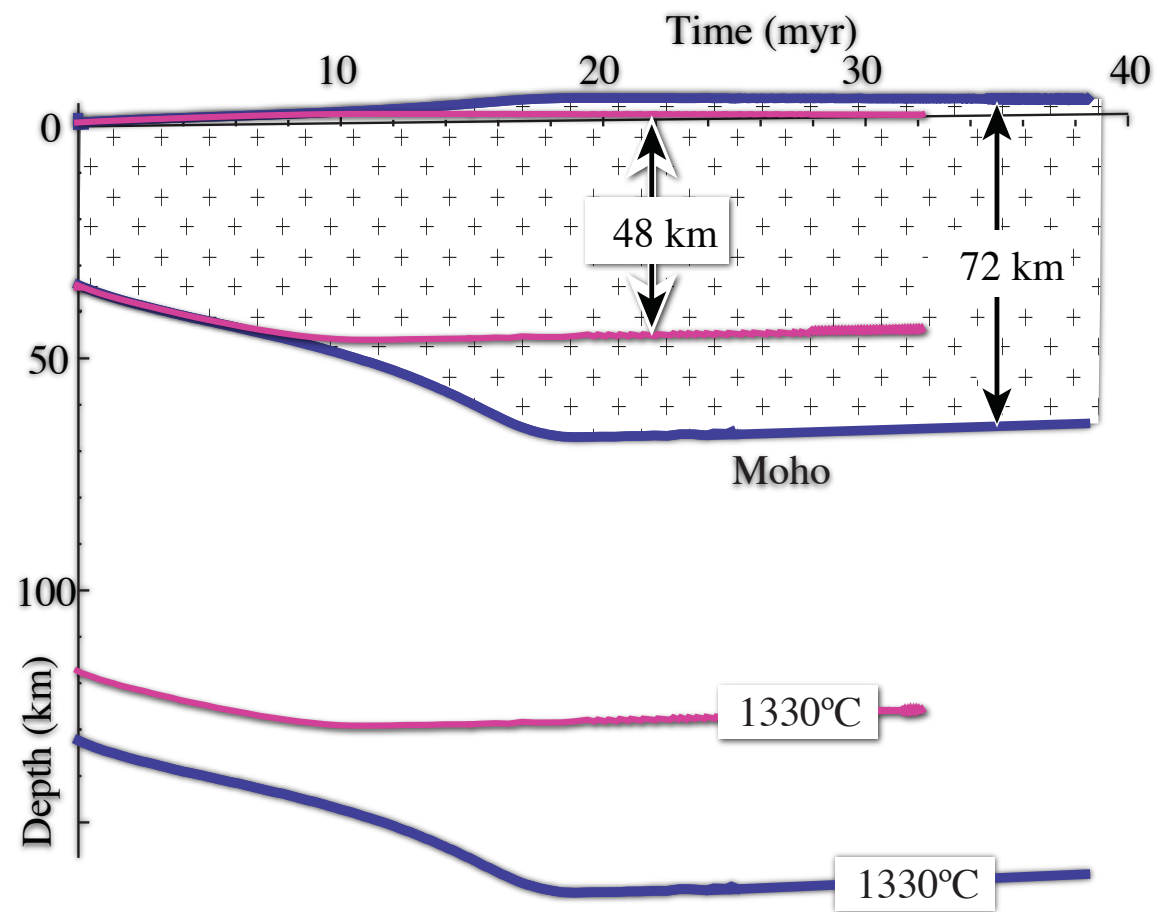
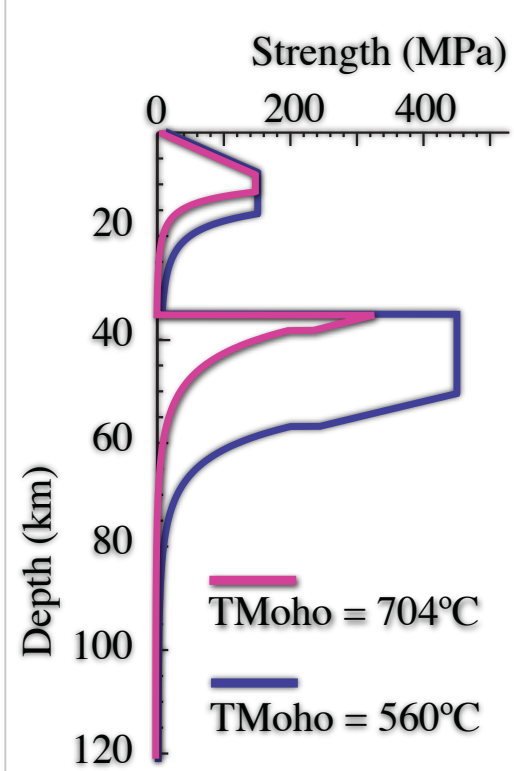
$$\dot{\epsilon}_{yy} = \frac{1}{6 \eta} (2 \sigma_{yy} - \sigma_{zz} - \sigma_{xx})$$

$$\dot{\epsilon}_{xx} = \frac{1}{6 \eta} (2 \sigma_{xx} - \sigma_{zz} - \sigma_{yy})$$

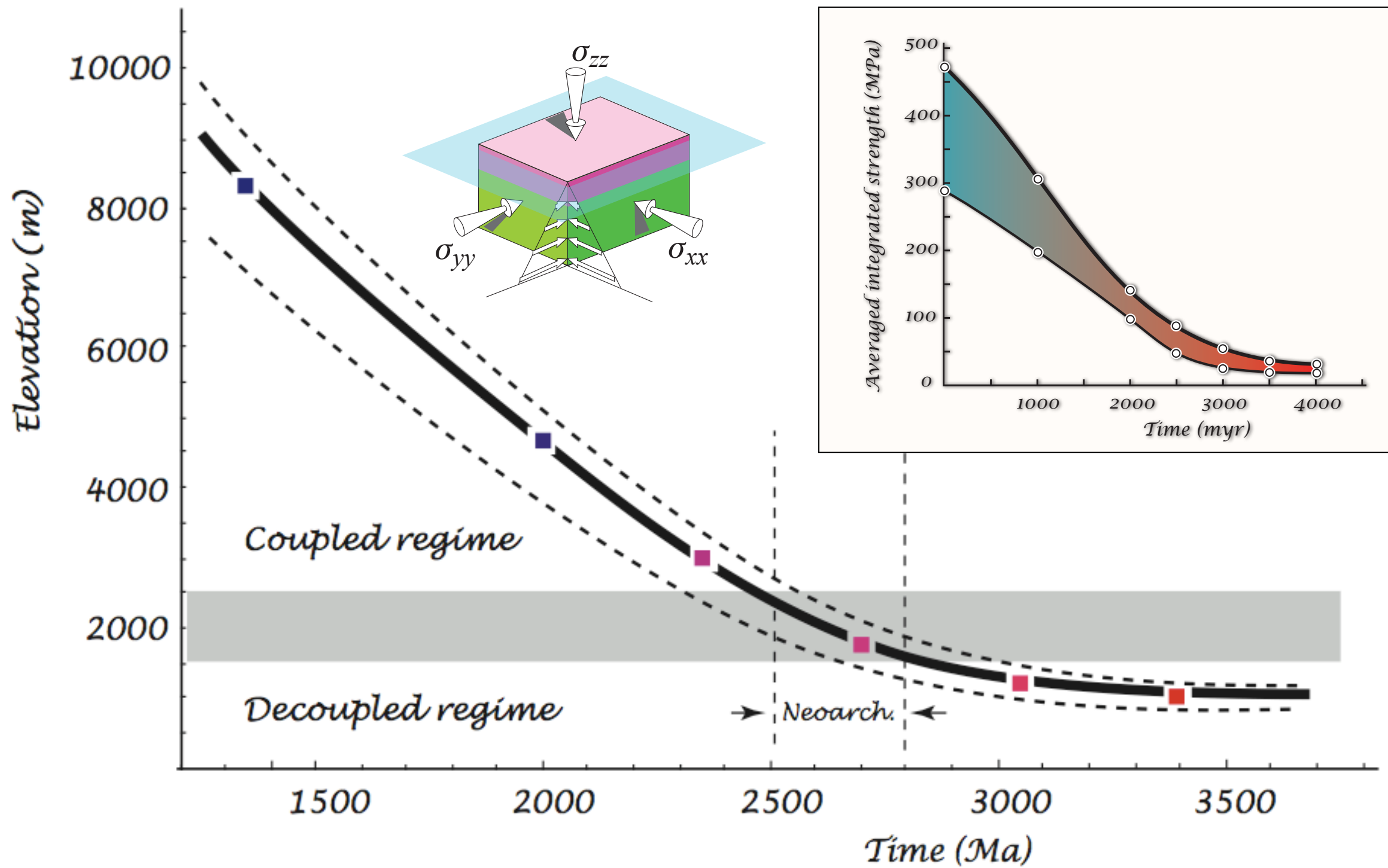
$$(\epsilon_{xx})_{init} = 5 \times 10^{-15} \text{ s}^{-1}$$



*Triaxial stress, Local isostasy, Radiogenic heating*







# The Archean Flat Earth Hypothesis

**Lithospheric scale gravitational flow: the impact of body forces on orogenic processes from Archean to Phanerozoic**

PATRICE F. REY<sup>1</sup> & GREGORY HOUSEMAN<sup>2</sup>

Geological Society, London, Special Publications, **253**, 2006.

**Orogen-parallel flow during continental convergence:  
Numerical experiments and Archean field examples**

Duclaux, Rey, Guillot, M  not

GEOLOGY, August 2007

**Neoarchean lithospheric strengthening and the  
coupling of Earth's geochemical reservoirs**

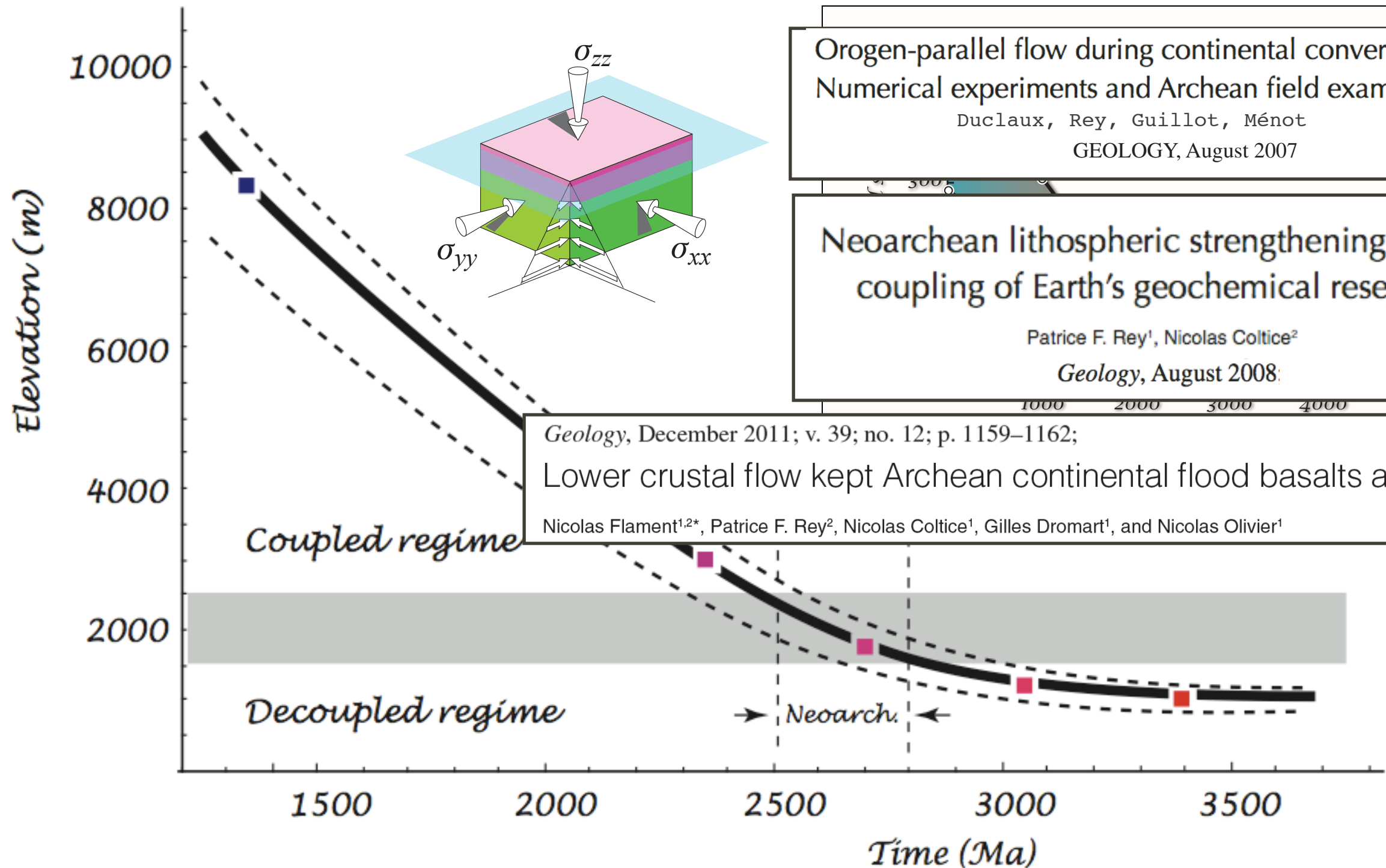
Patrice F. Rey<sup>1</sup>, Nicolas Coltice<sup>2</sup>

Geology, August 2008;

Geology, December 2011; v. 39; no. 12; p. 1159–1162;

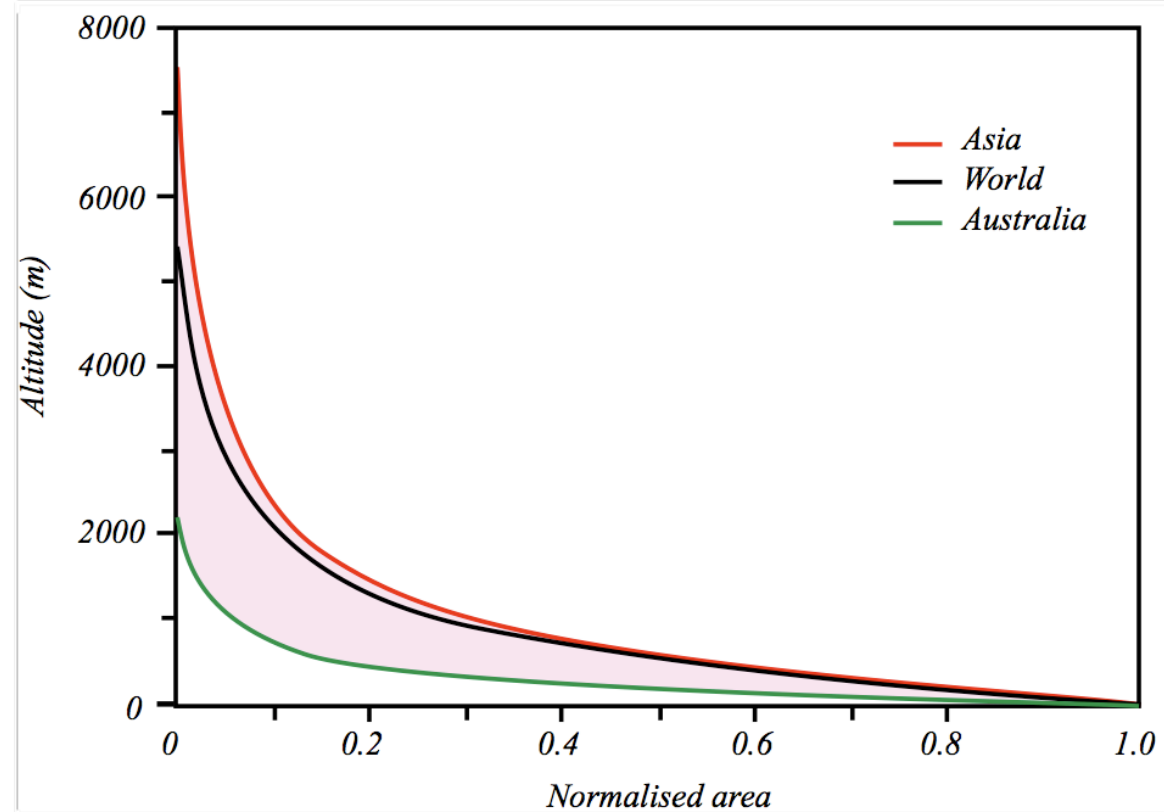
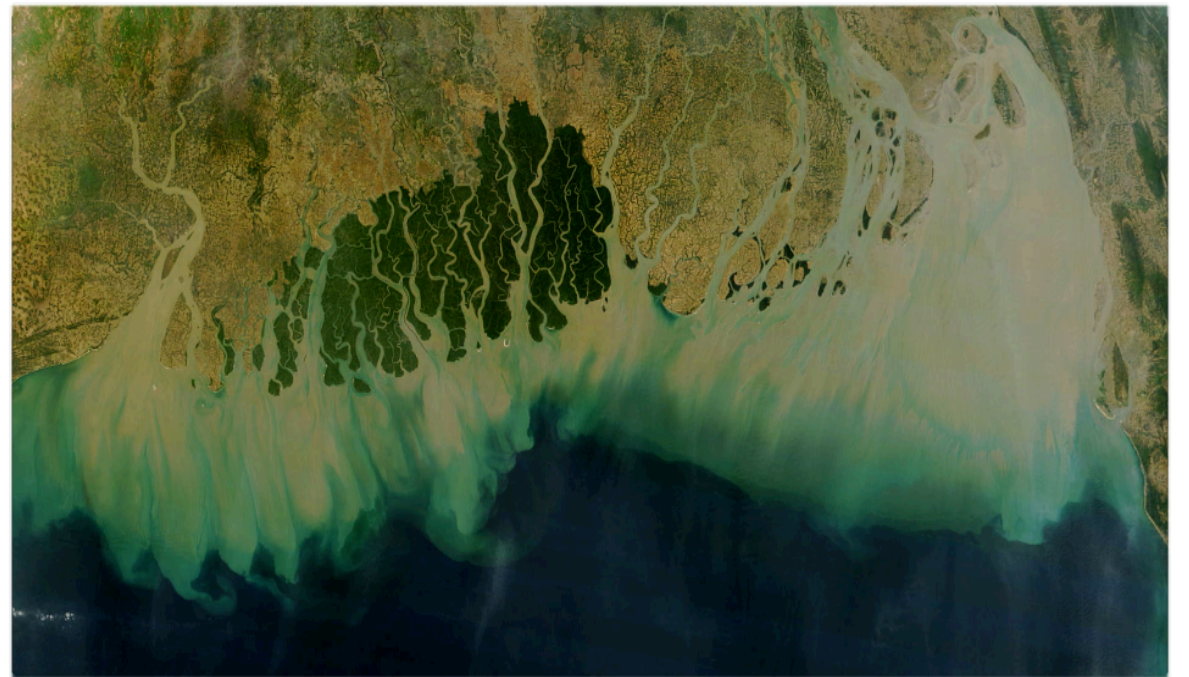
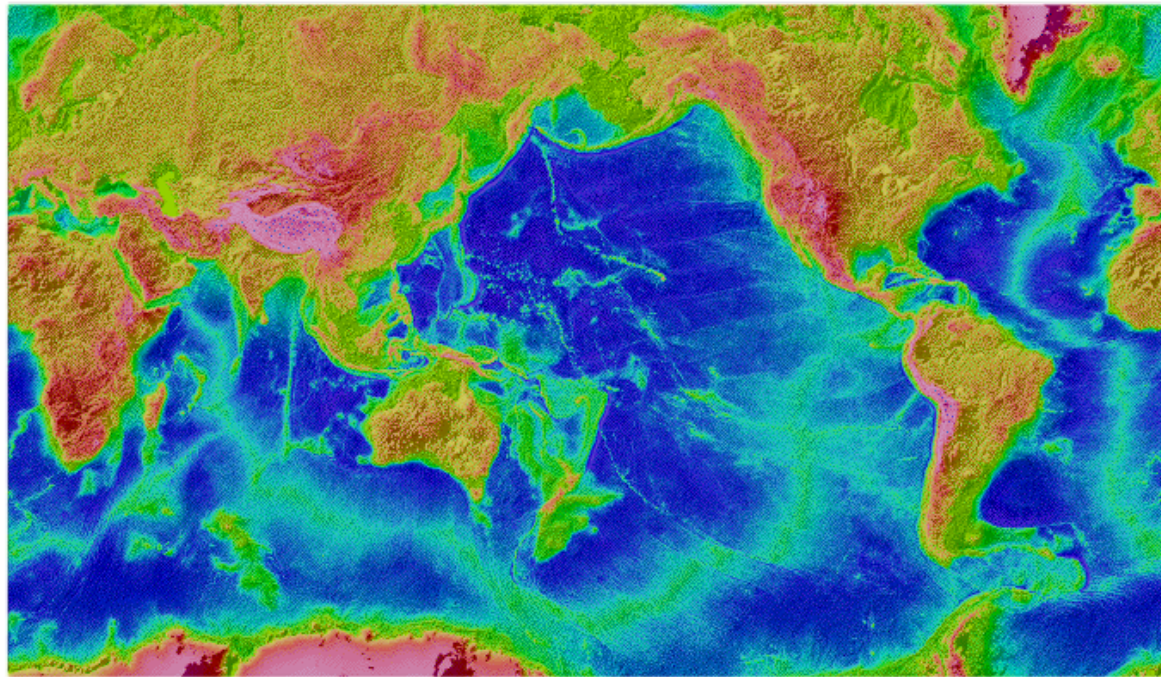
**Lower crustal flow kept Archean continental flood basalts at sea level**

Nicolas Flament<sup>1,2\*</sup>, Patrice F. Rey<sup>2</sup>, Nicolas Coltice<sup>1</sup>, Gilles Dromart<sup>1</sup>, and Nicolas Olivier<sup>1</sup>



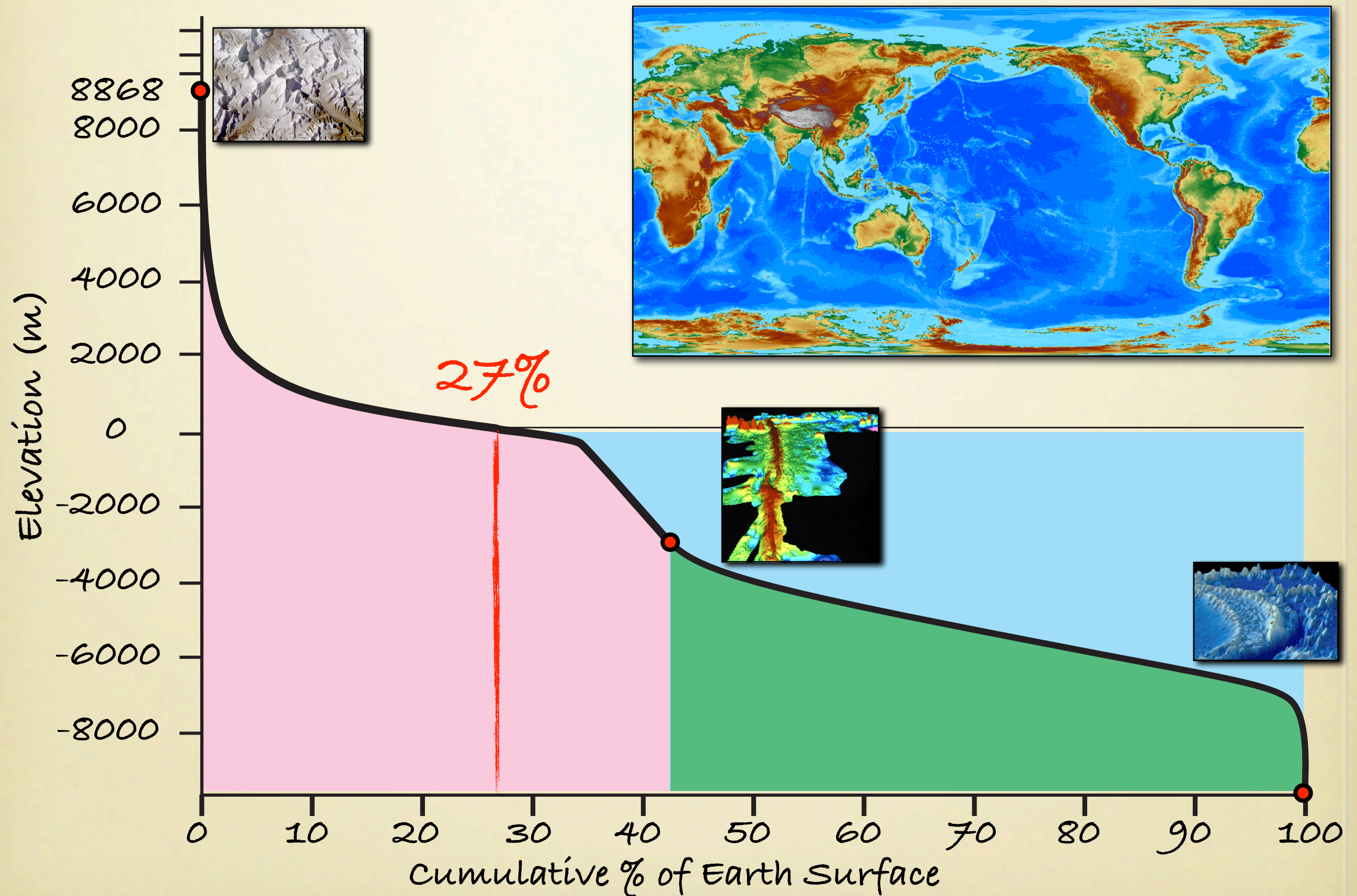


# Earth's hypsometry through times: The Archean Flat Earth's Hypothesis



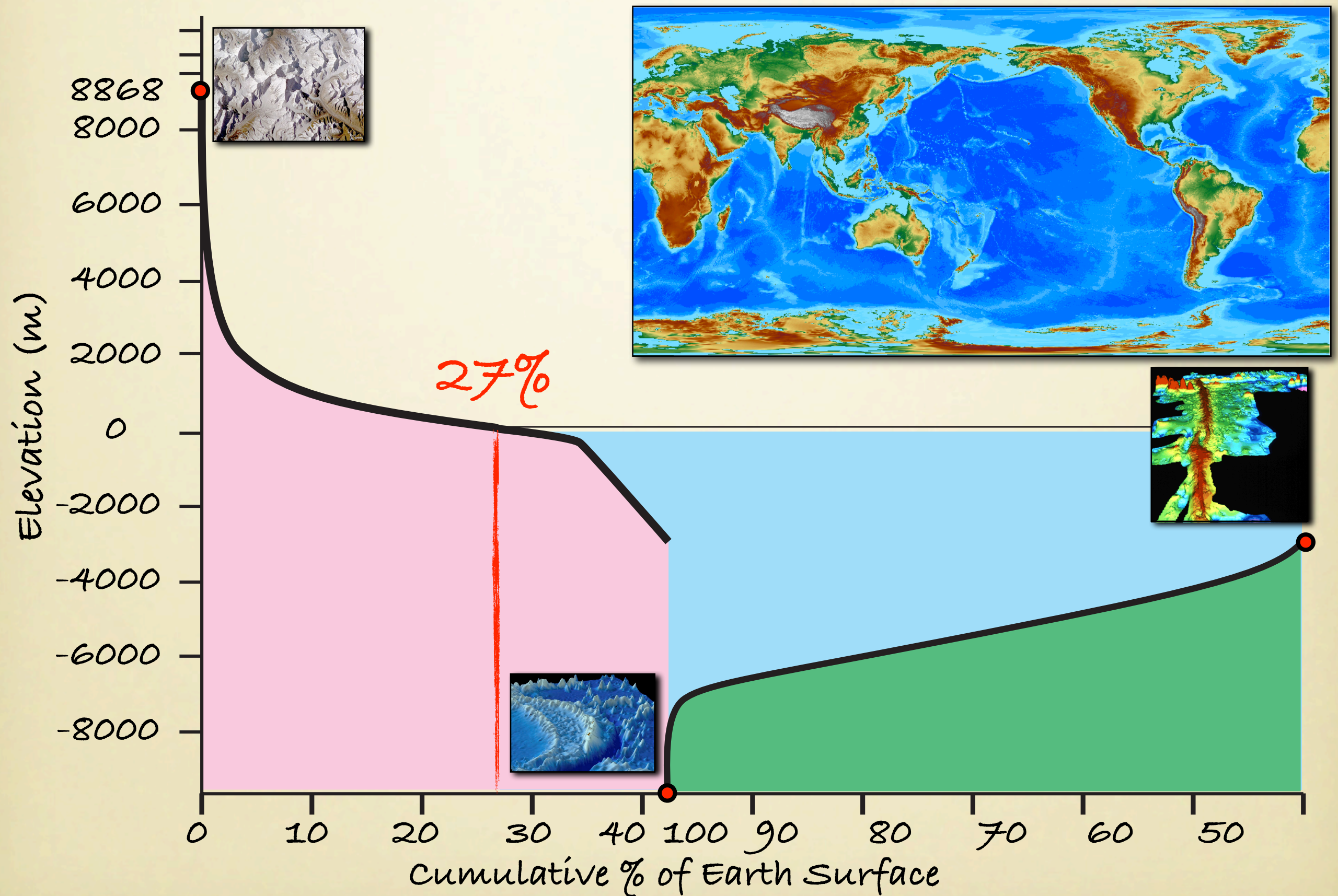


# Present-day Earth Hypsometry



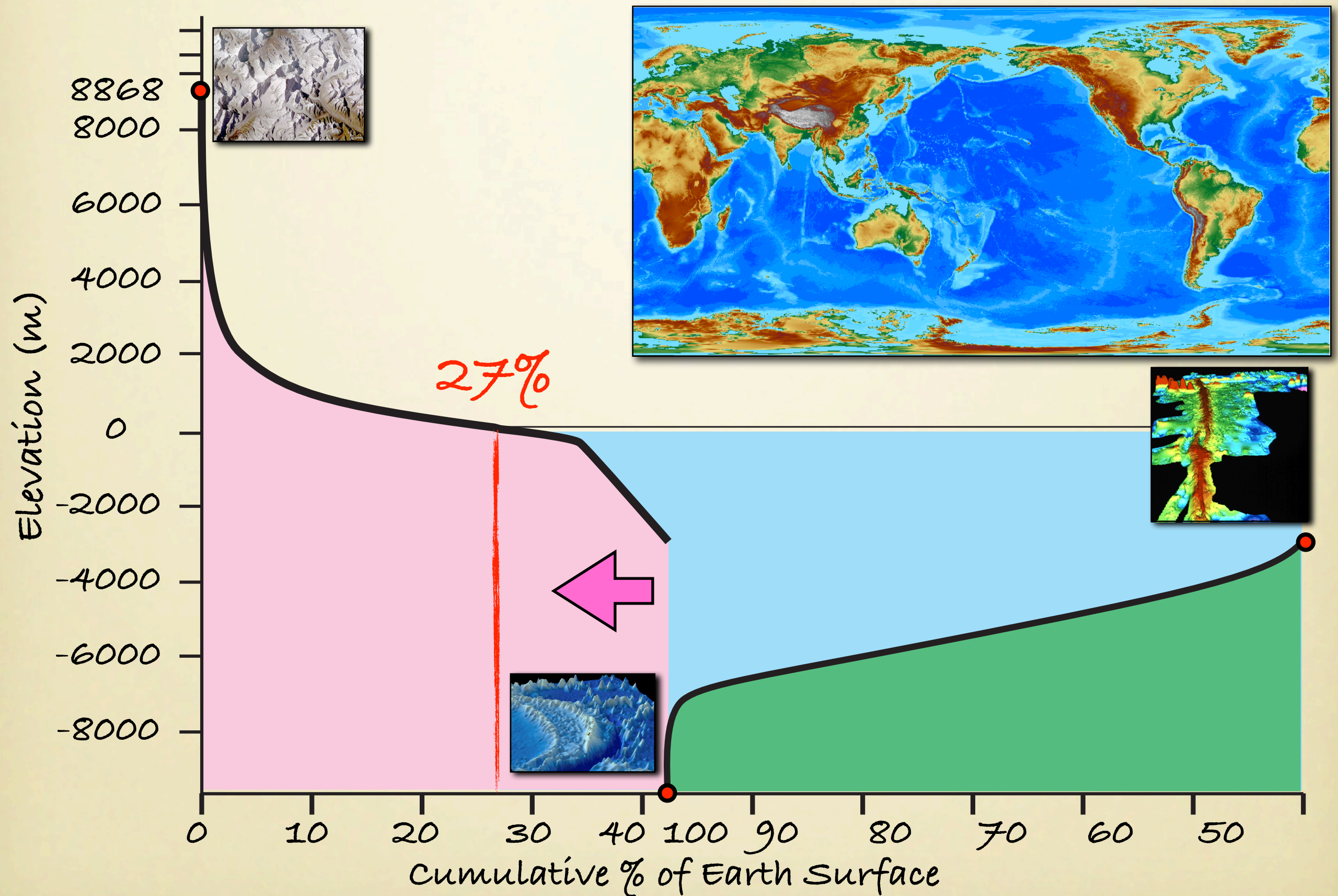


# Present-day Earth Hypsometry



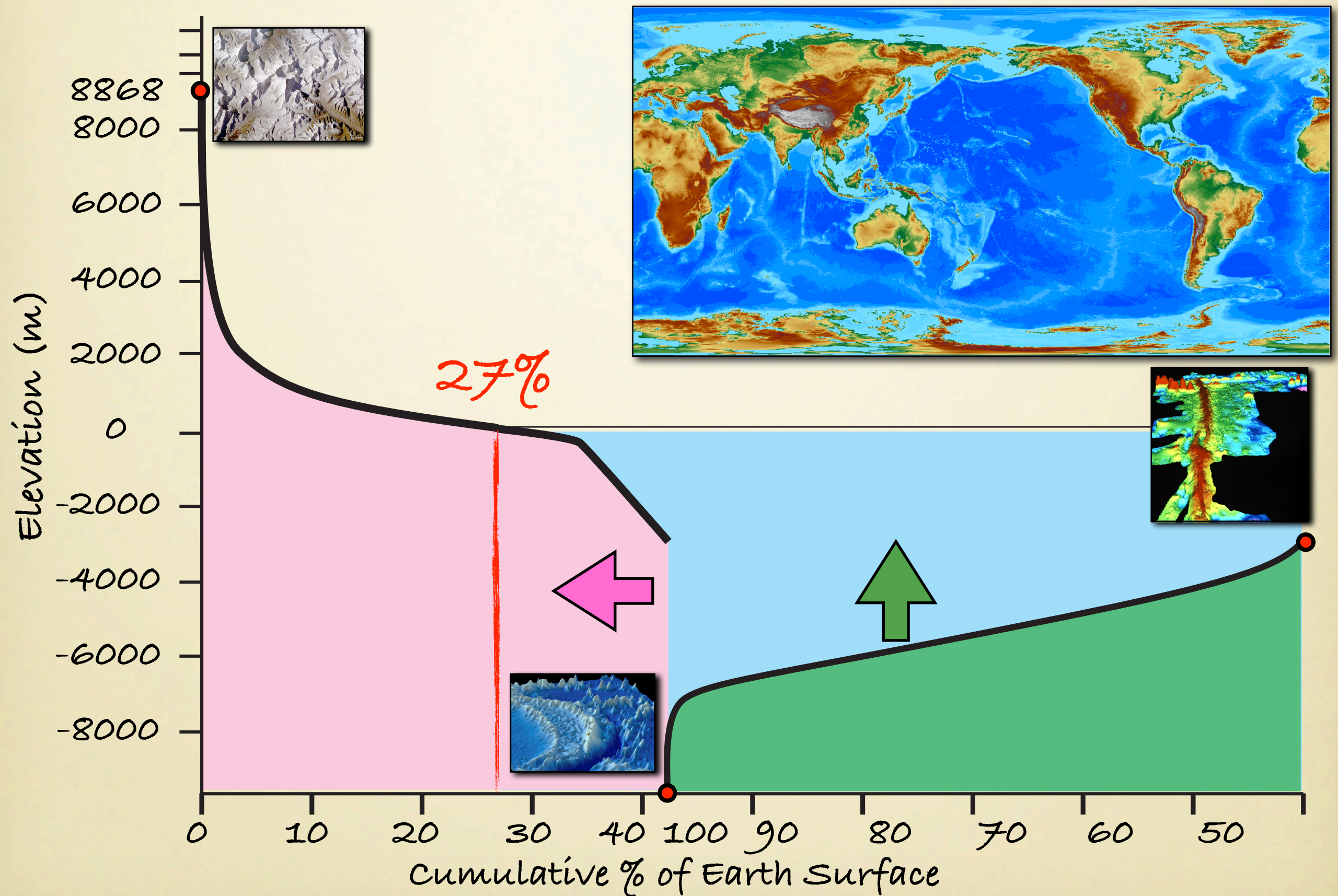


# Present-day Earth Hypsometry



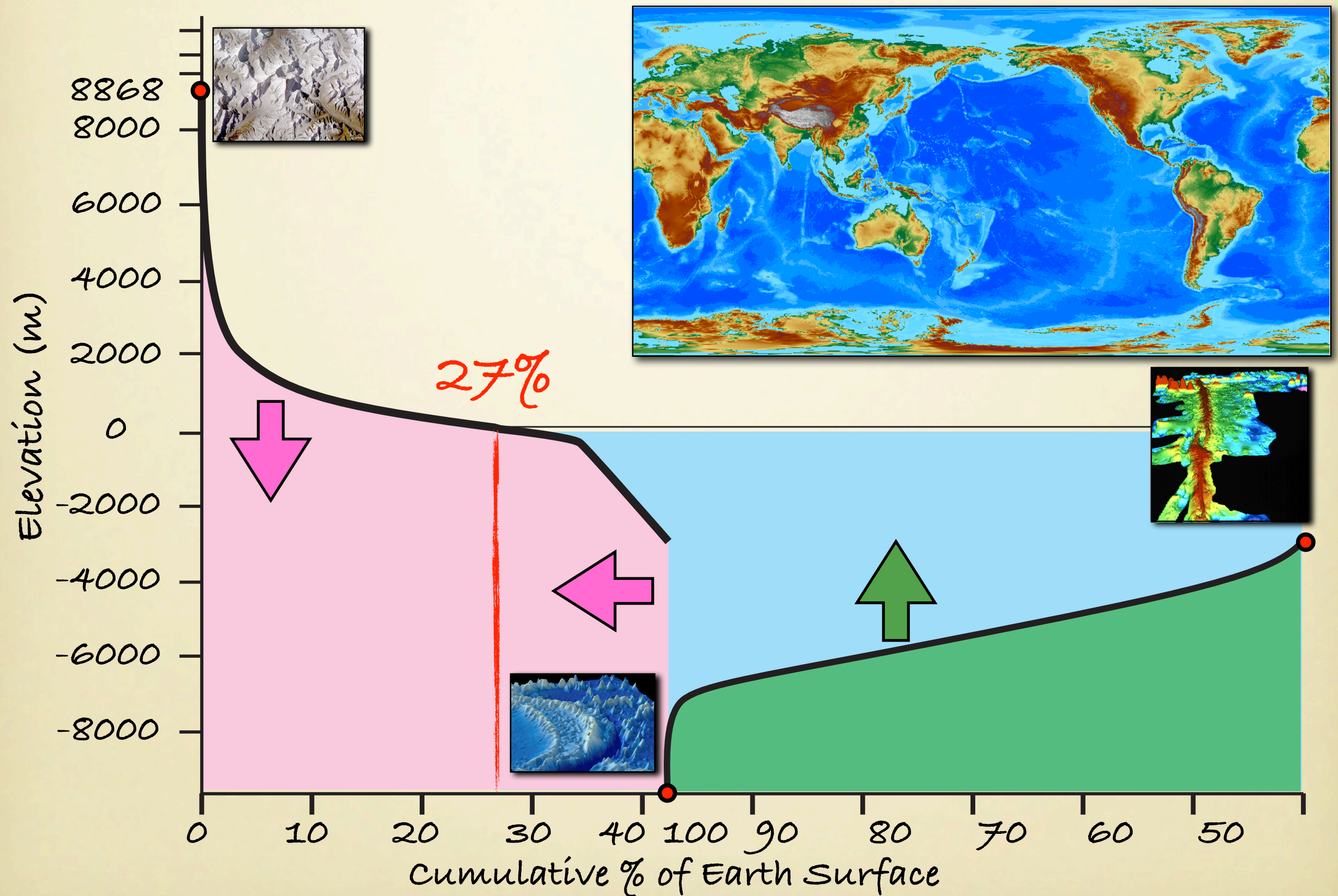


# Present-day Earth Hypsometry



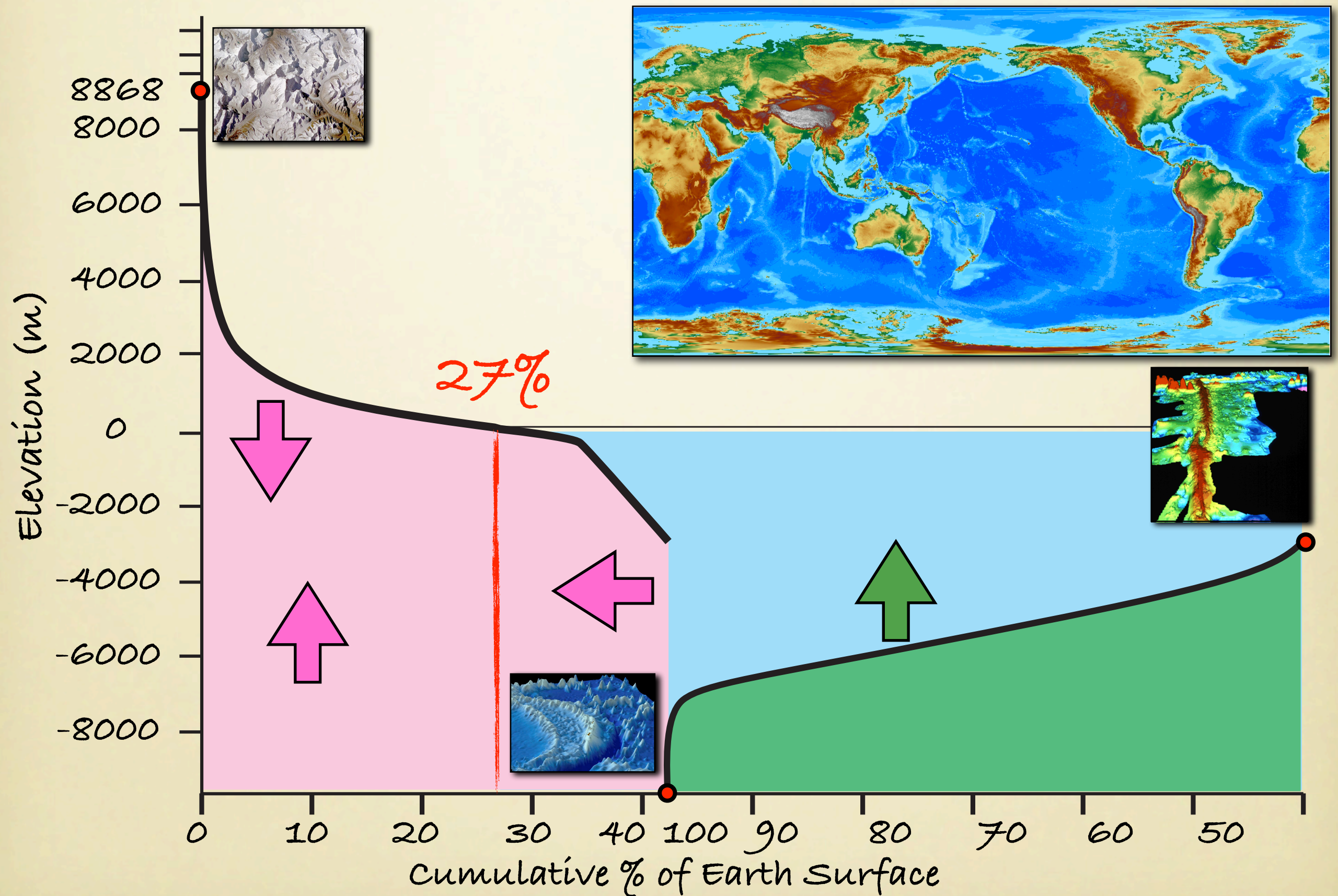


# Present-day Earth Hypsometry

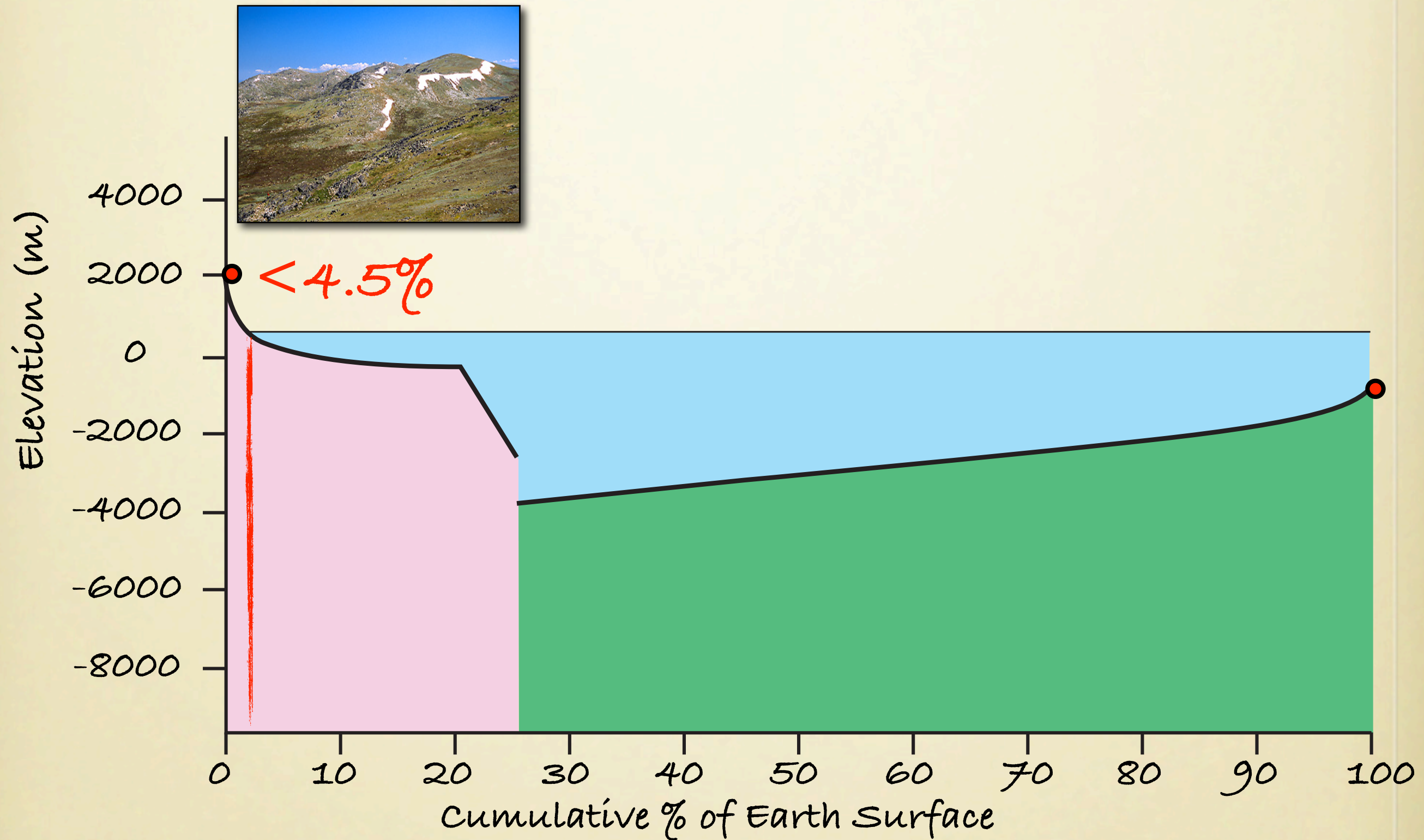




# Present-day Earth Hypsometry

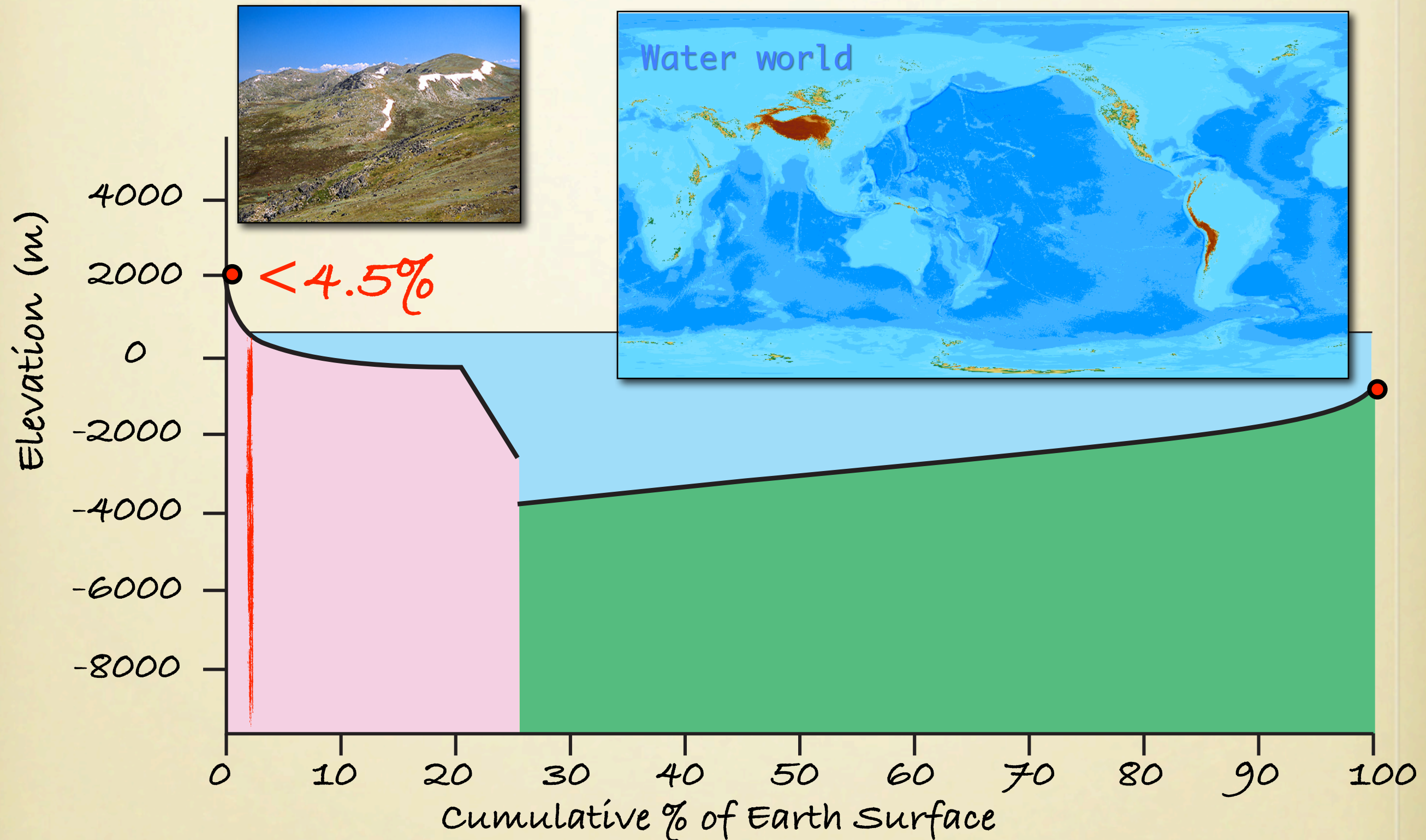


# Late Archean Hypsometry: The Archean Water World Hypothesis

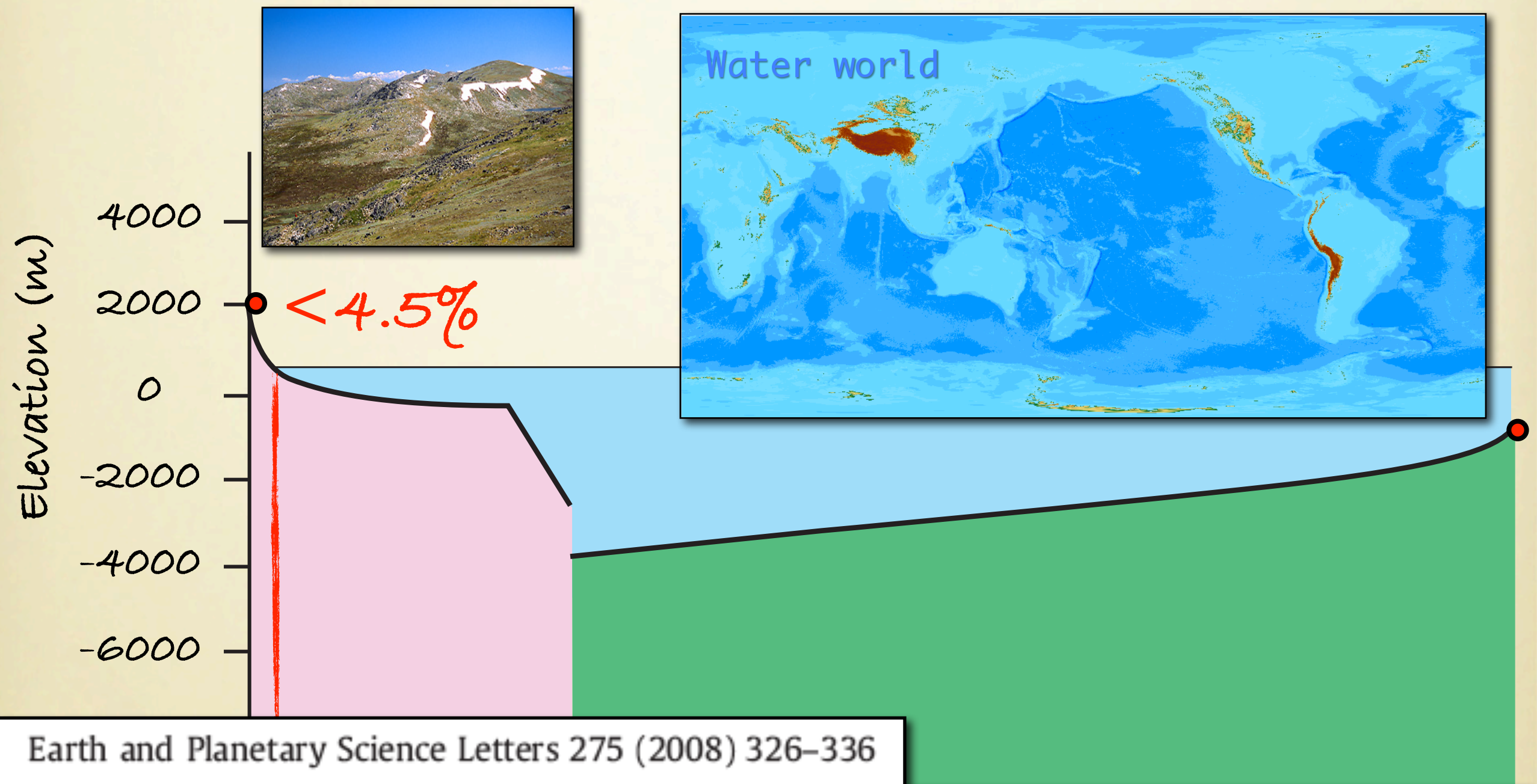




# Late Archean Hypsometry: The Archean Water World Hypothesis



# Late Archean Hypsometry: The Archean Water World Hypothesis

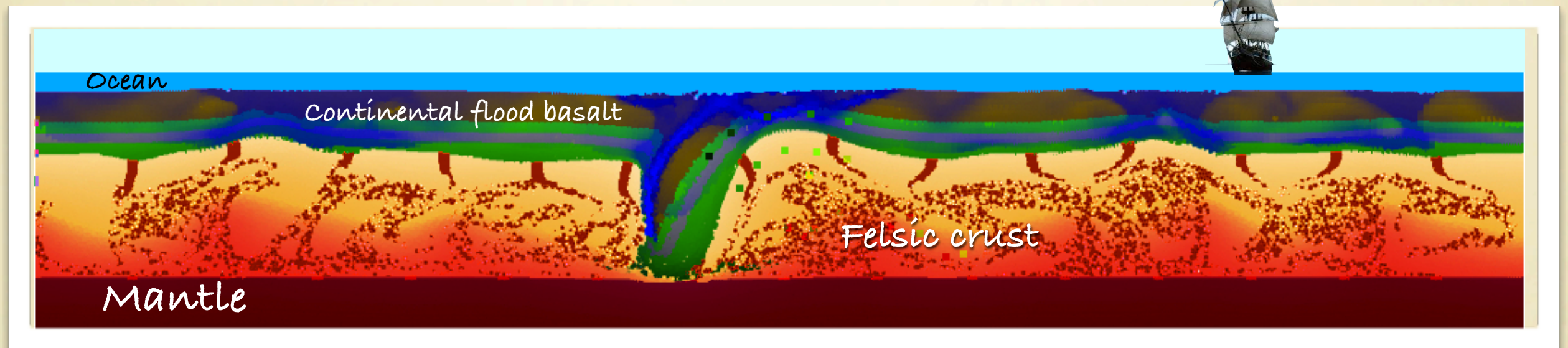


A case for late-Archaean continental emergence from thermal evolution models and hypsometry

Nicolas Flament<sup>a,b,\*</sup>, Nicolas Coltice<sup>a</sup>, Patrice F. Rey<sup>b</sup>

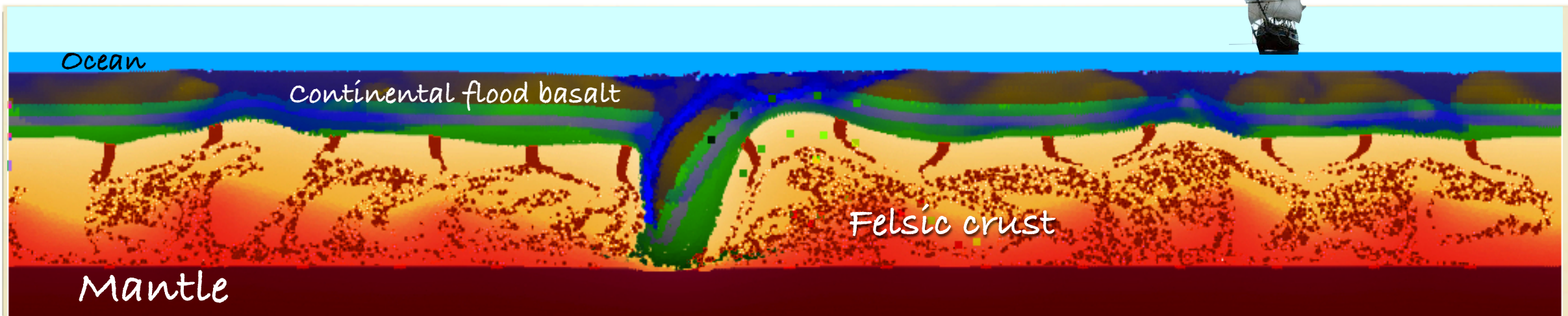


# CONSEQUENCES FOR ECONOMIC GEOLOGY





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Precambrian Research 229 (2013) 93–104



Contents lists available at SciVerse ScienceDirect

Precambrian Research

journal homepage: [www.elsevier.com/locate/precamres](http://www.elsevier.com/locate/precamres)

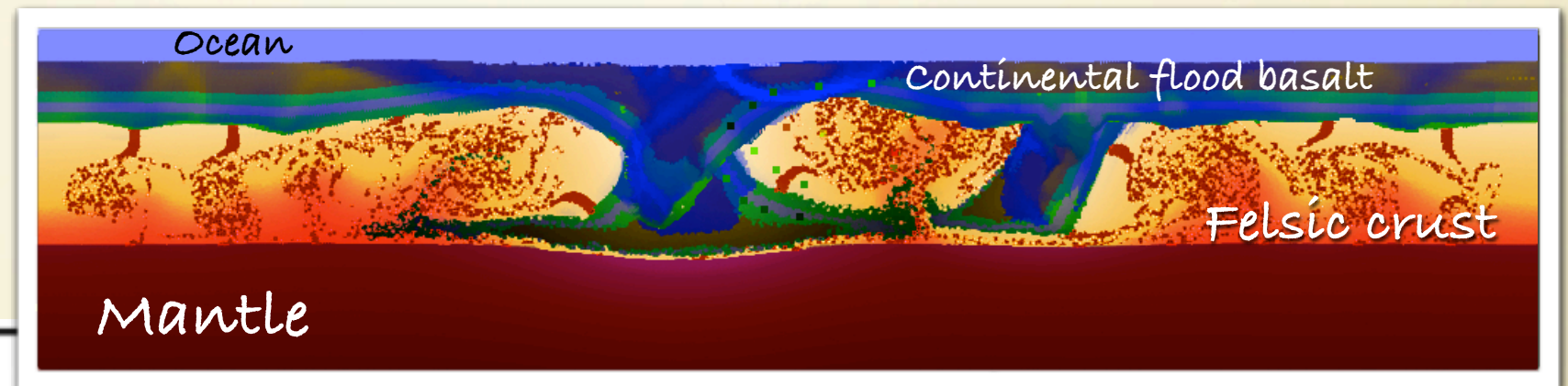


Archean gravity-driven tectonics on hot and flooded continents: Controls on long-lived mineralised hydrothermal systems away from continental margins

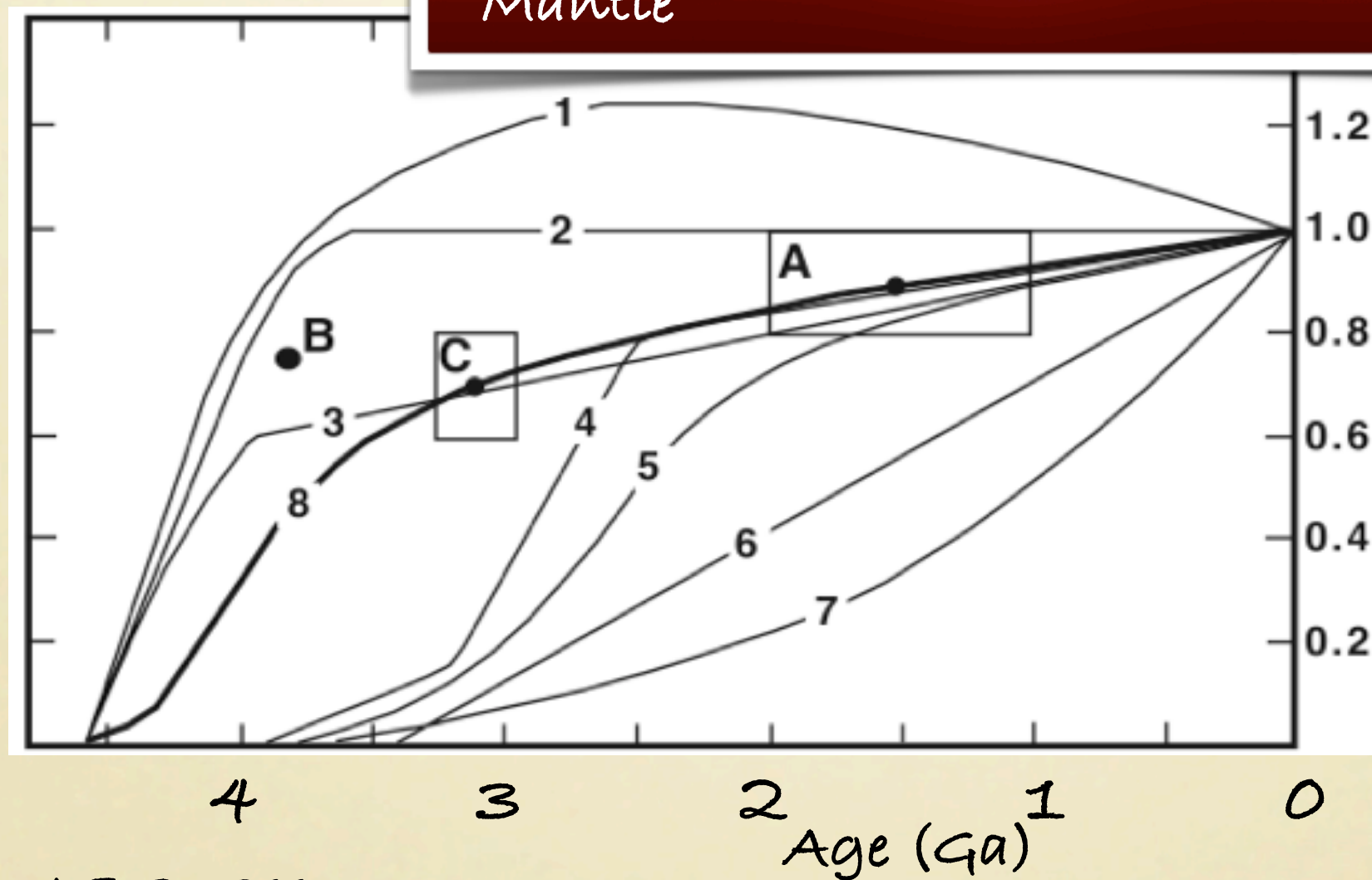
N. Thébaud<sup>a,\*</sup>, P.F. Rey<sup>b</sup>



# CONSEQUENCES FOR EVERYONE ELSE

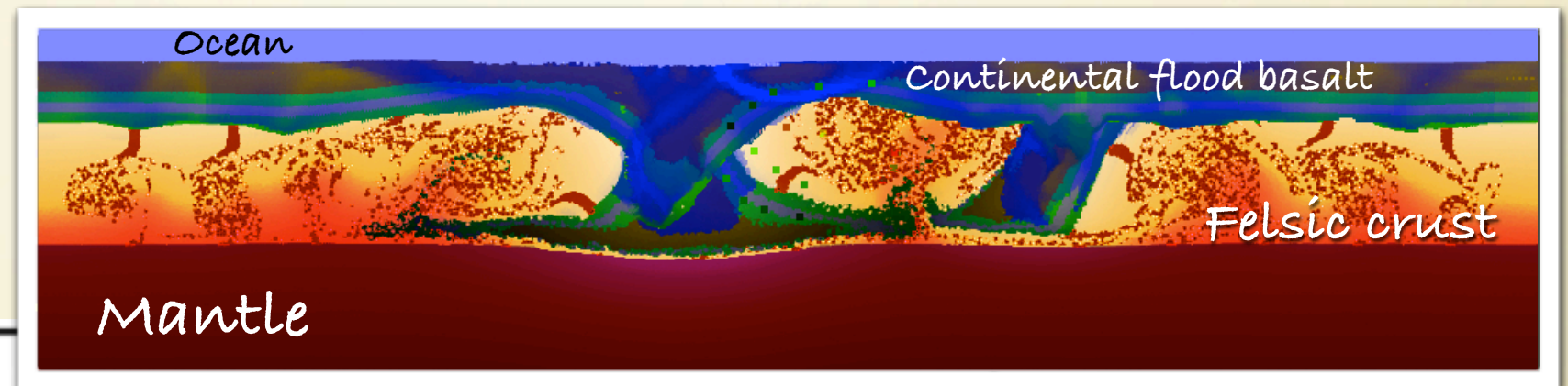


Crust volume  
(% present day)

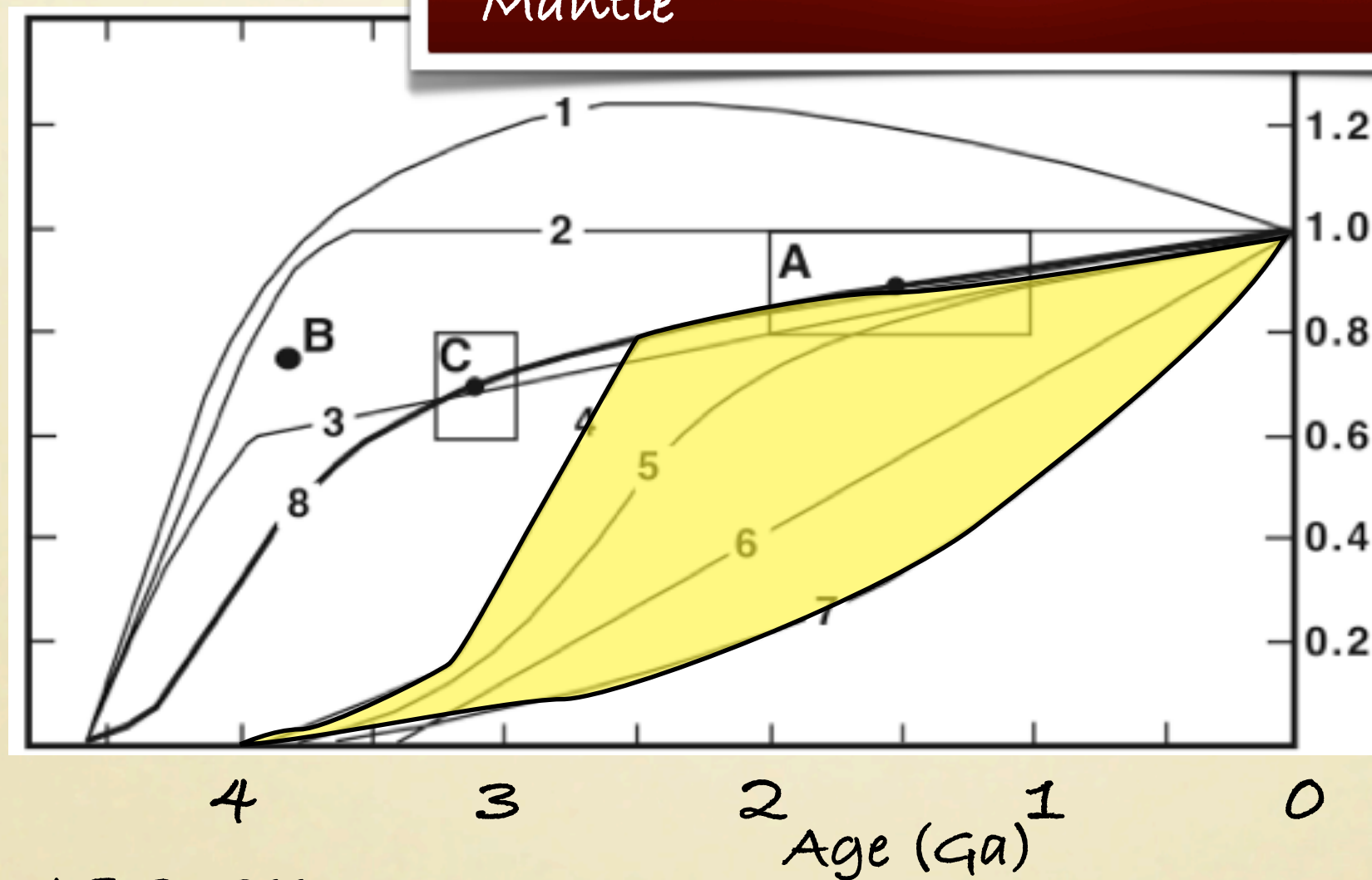


Campbell, Am. J. E. Sc., 2003

# CONSEQUENCES FOR EVERYONE ELSE



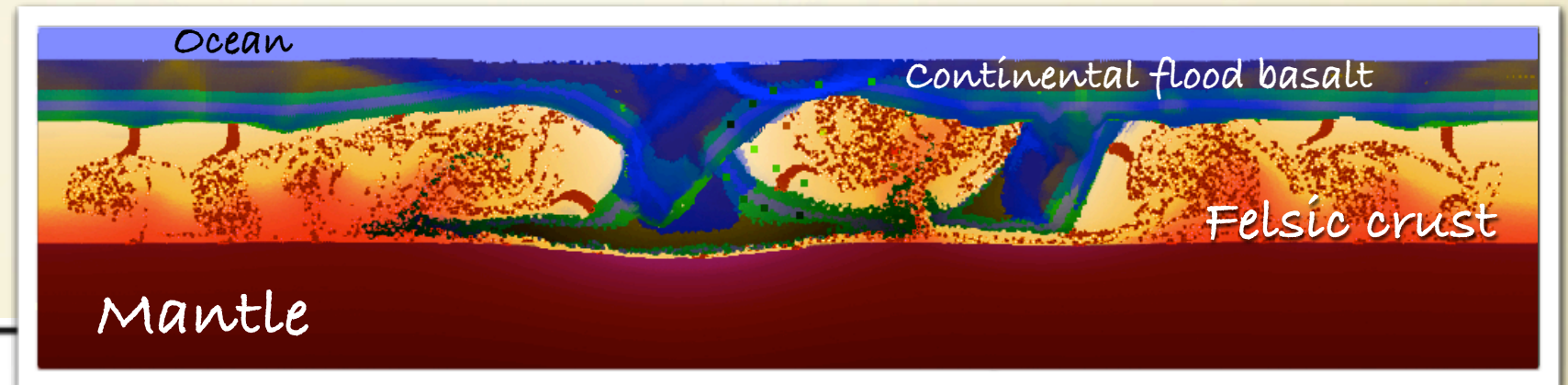
Crust volume  
(% present day)



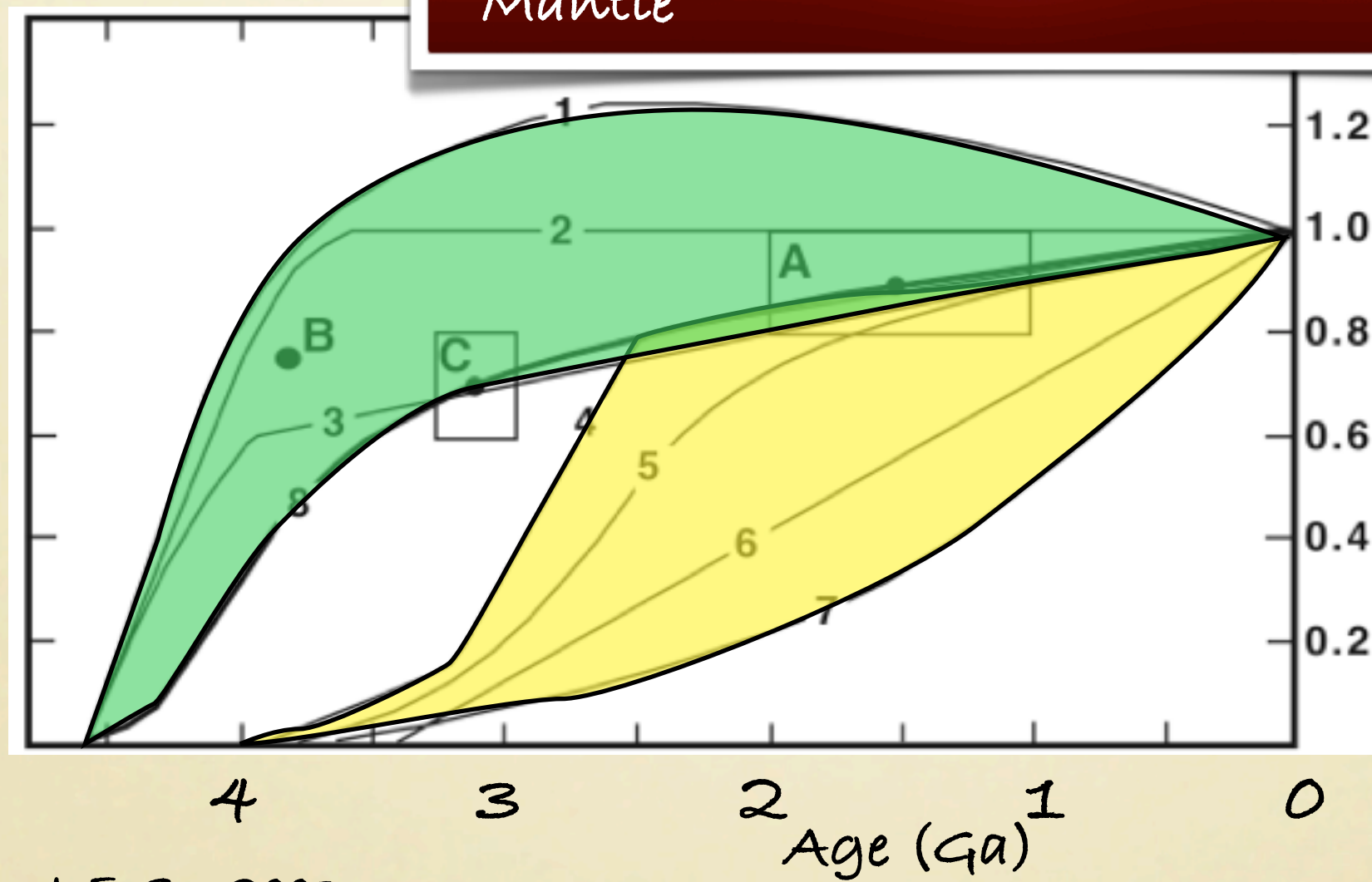
Campbell, Am. J. E. Sc., 2003



# CONSEQUENCES FOR EVERYONE ELSE

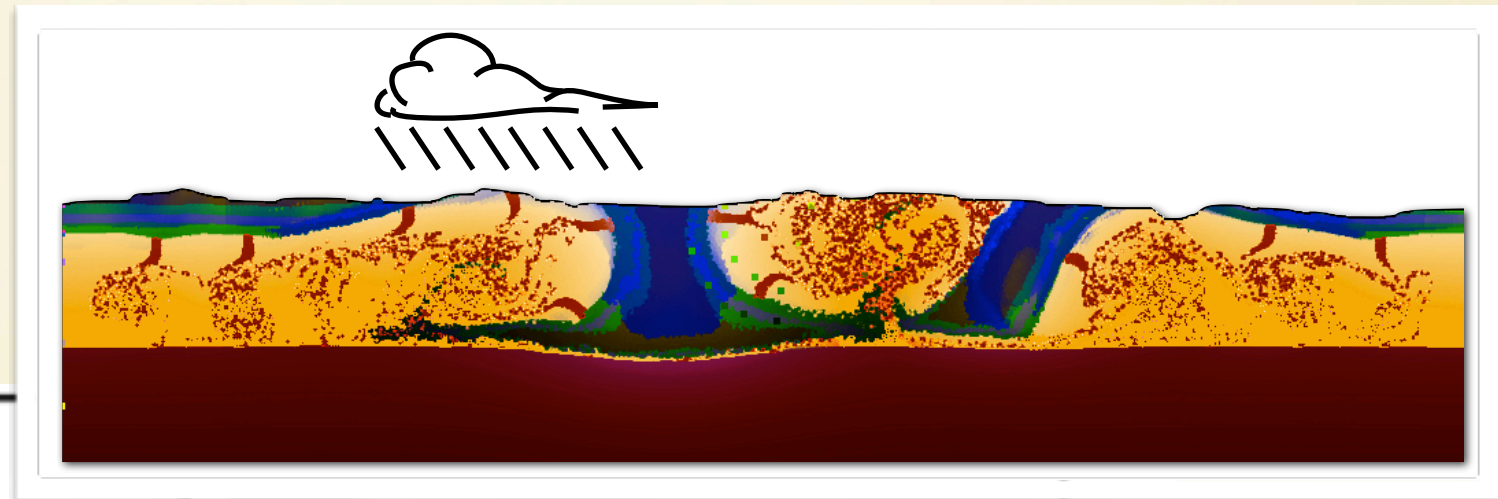


Crust volume  
(% present day)

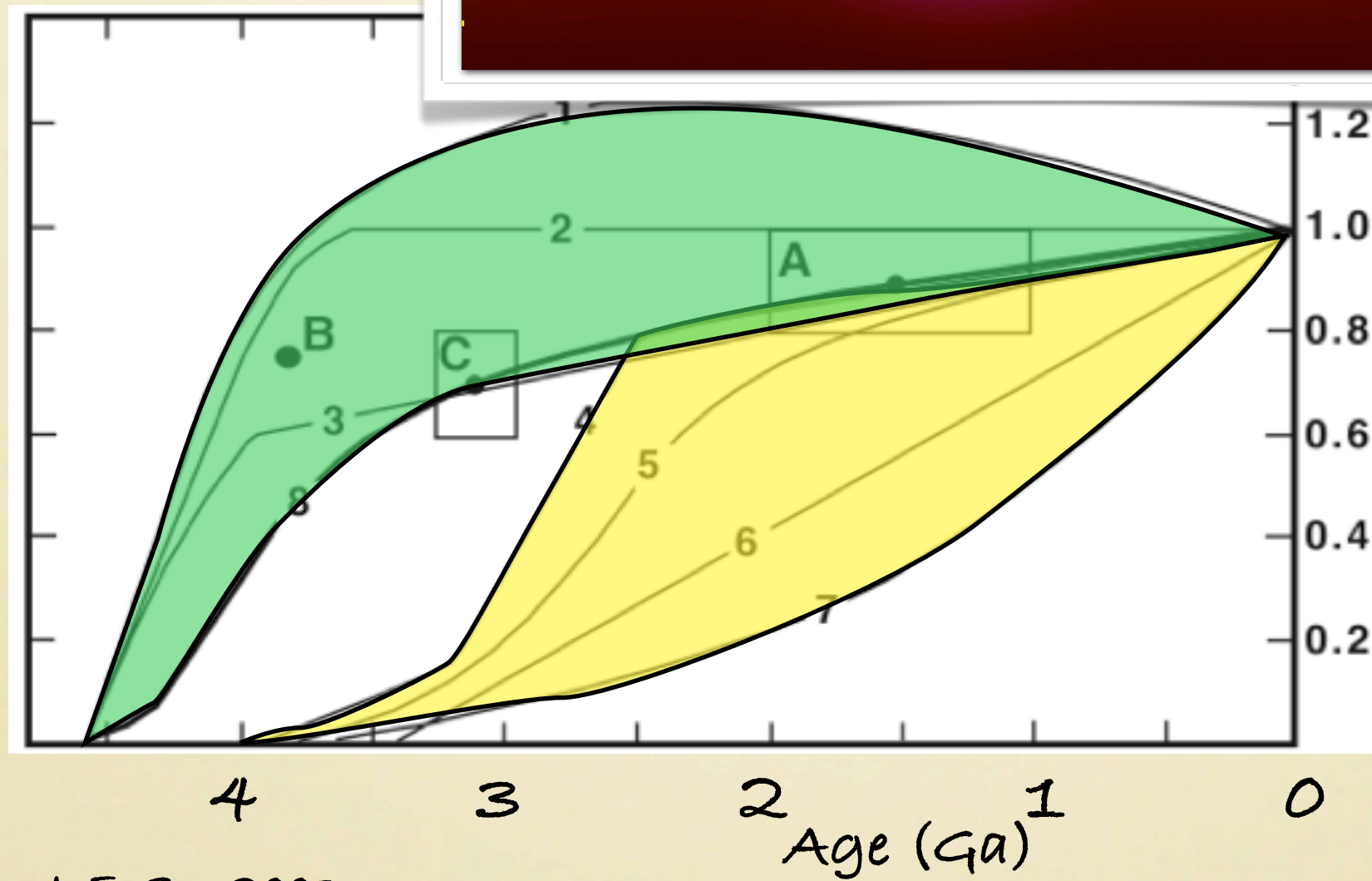


Campbell, Am. J. E. Sc., 2003

# CONSEQUENCES FOR EVERYONE ELSE



Crust volume  
(% present day)



Campbell, Am. J. E. Sc., 2003



# **HOT CONTINENTS CHANGE EVERYTHING...**

# HOT CONTINENTS CHANGE EVERYTHING...

HYPSOMETRY  
&  
TECTONICS



# HOT CONTINENTS CHANGE EVERYTHING...

HYPSONOMETRY  
&  
TECTONICS



EARTH'S GEOCHEM



BIOSPHERE

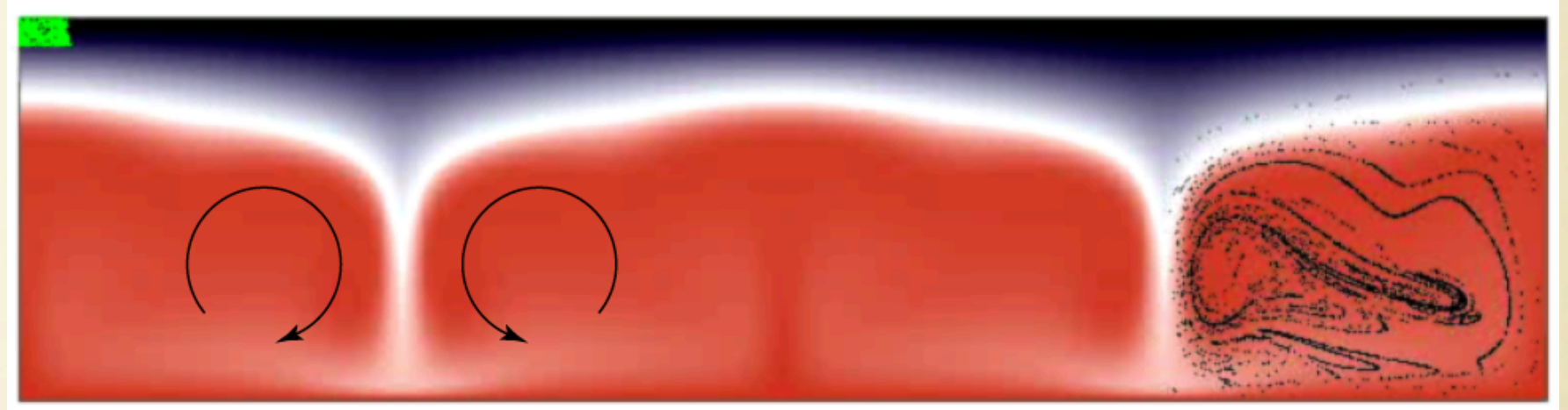


ECONOMIC GEOLOGY

# EARLY CONTINENTS AND PLATE TECTONICS

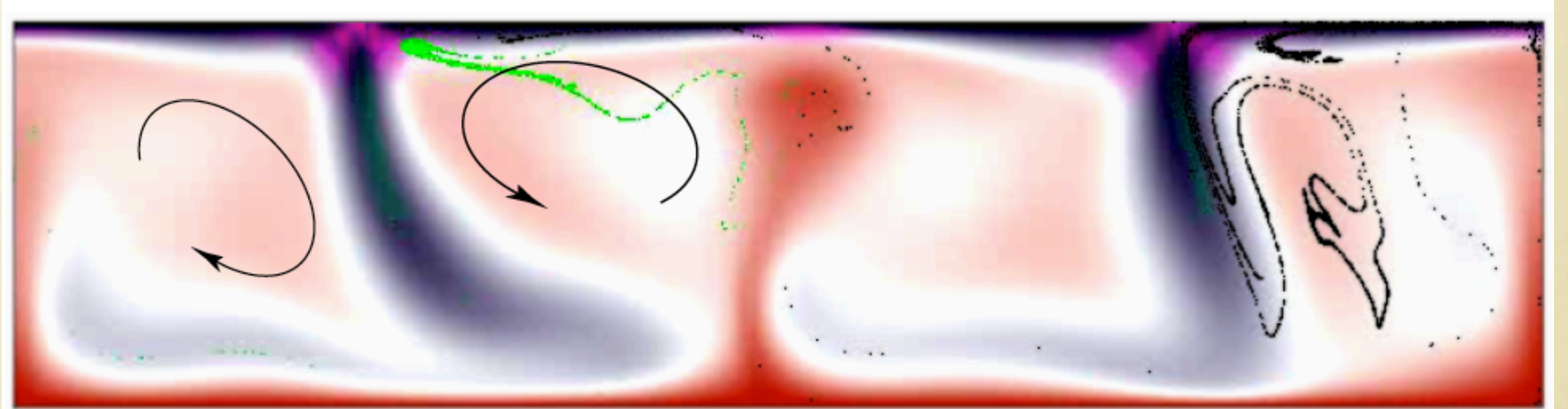
## STAGNANT LID

Convective stress  
<  
Yield stress



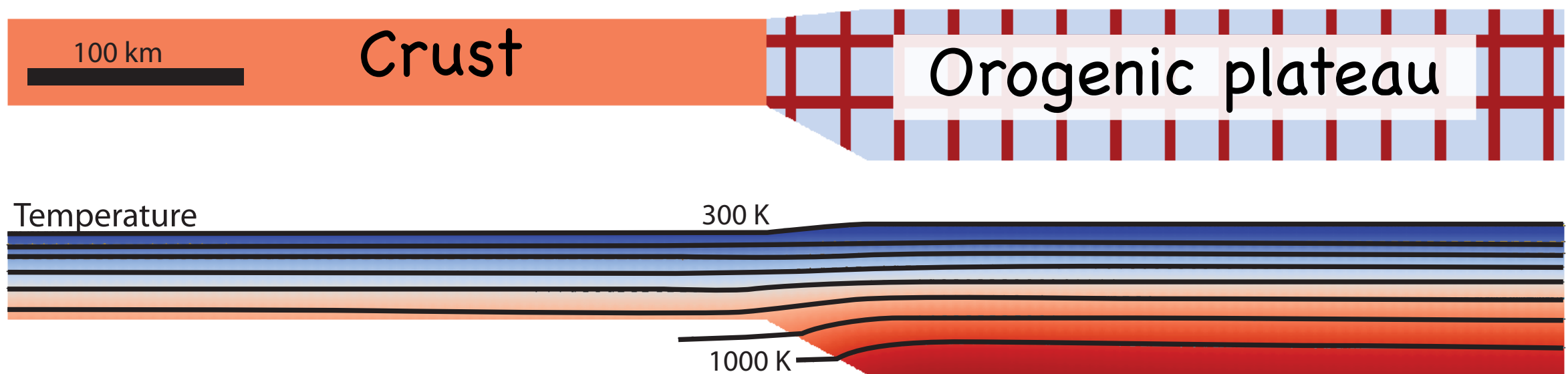
## MOBILE LID

Convective stress  
>  
Yield stress

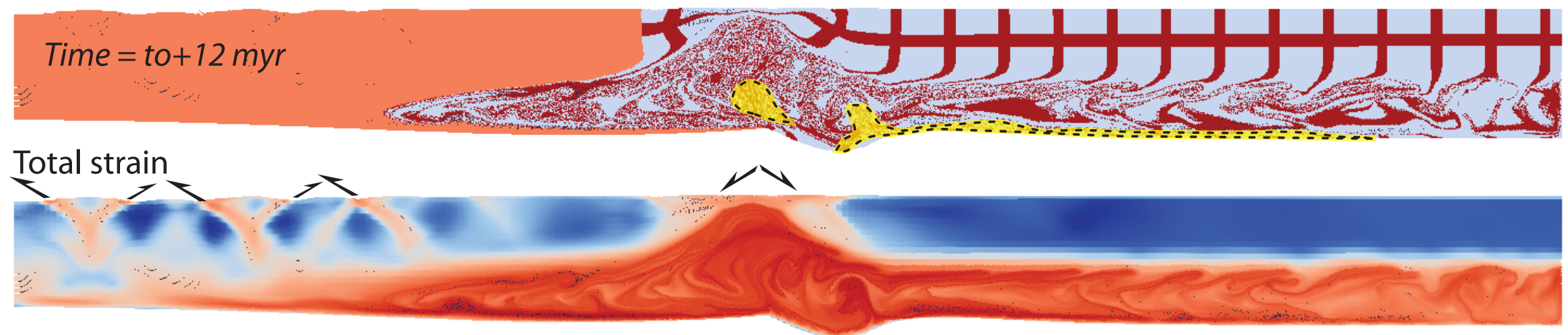


Moresi, Zhong & Gurnis, 2000

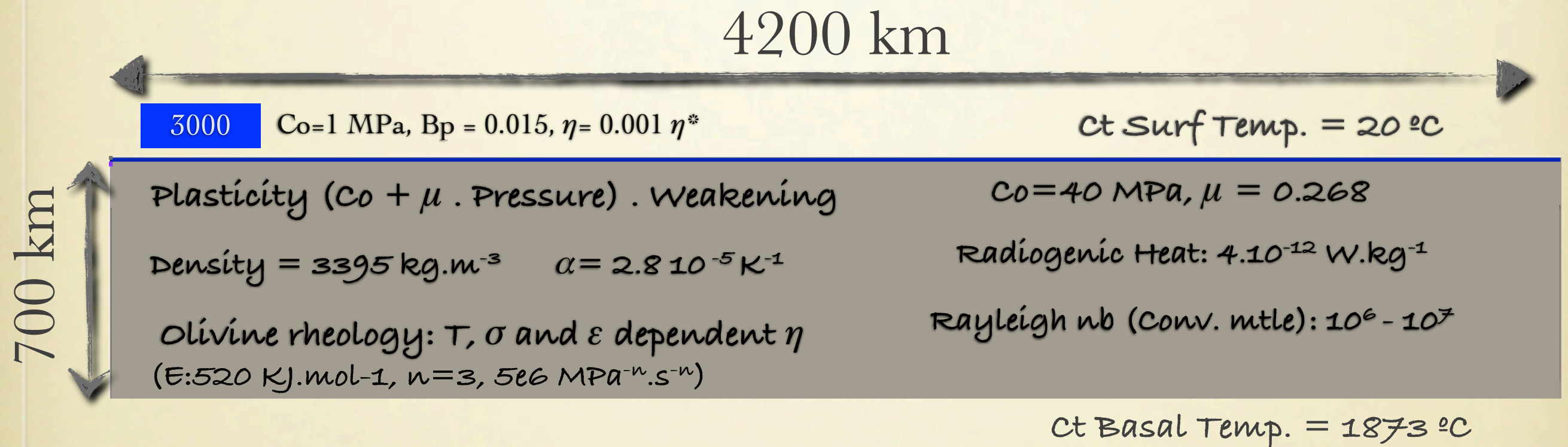




Fix foreland, melt with buoyancy



# Convective system without and with continents ...



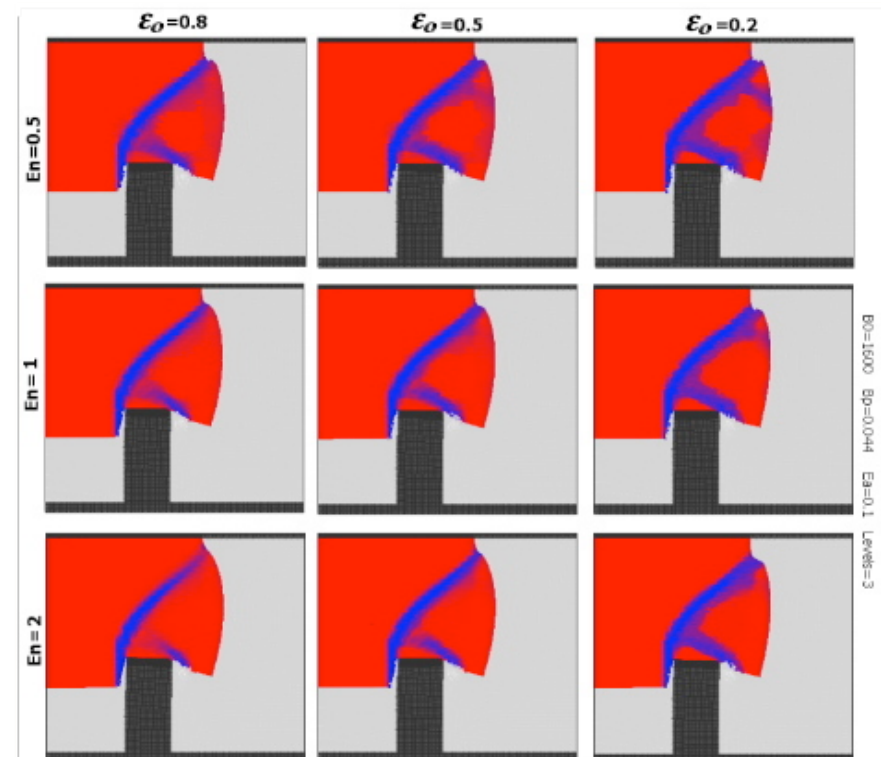
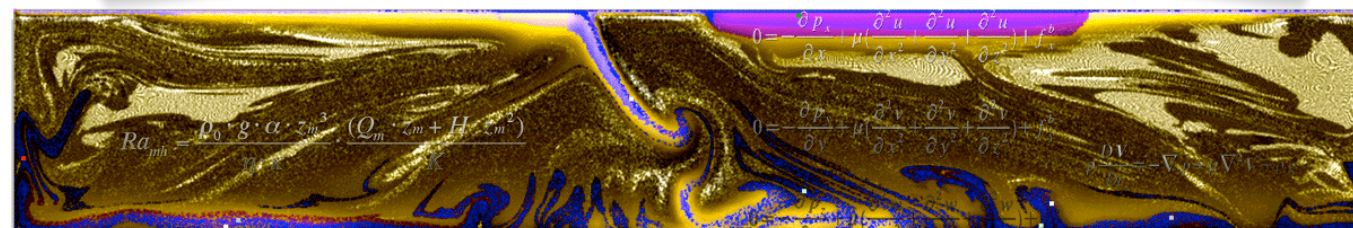
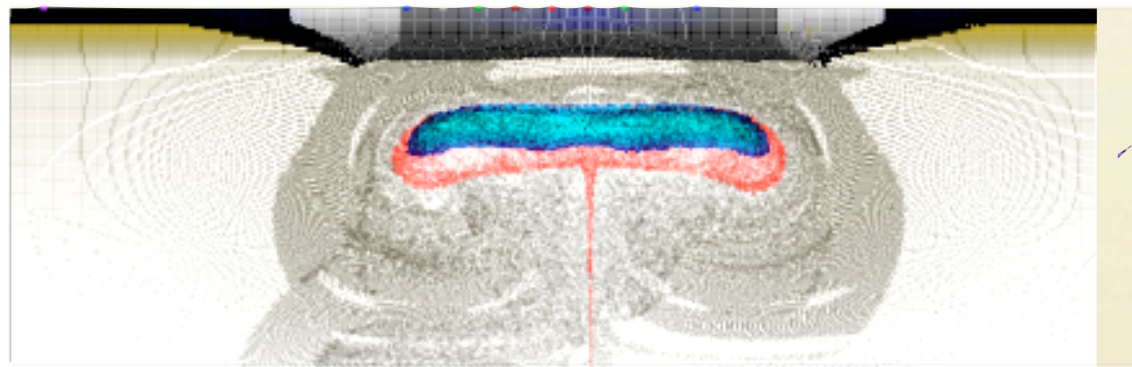
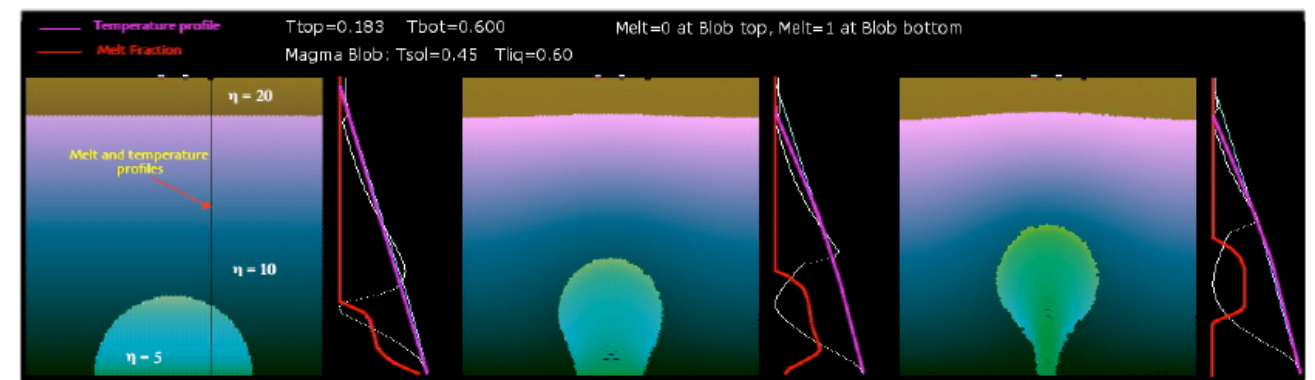
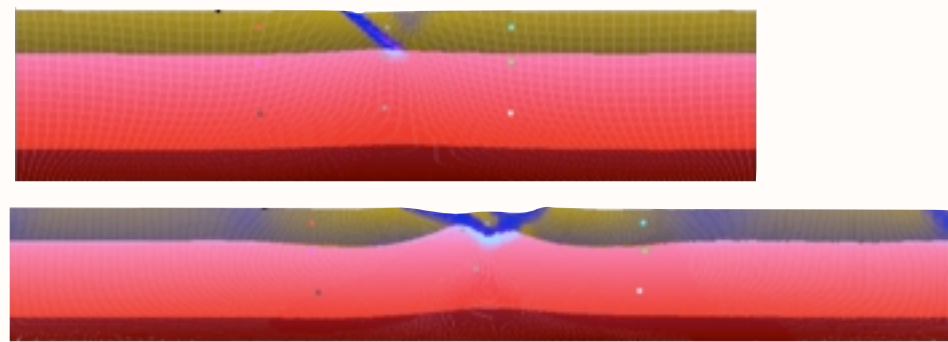
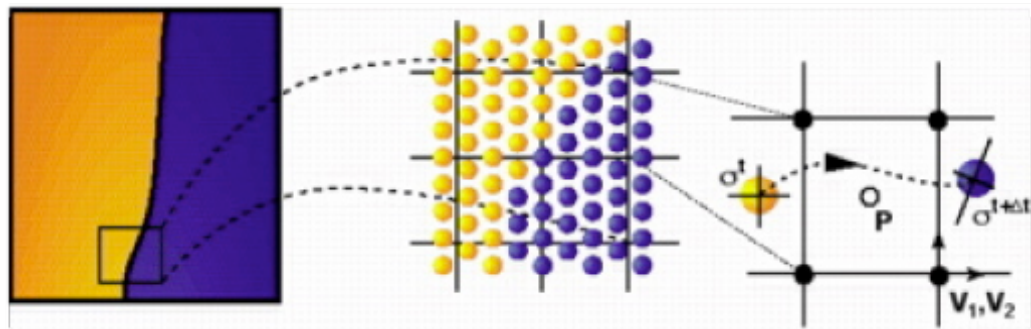


# Open source codes: *Ellipsis*, *Underworld*

Coupled thermal-mechanical  $\rho(T, \text{Metam.})$   
 $\eta(\sigma, \epsilon, \dot{\epsilon}, T)$

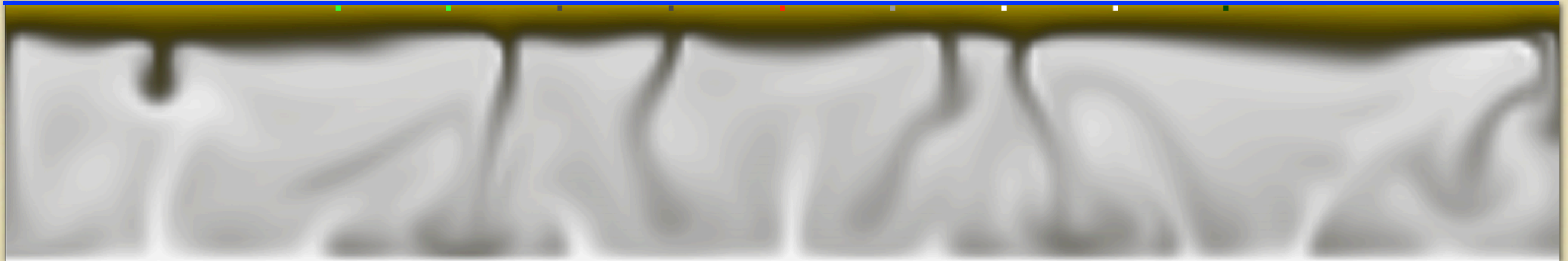
Visco-plastic rheology with strain weakening

Radiogenic heat, partial melting, eclogitization...



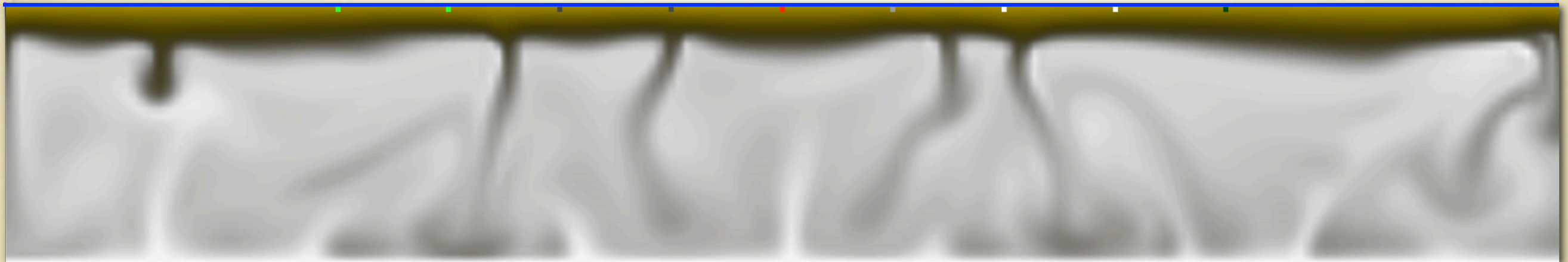


Convective system with no continent ...

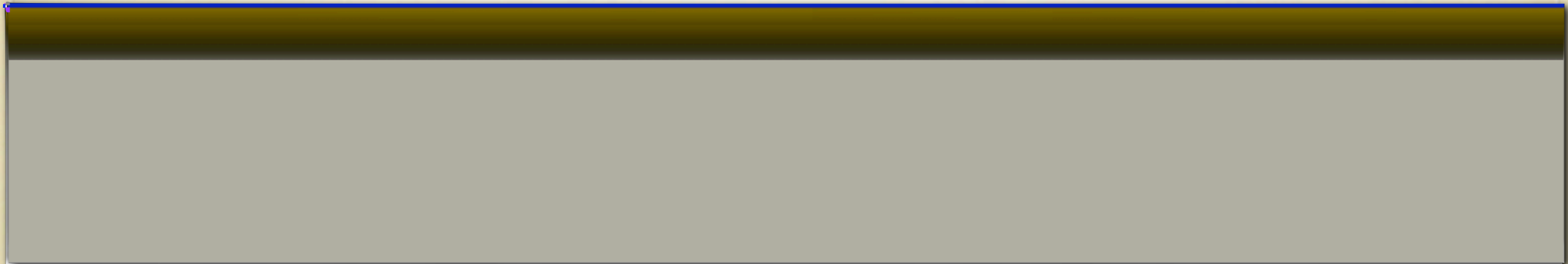




Convective system with continent ...

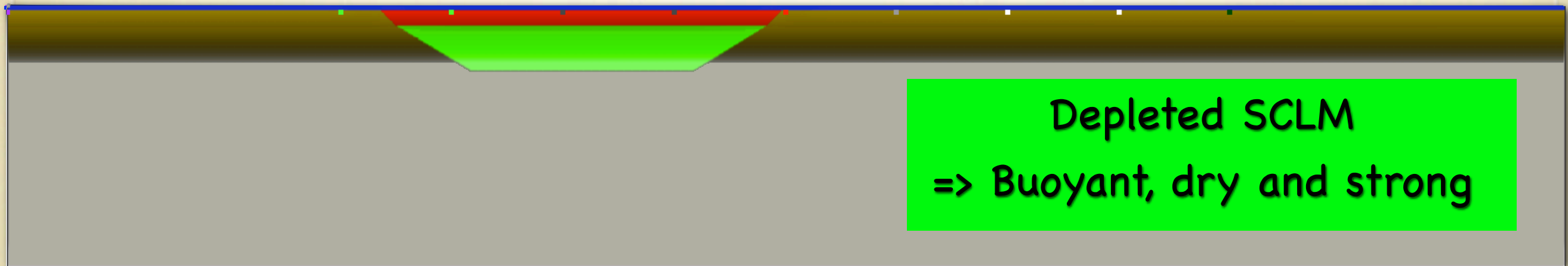


Convective system with continent ...





# Convective system with continent ...



15 km		3000 kg.m <sup>-3</sup>
40 km		2850 kg.m <sup>-3</sup>
120 km		3310 kg.m <sup>-3</sup>

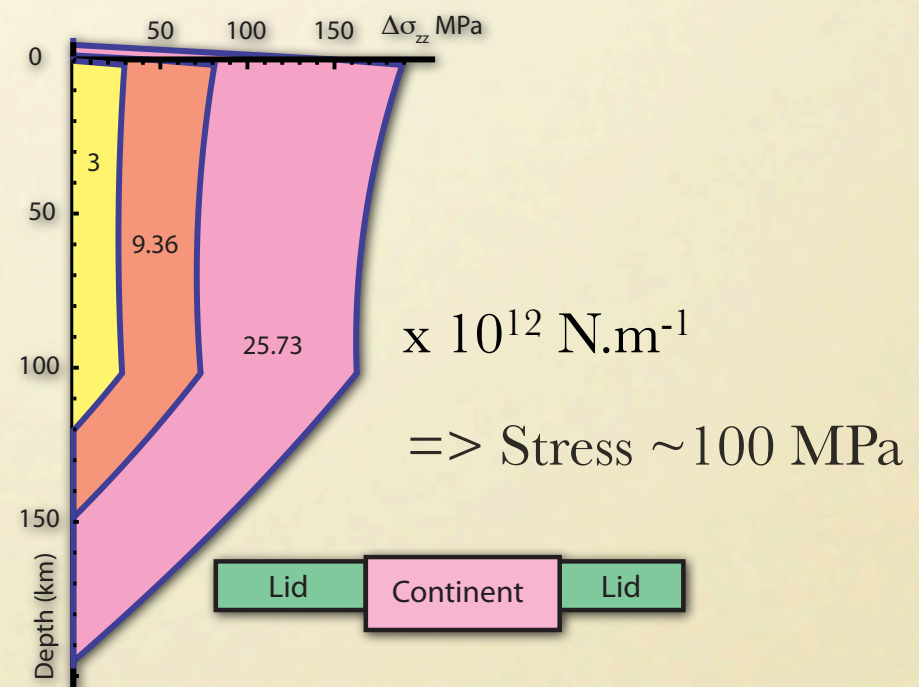
Total thickness: 175 km

# Convective system with continent ...

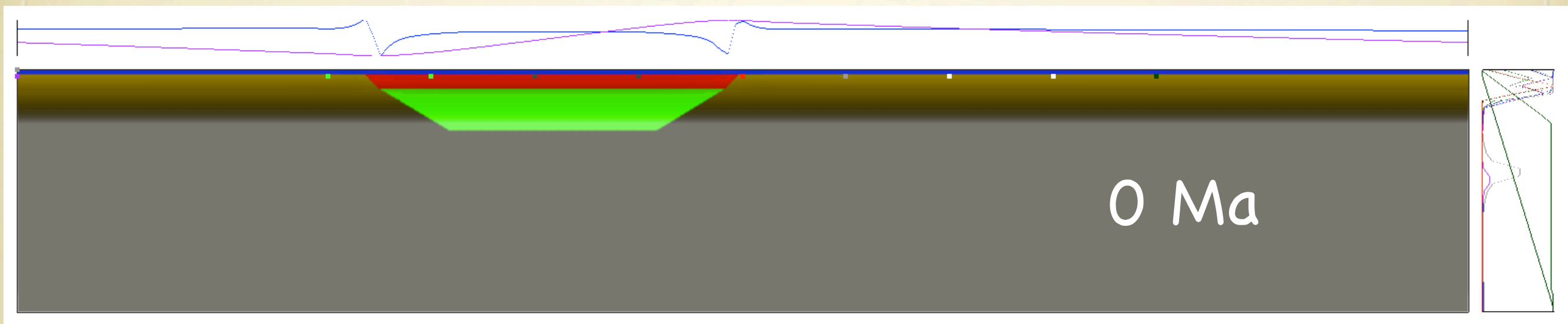
Depleted SCLM  
=> Buoyant, dry and strong

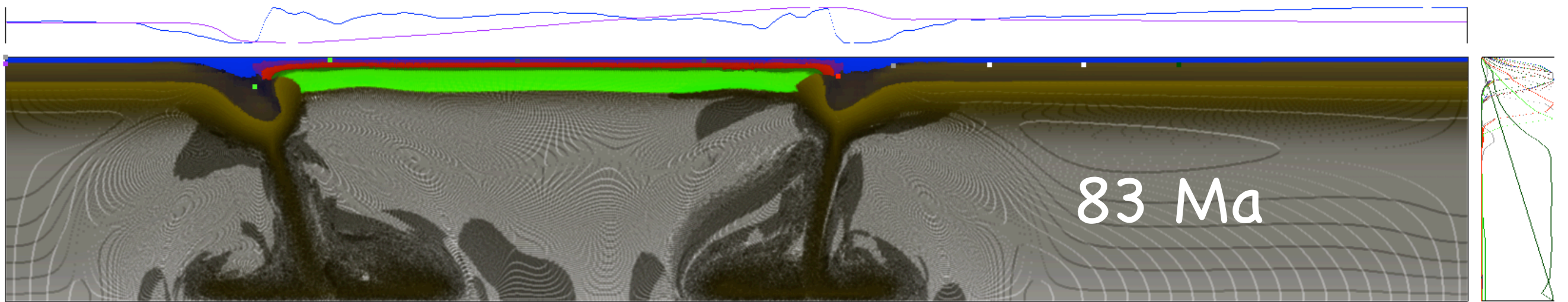
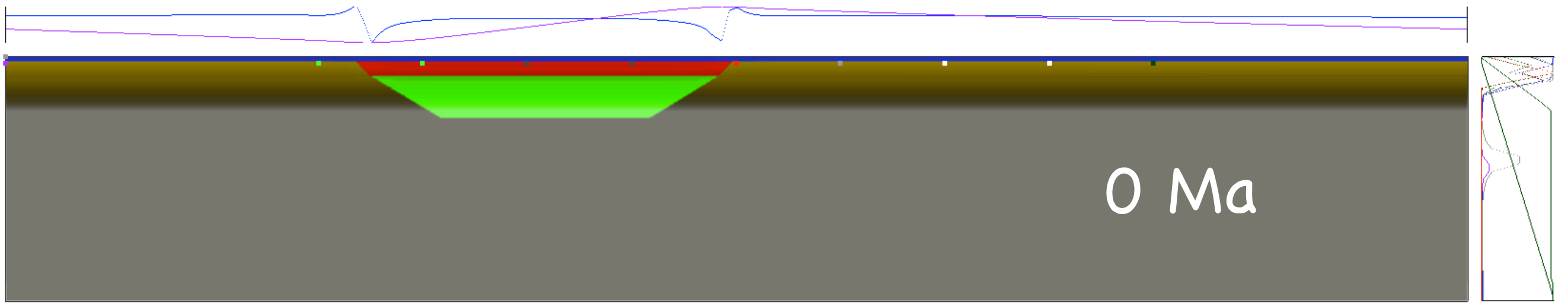
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Total thickness: 175 km

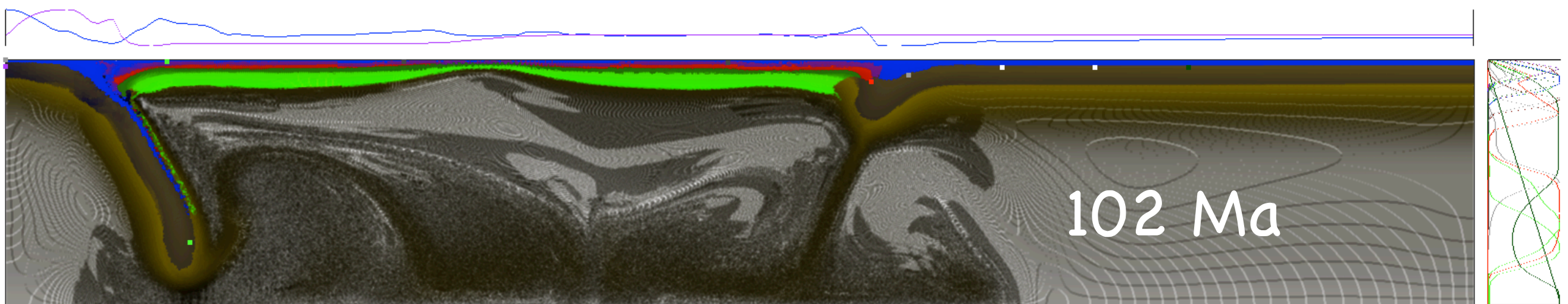
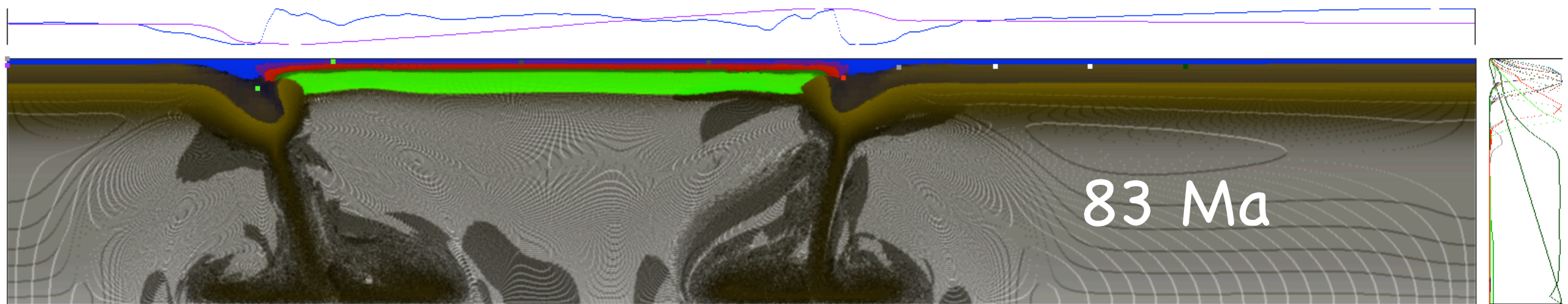
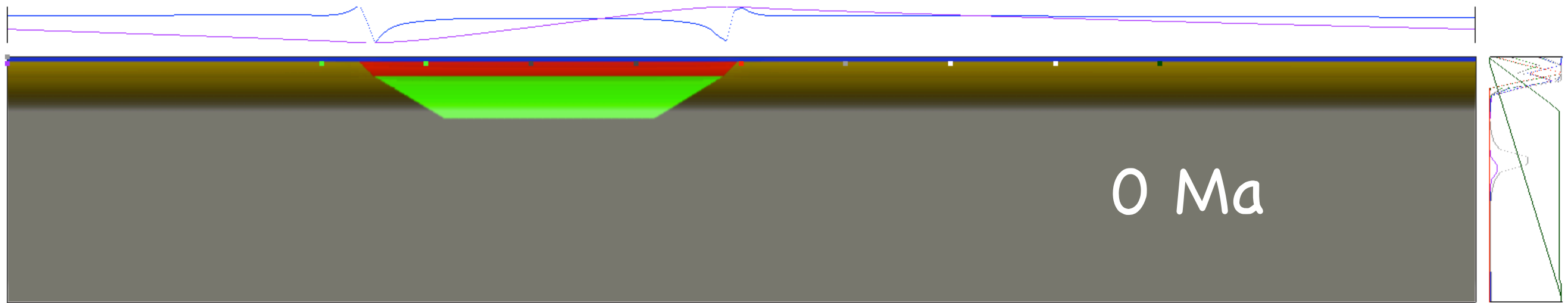




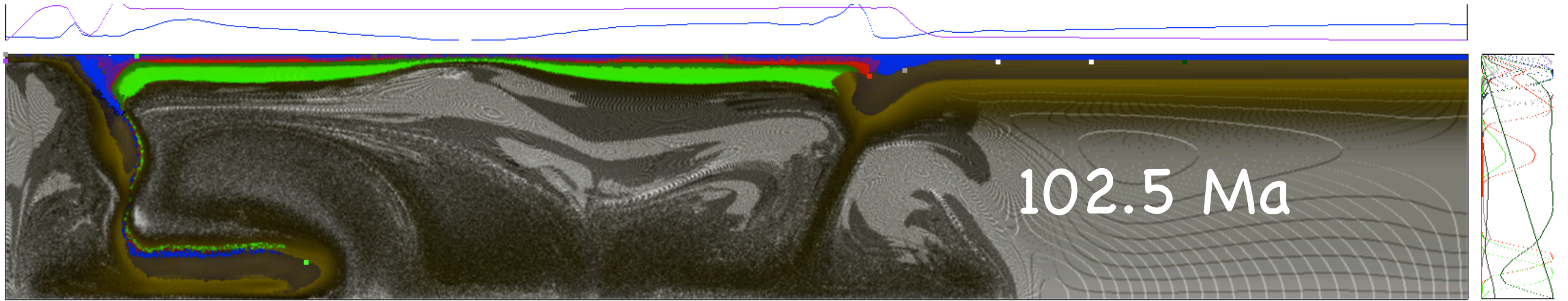




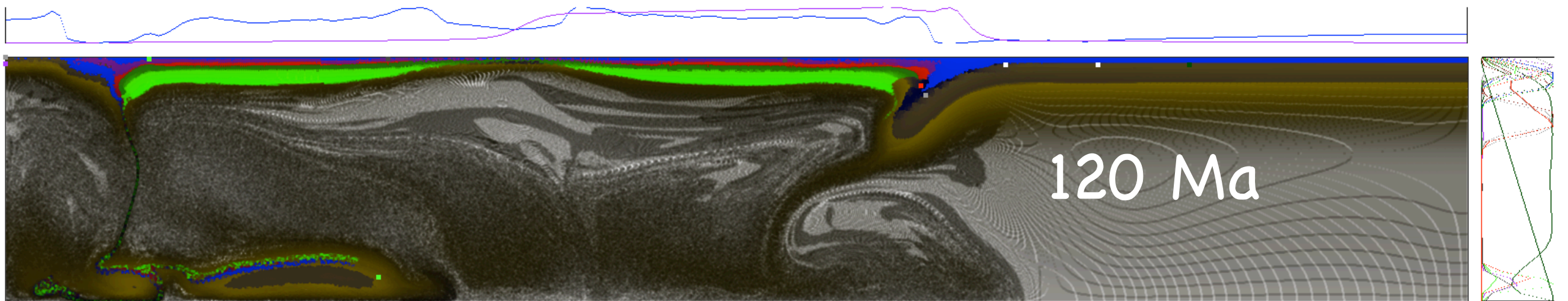
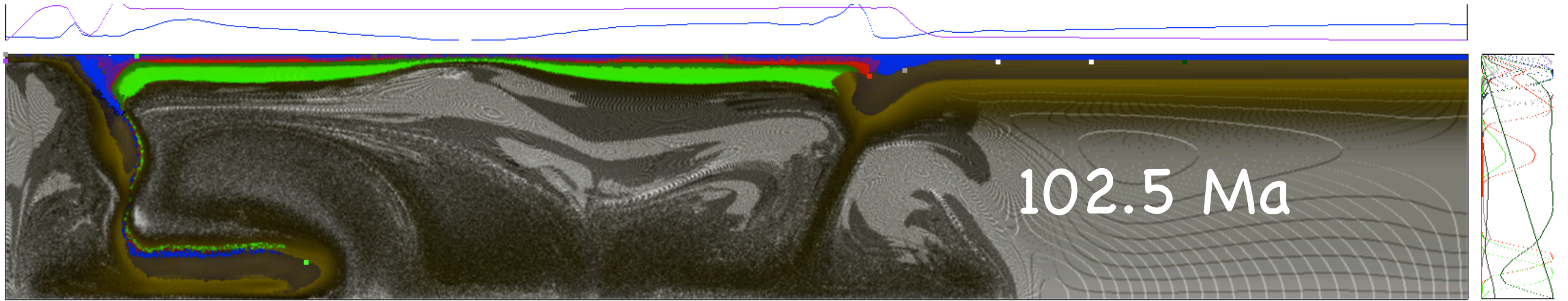




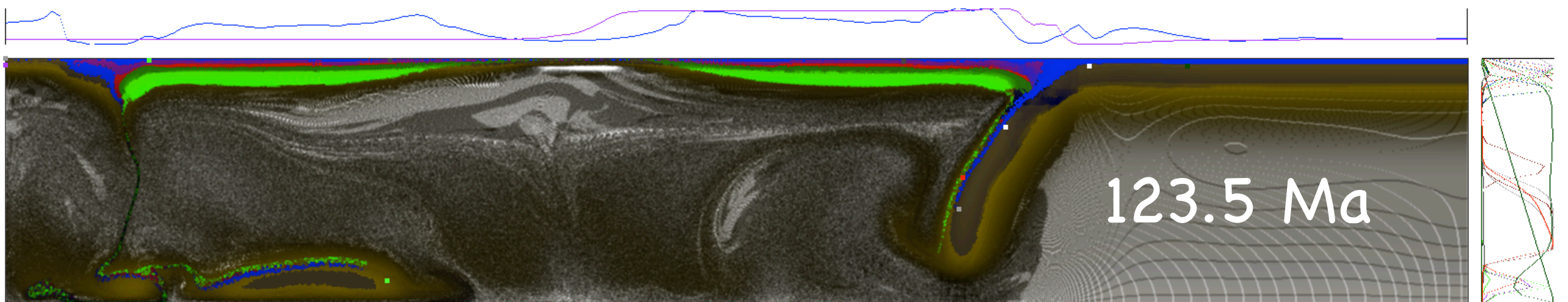
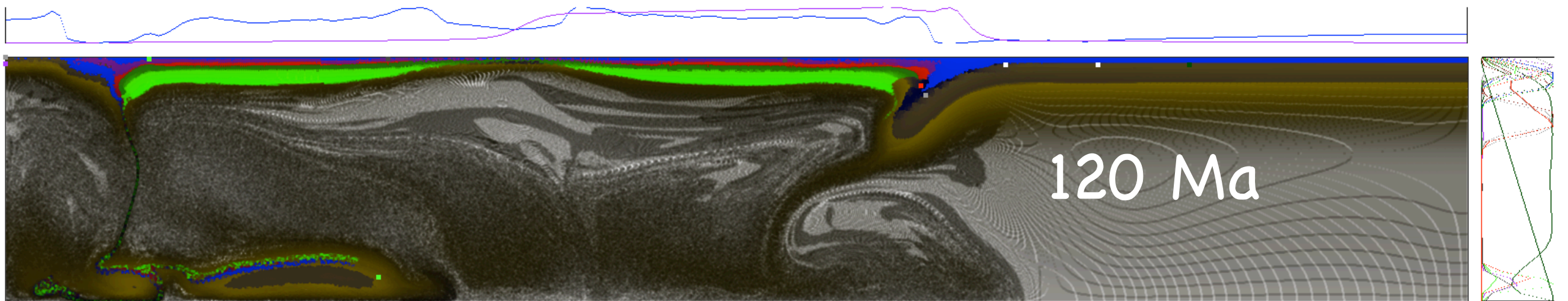
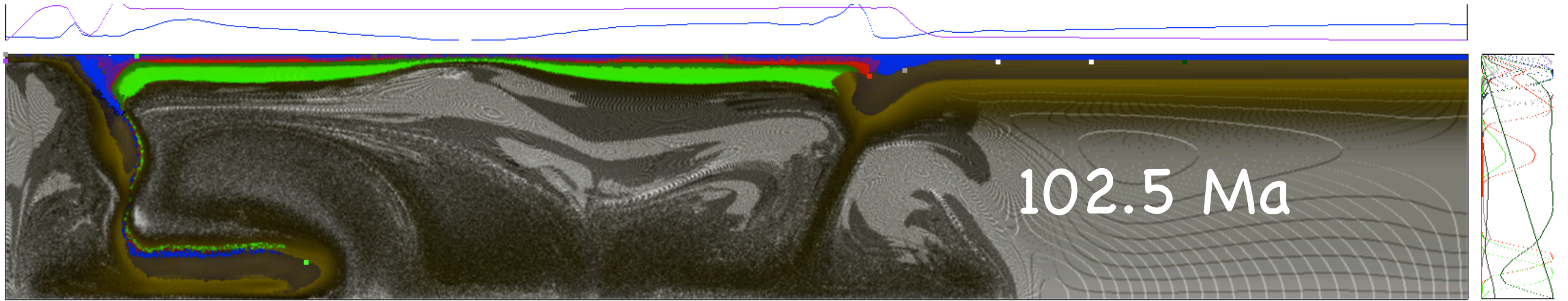




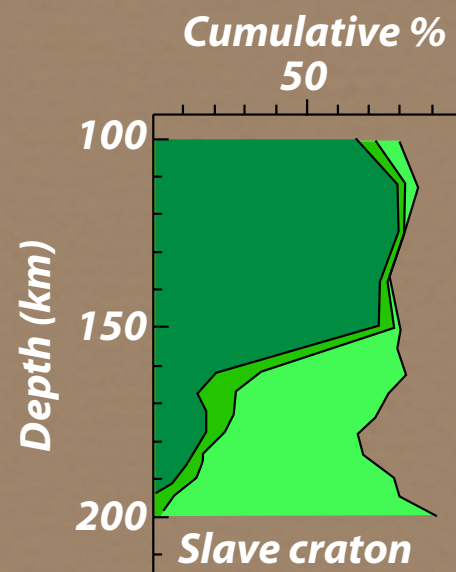
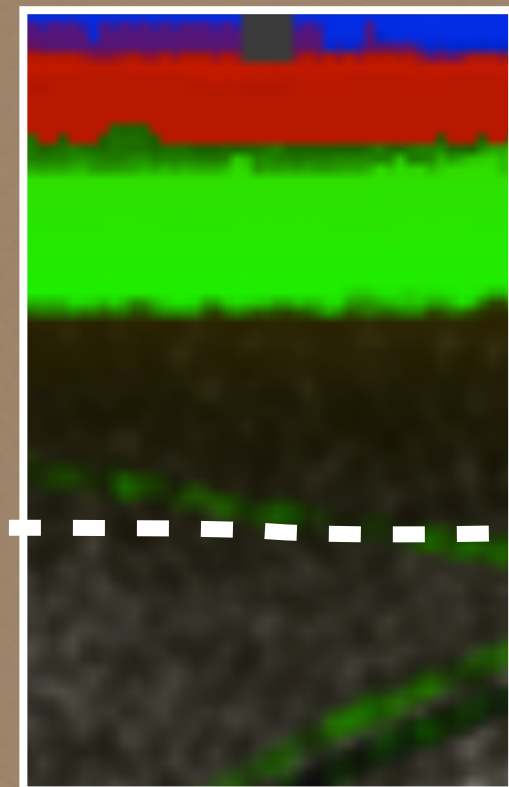
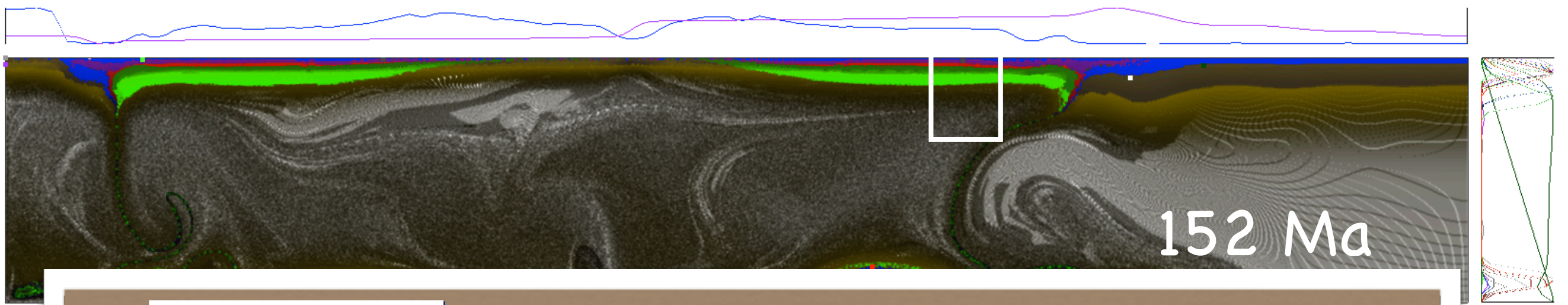
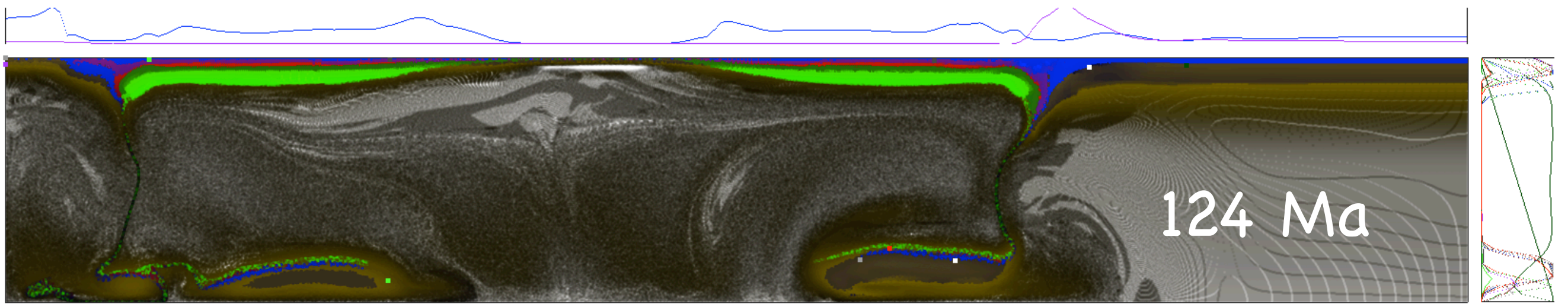












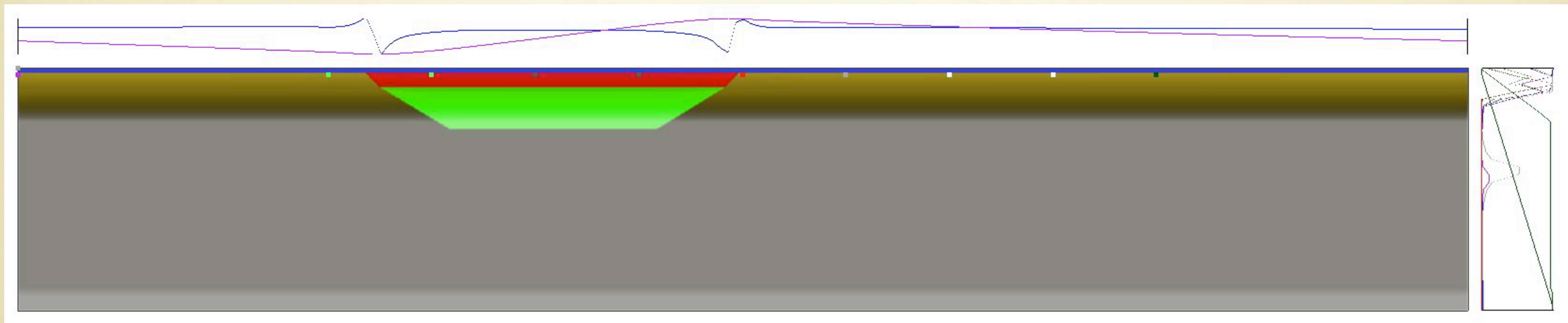
- Depleted Harzburgite & Lherzolite
- Fertile Lherzolite
- Melt metasomatized

Griffin et al. 1998

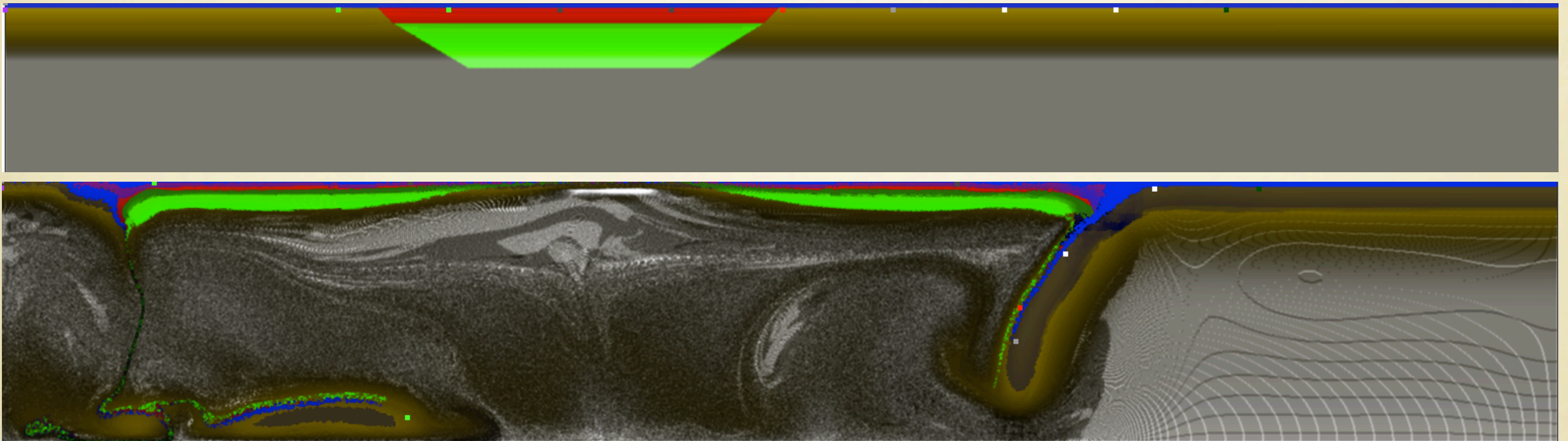








Did early continents  
crank-start plate tectonics?

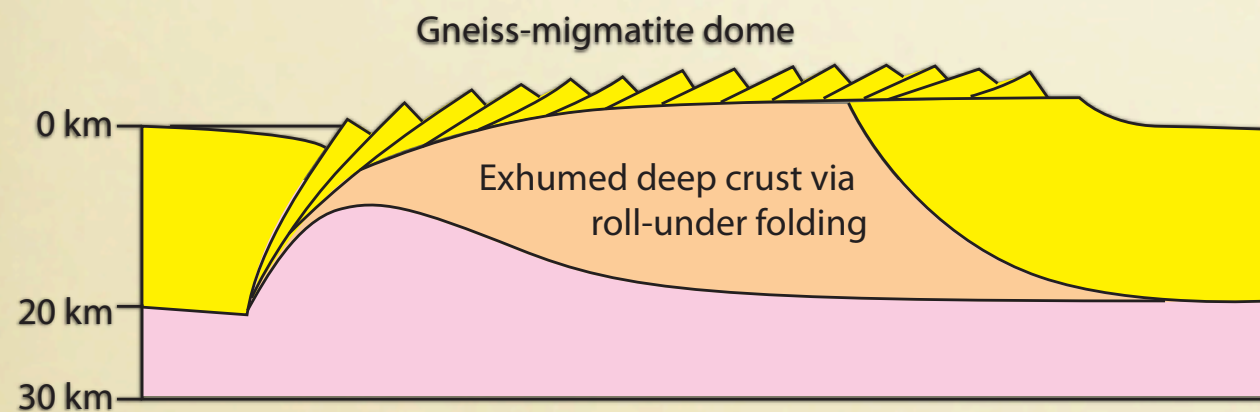
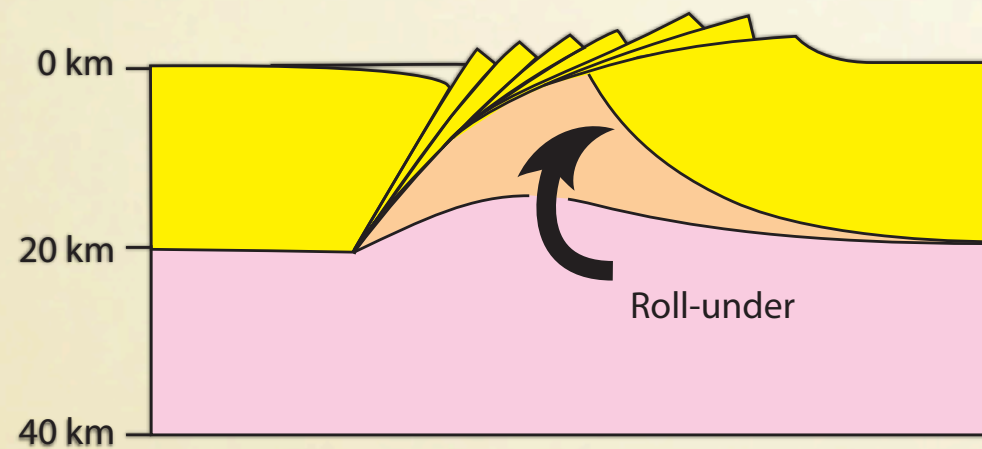
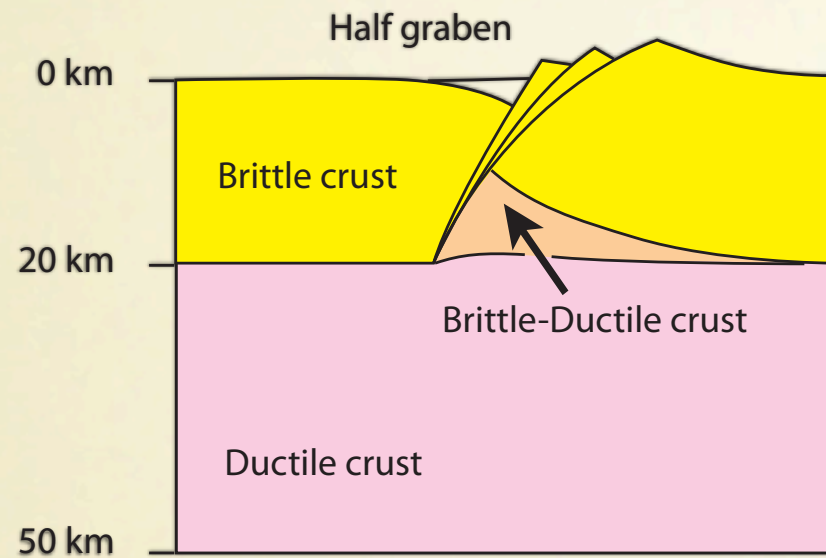




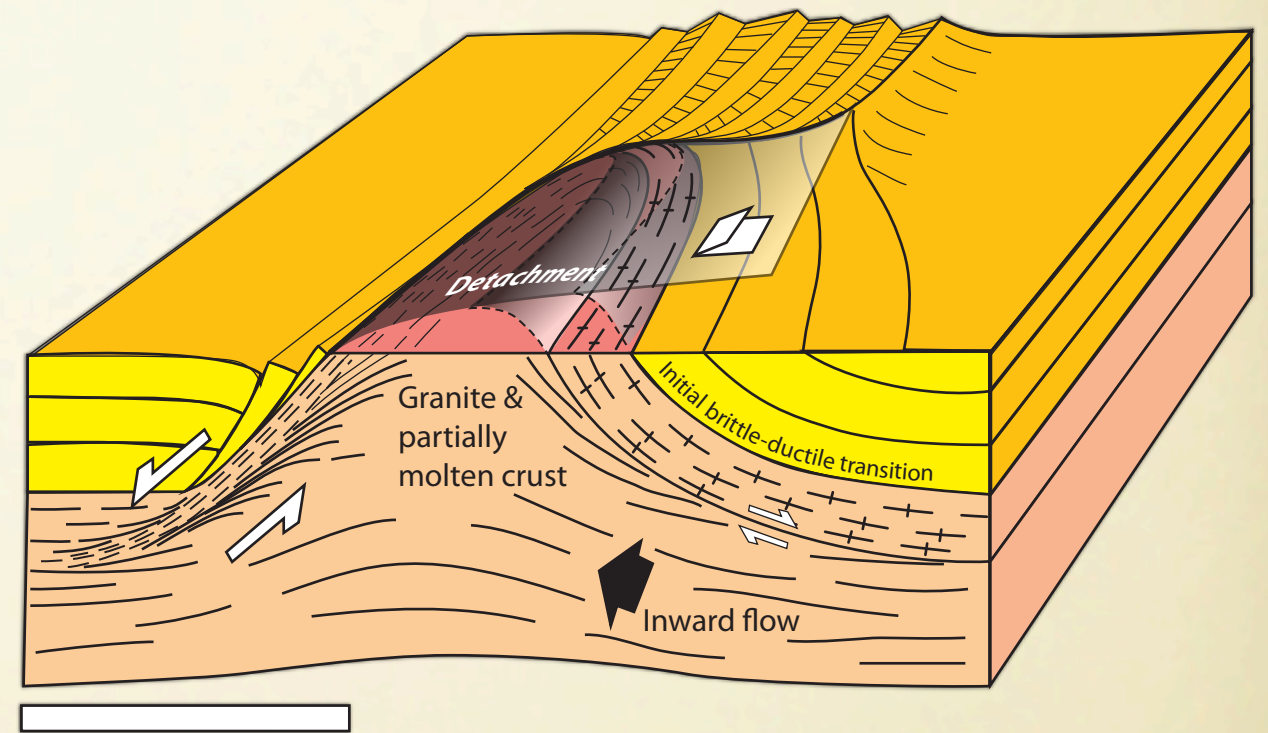
# STRAIN REGIMES IN HOT CRUSTS





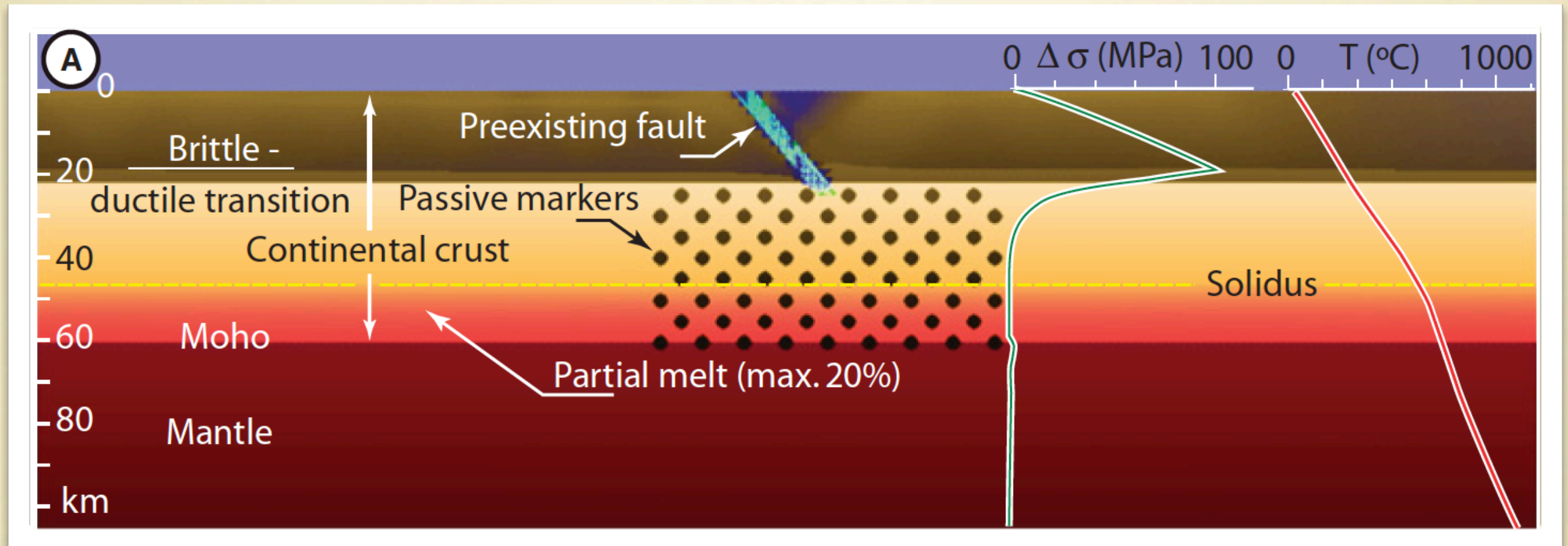


Van den Driessche & Brun, 1992



Brun & Van den Driessche, 1994

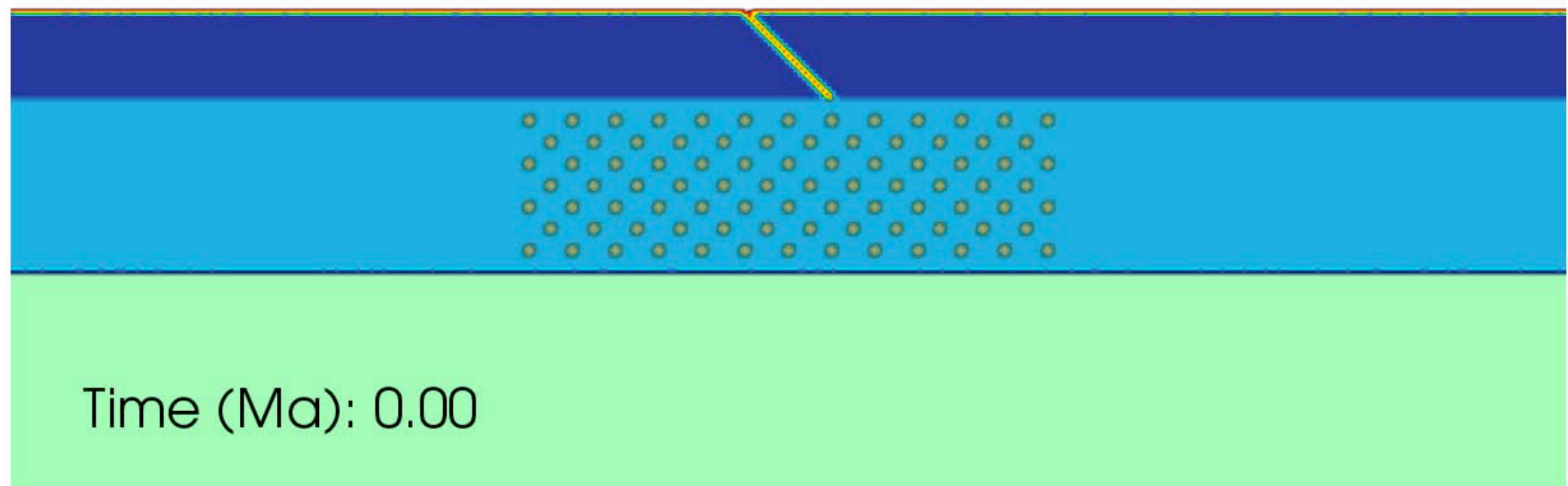




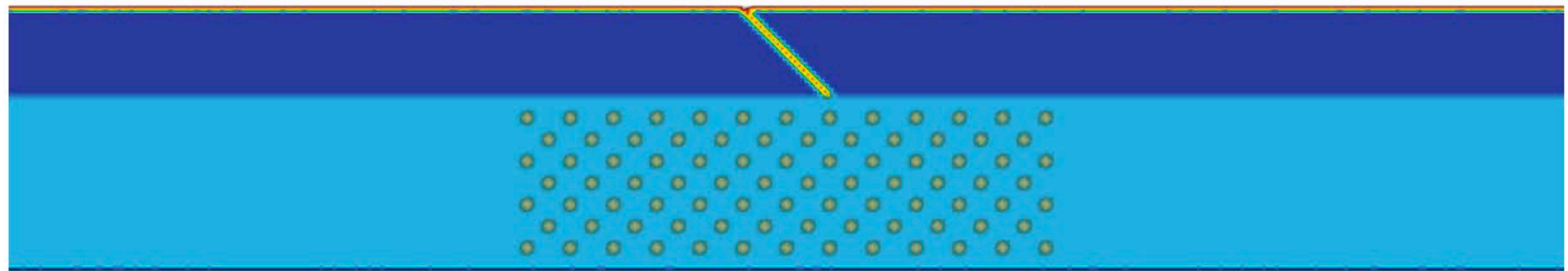


1.8 cm / yr



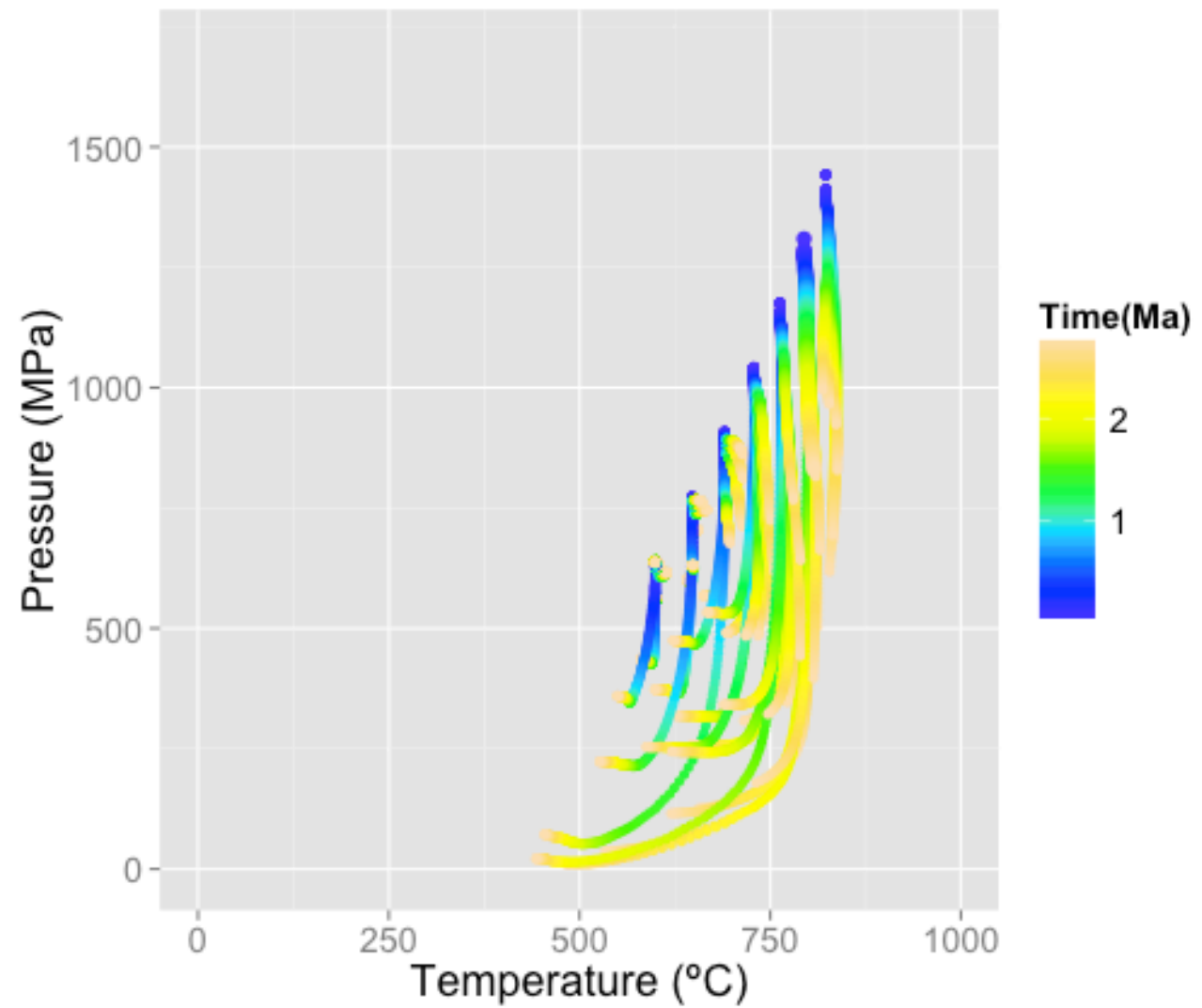


1.8 cm / yr



Time (Ma): 0.00

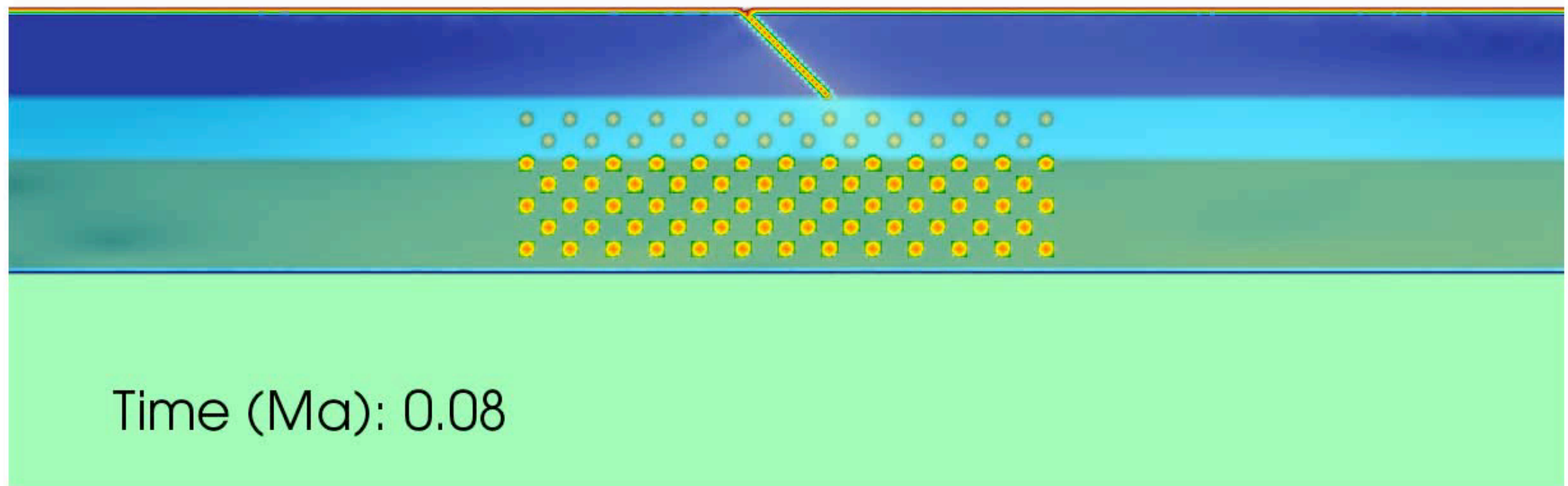
1.8 cm / yr





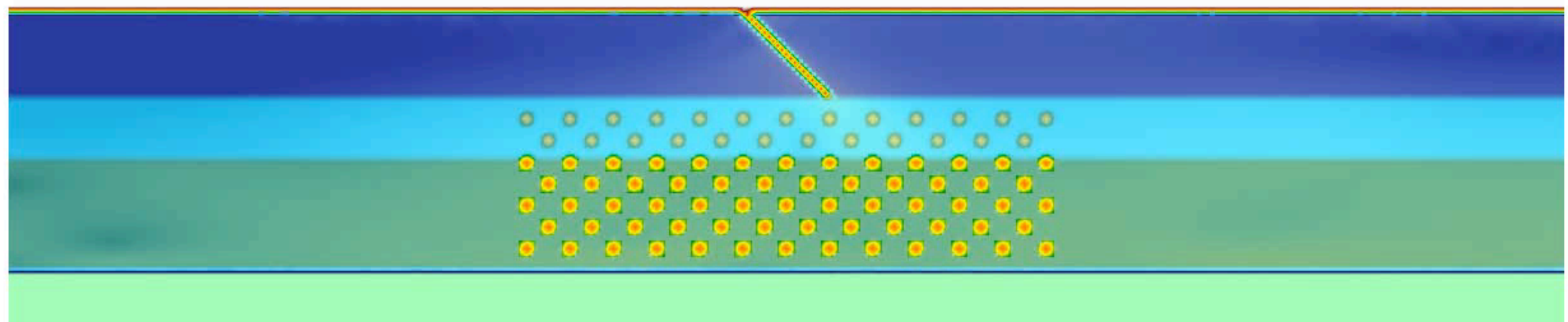


1.8 mm / yr



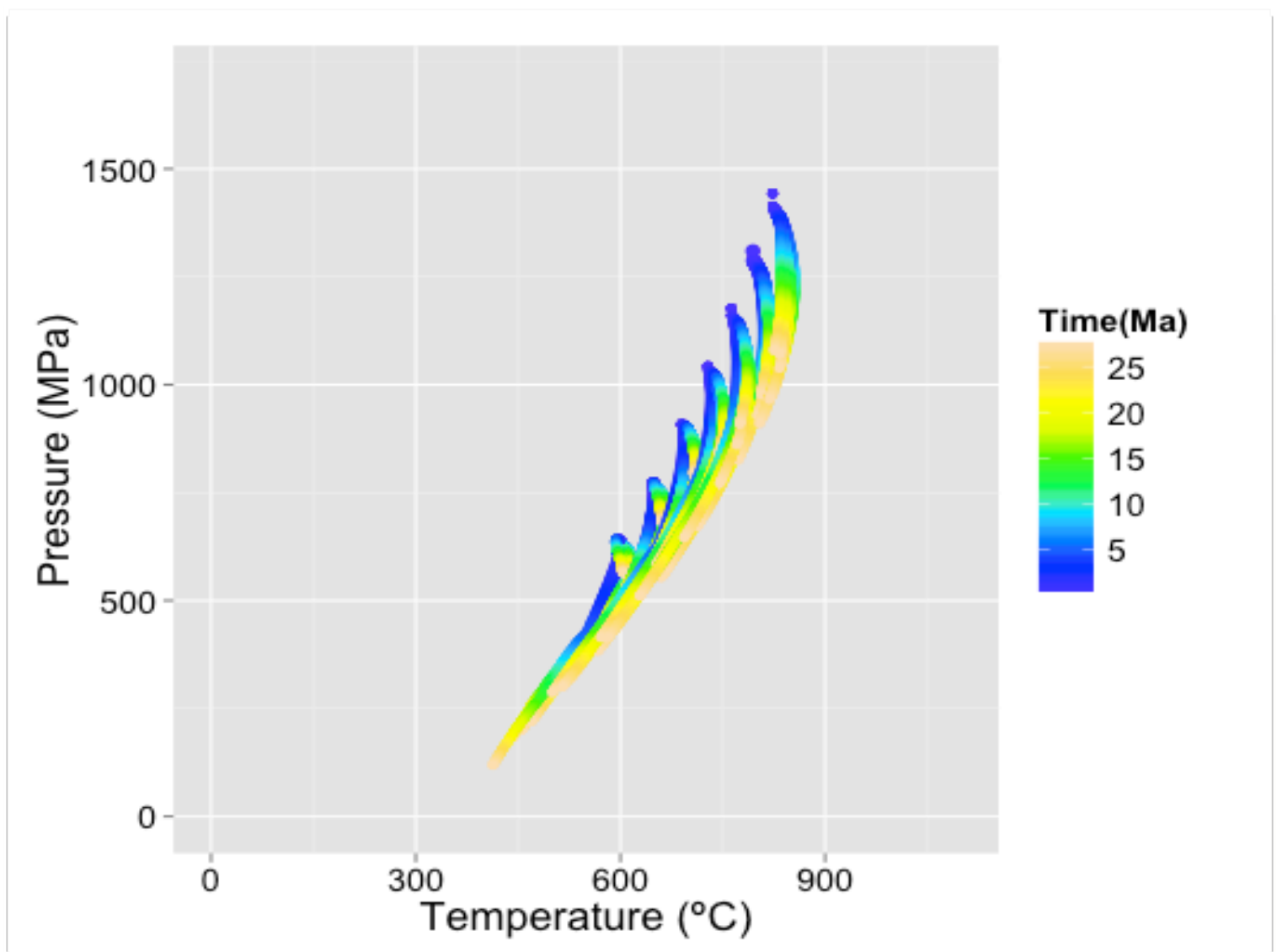
1.8 mm / yr

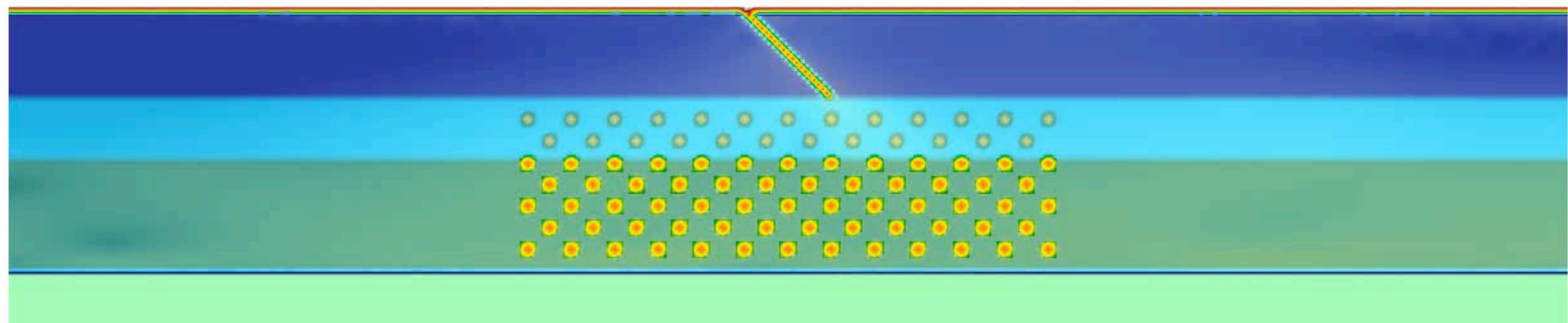




Time (Ma): 0.08

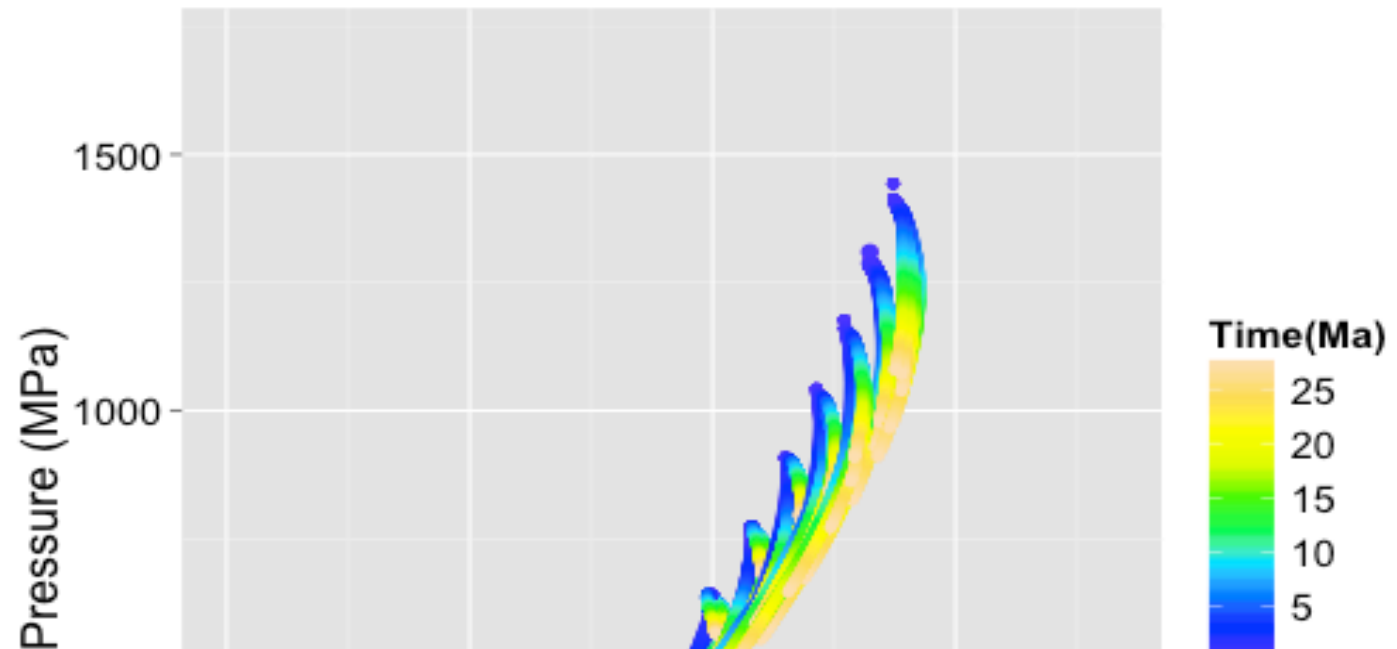
1.8 mm / yr





Time (Ma): 0.08

1.8 mm / yr



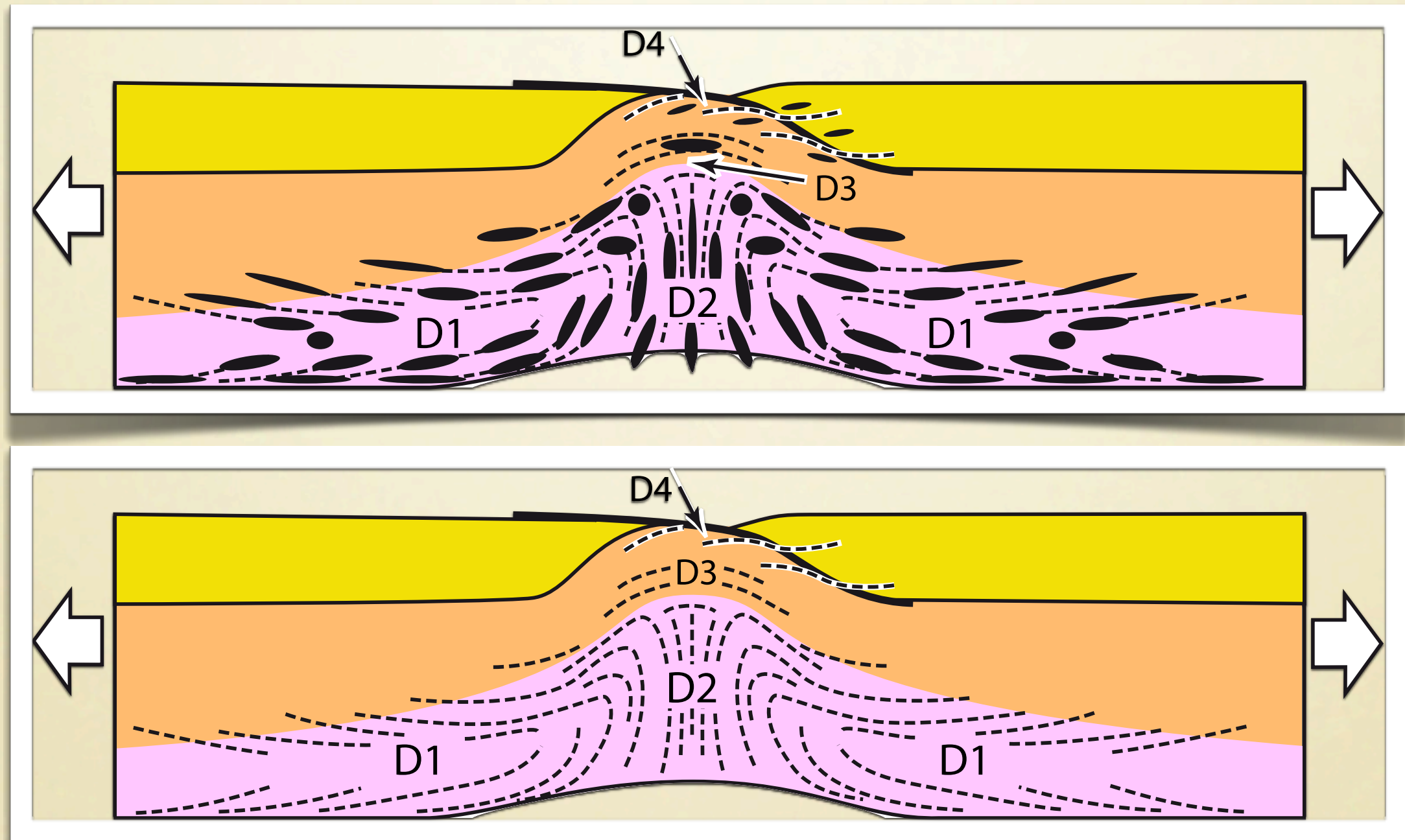
Viscous collision in channel explains double domes in metamorphic core complexes

Patrice F. Rey<sup>1</sup>, Christian Teyssier<sup>2</sup>, Seth C. Kruckenberg<sup>3</sup>, and Donna L. Whitney<sup>2</sup>

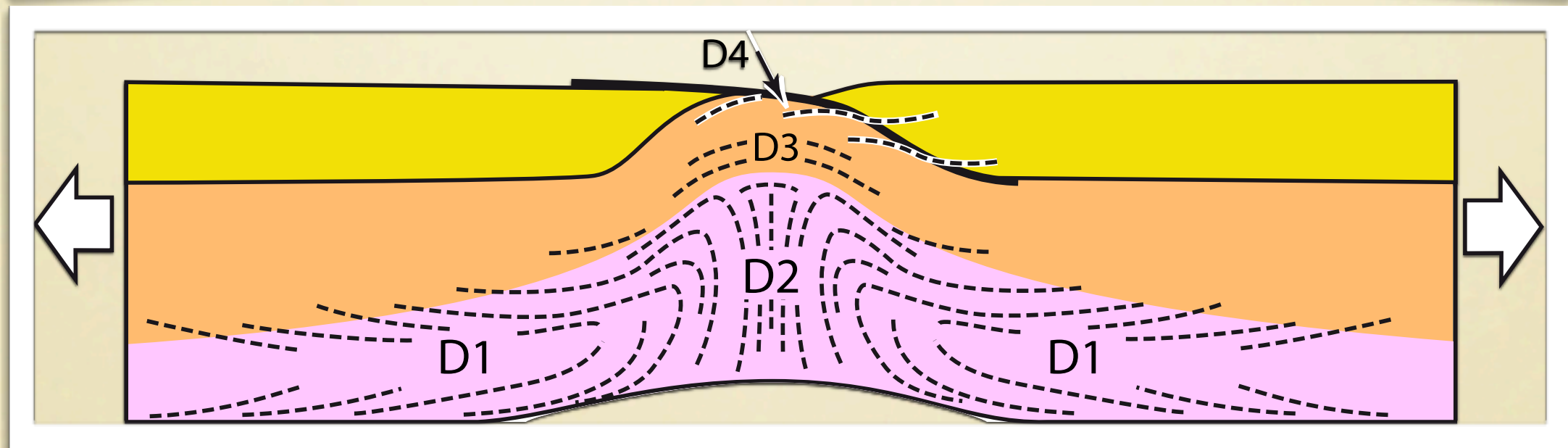
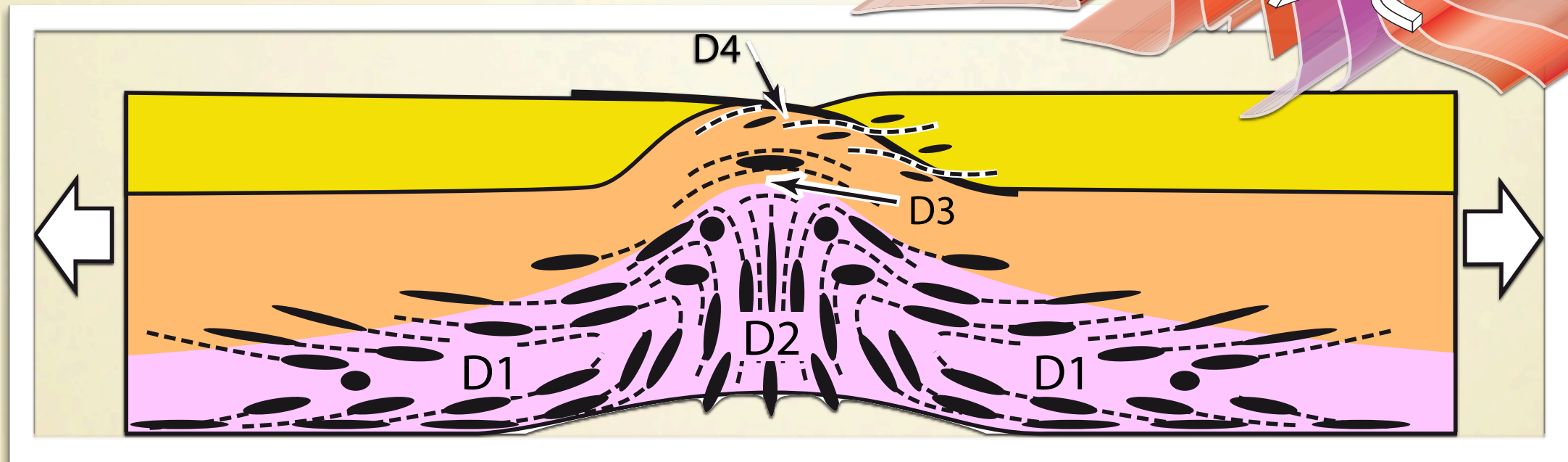
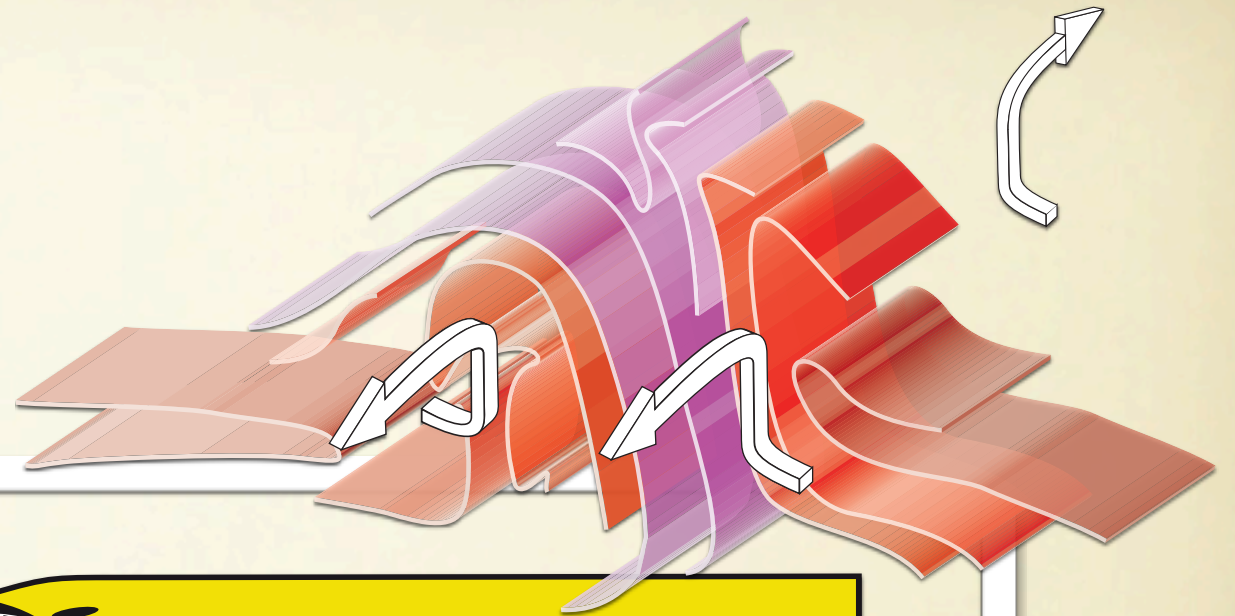
**GEOLOGY**, April 2011



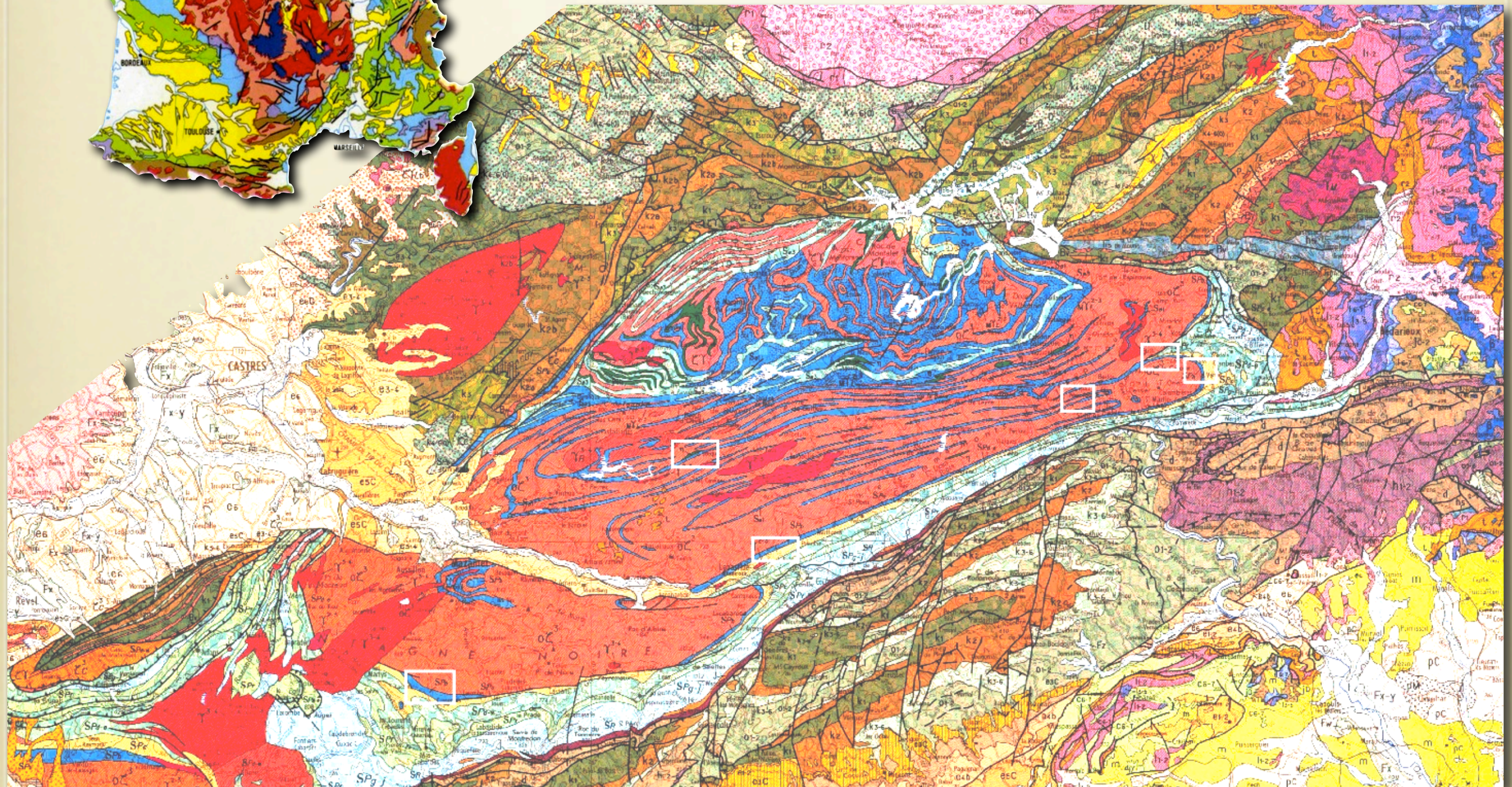
# Strain regime partitioning



# Strain regime partitioning

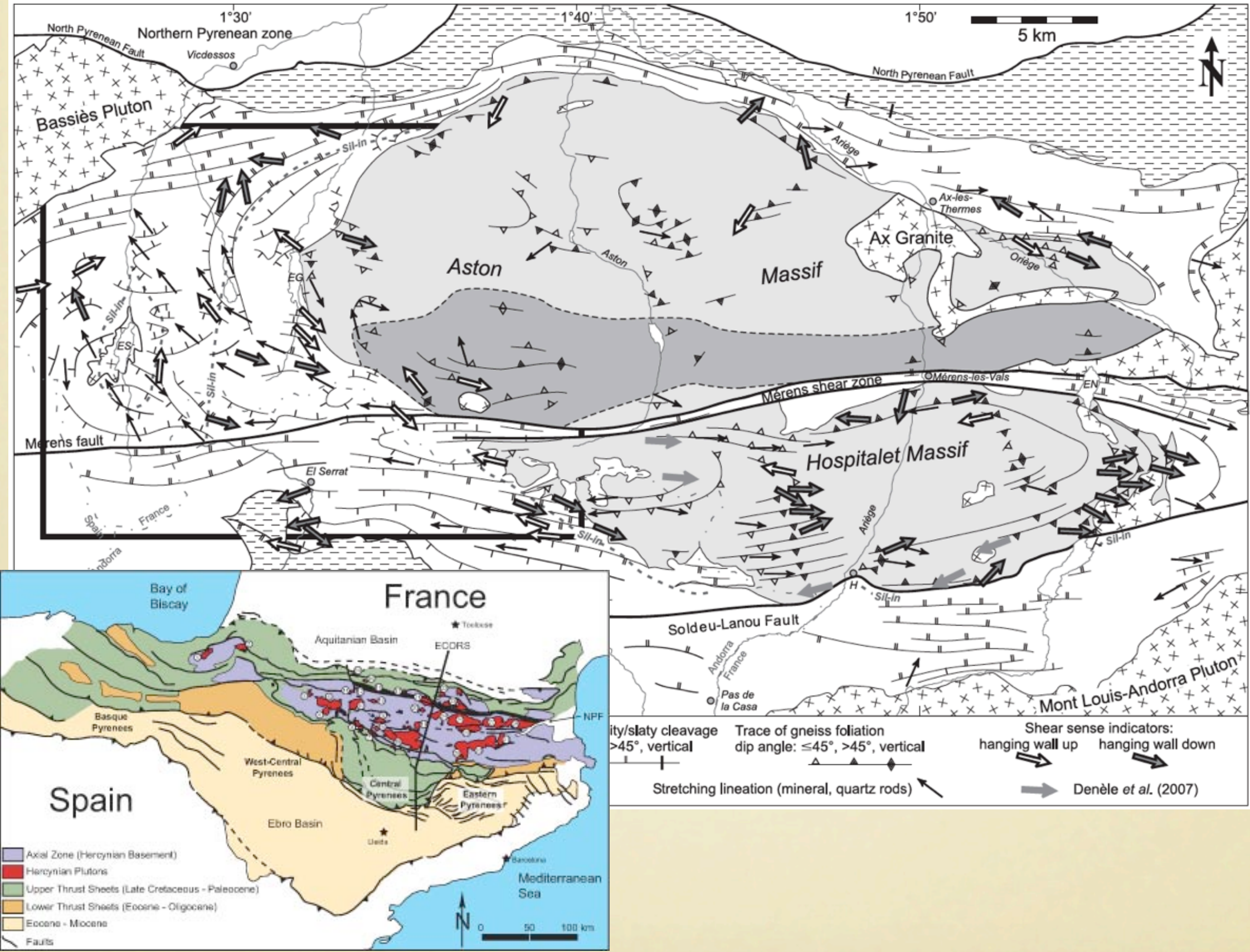




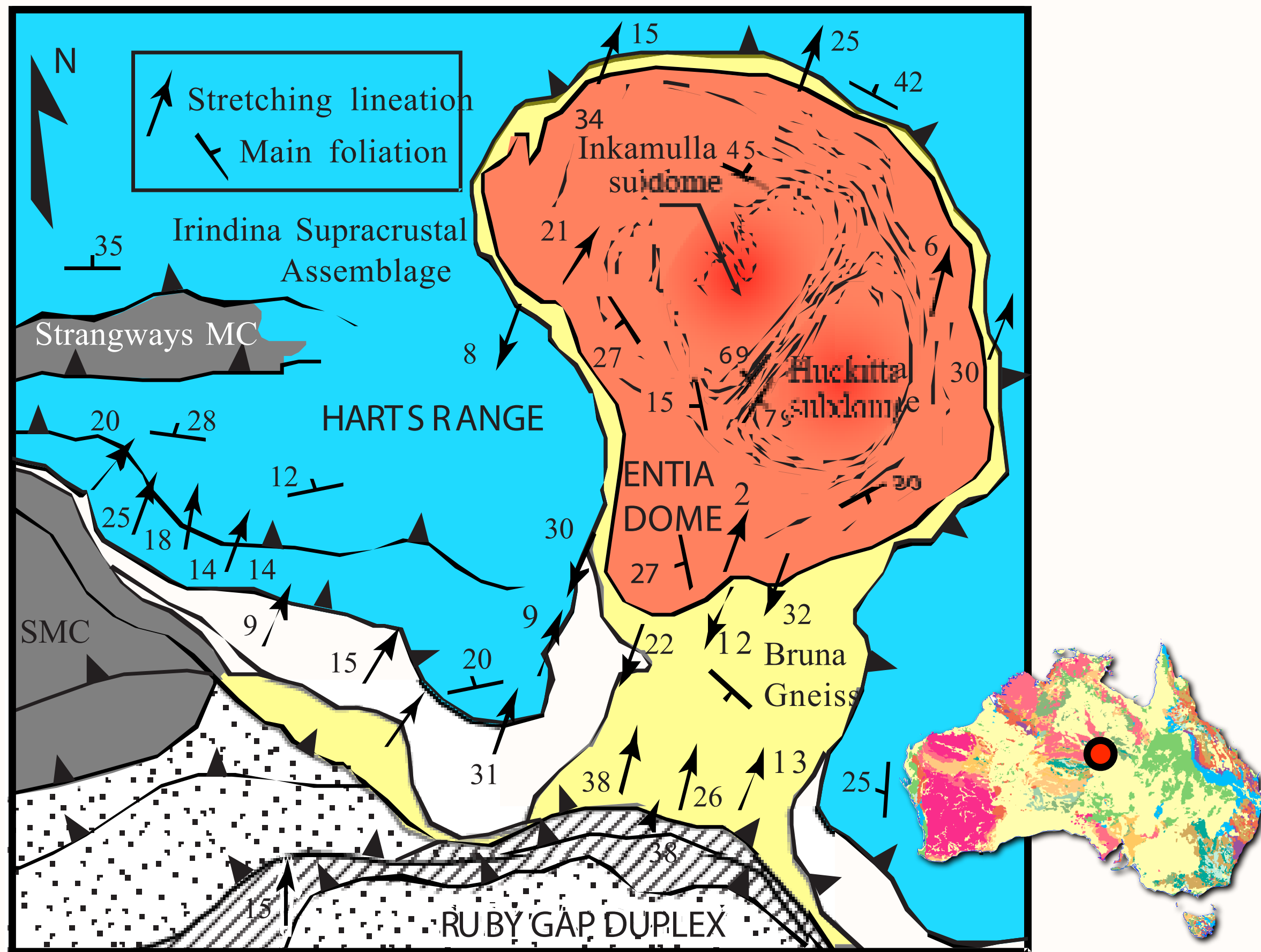




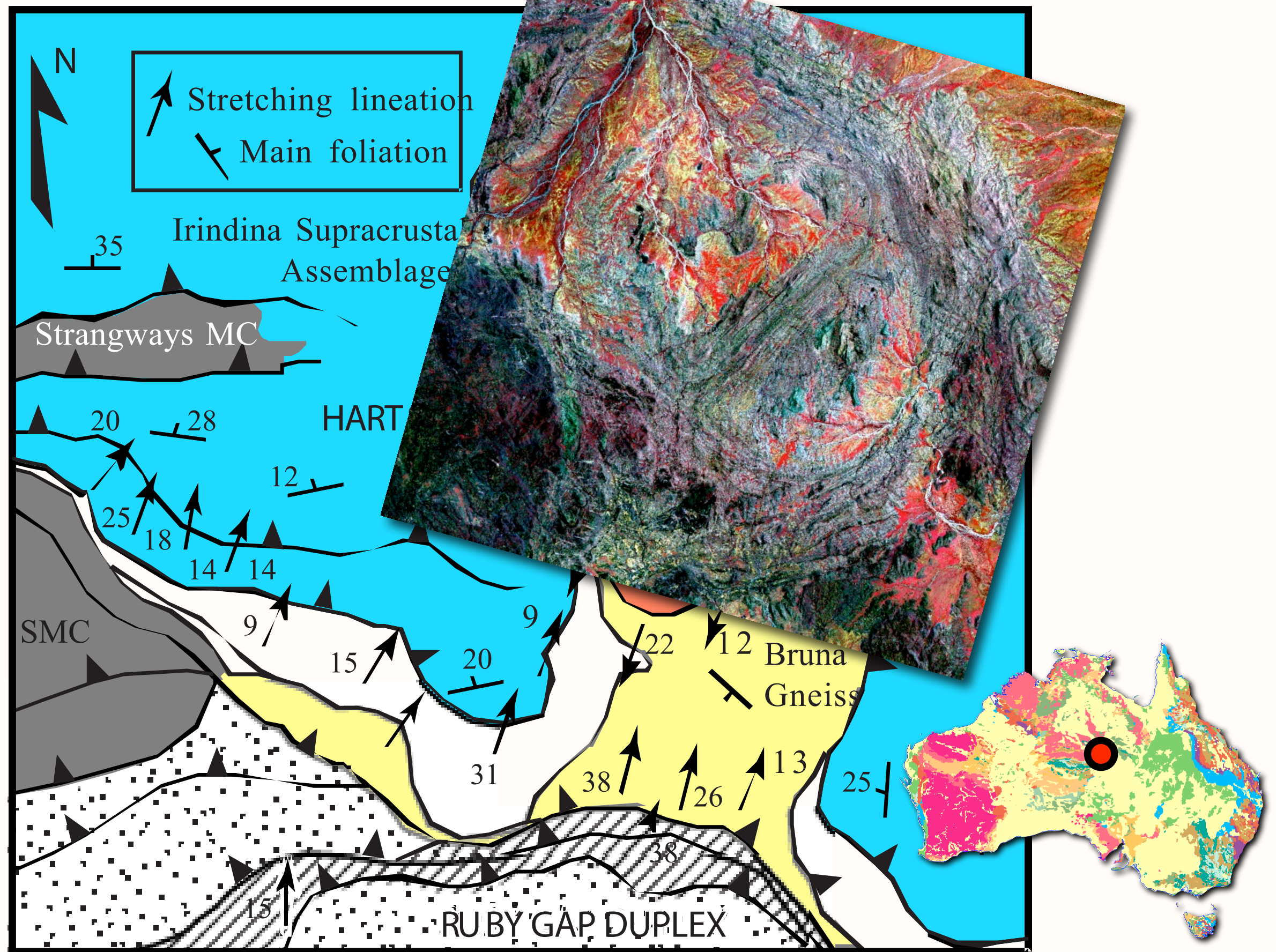
## Mezger, 2004













## Strain regimes in hot crusts:

Coeval contractional, extensional and shear fabrics develop in various parts of hot extending crusts.

During extension, hot rocks are advected through regions of contrasting “tectonic regimes”.

